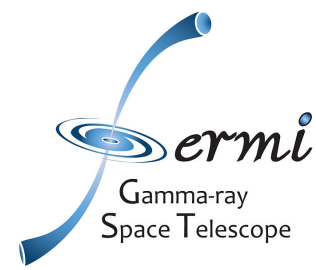




UC SANTA CRUZ



Stefano Profumo

UC Santa Cruz

Santa Cruz Institute for Particle Physics

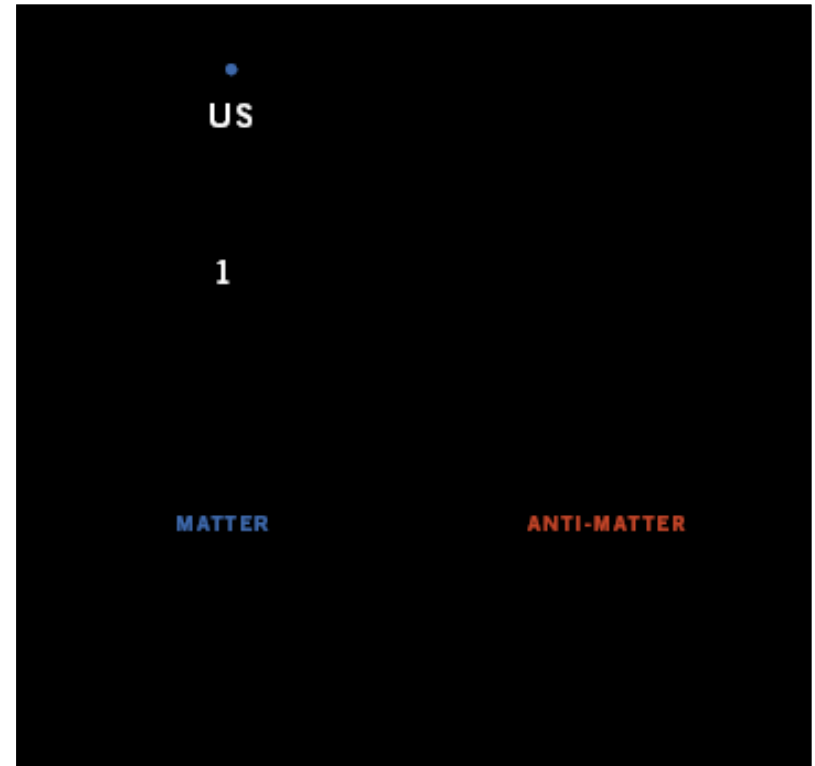
Affiliated Member, Fermi Collaboration

***Terascale: key to the origin of
baryonic and dark matter?***

UCSC - Physics 205

Monday February 11, 2013

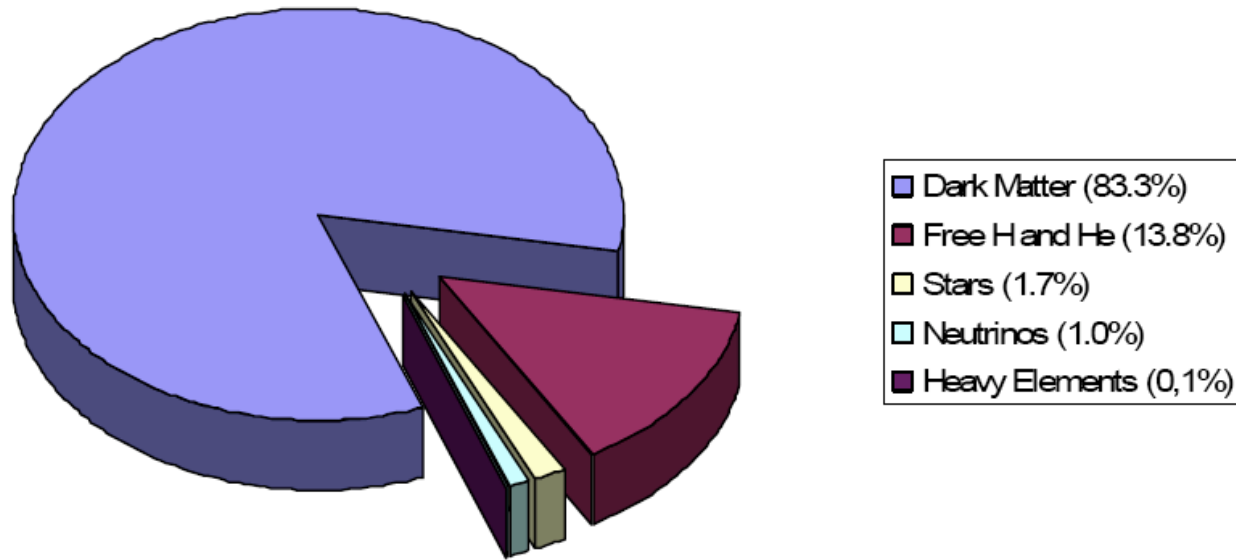
- ✓ PhD **Theoretical Particle Physics** (2004)
International School for Advanced Studies (SISSA-ISAS), Trieste, Italy
- ✓ Postdoc, FSU and California Institute of Technology (2005-2007)
Theoretical Astrophysics and Particle Physics
- ✓ Joined **UCSC Physics** Faculty (Assistant Professor, 2007-2011,
Associate Professor, July 2011-)
- ✓ Research funded by Department of Energy (Outstanding Junior Investigator Award), National Science Foundation, NASA
- ✓ **Fermi**-LAT affiliated scientist; collaborate with **ATLAS** group
- ✓ **SCIPP Deputy Director** for **Theory** (July 2011-)



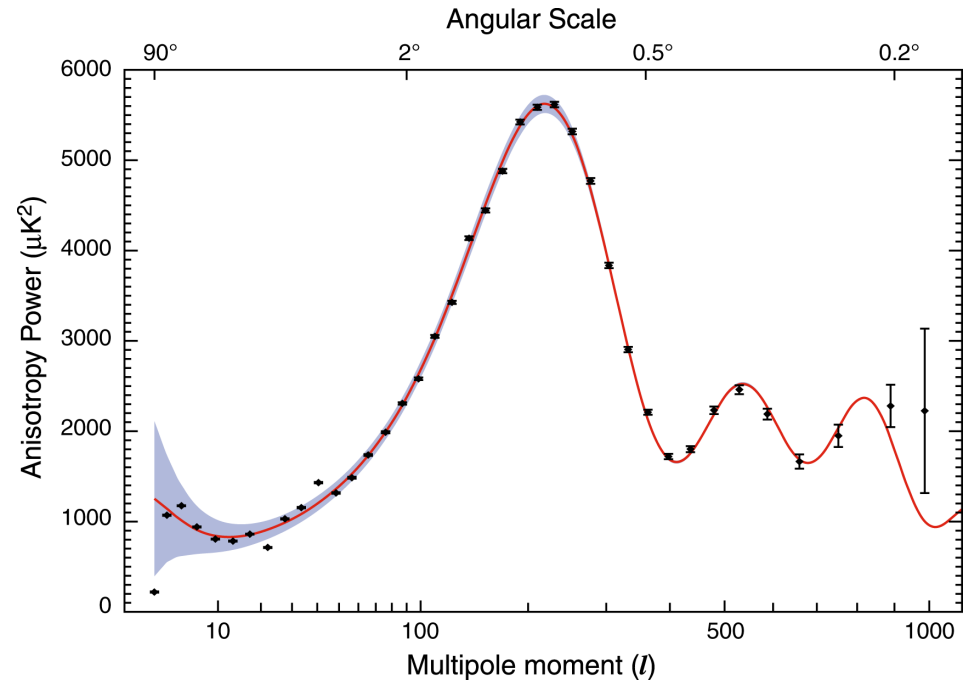
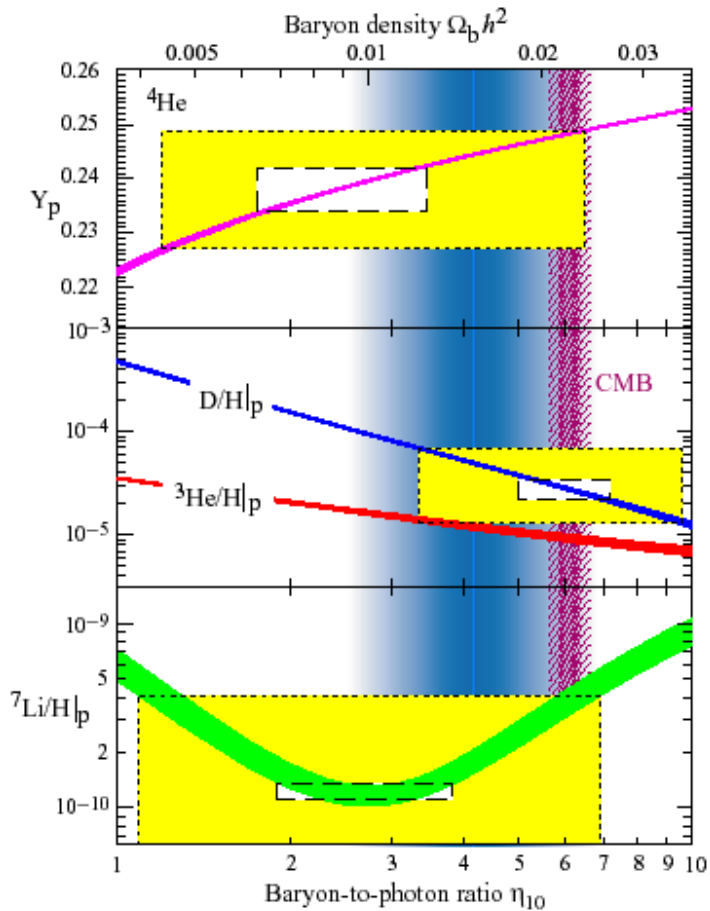
1. What is the origin of the tiny excess of matter over anti-matter?

2. What is the fundamental particle physics nature of Dark Matter?

The Matter Content of the Universe

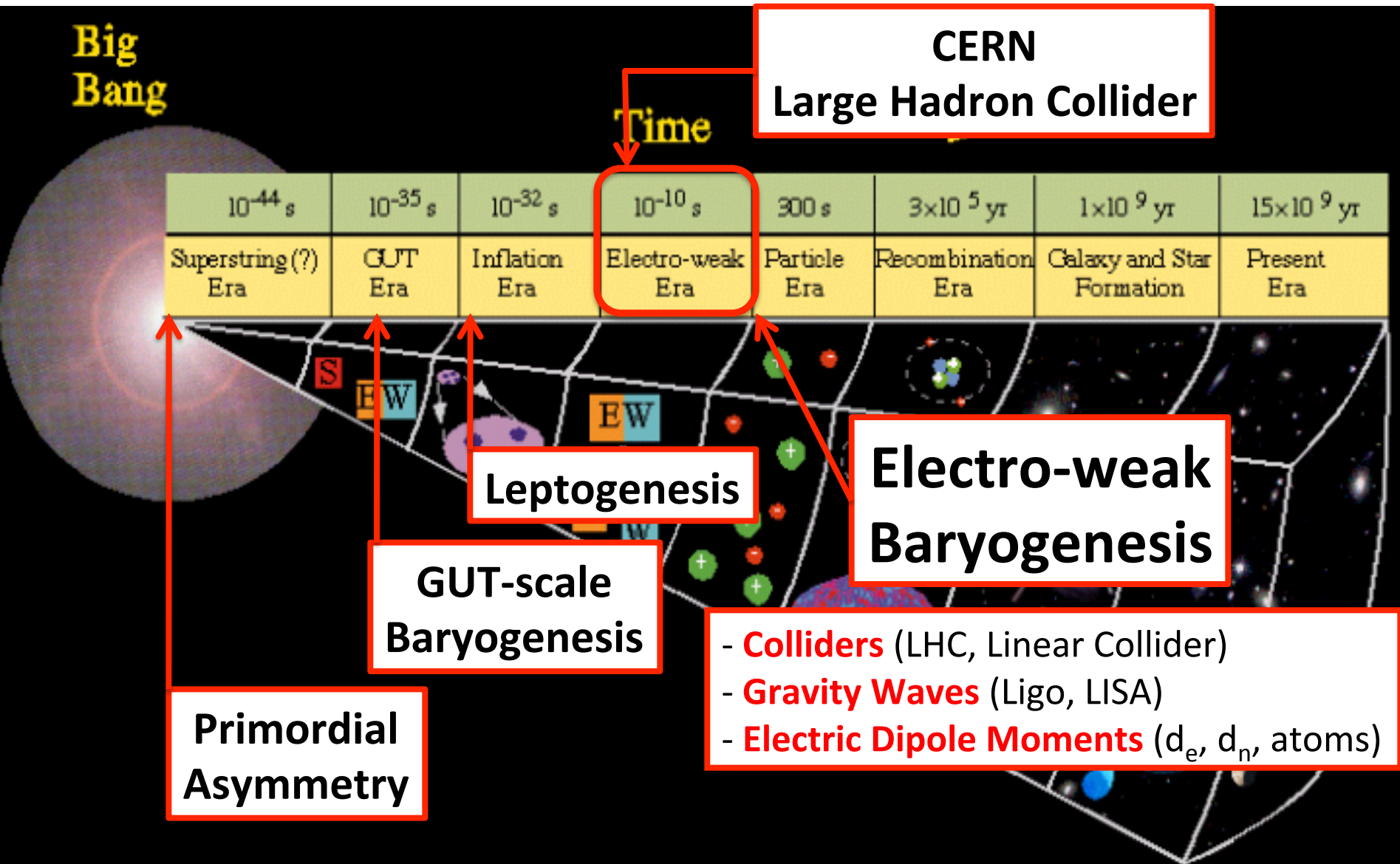


The Matter-Antimatter (Baryon) Asymmetry

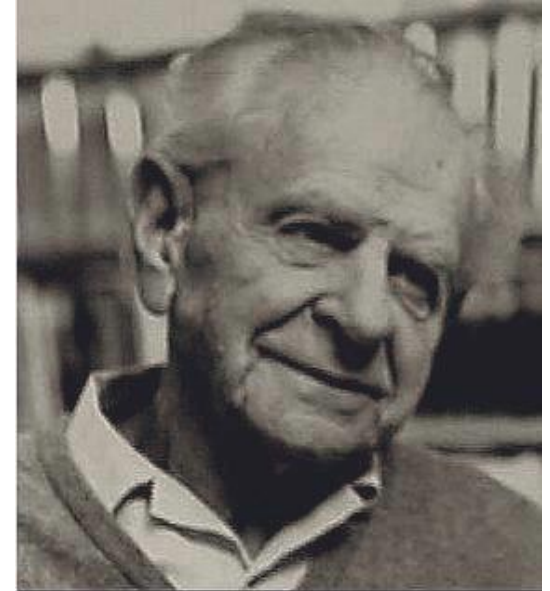


$$\frac{n_{\text{Baryons}} - n_{\text{Antibaryons}}}{n_\gamma} \approx 10^{-10}$$

No “Standard Model” of Baryogenesis!



“In so far as a scientific statement
speaks about **reality**,
it must be **falsifiable**:
And in so far as it is not falsifiable
it does not speak about reality”



Sir Karl Popper (1902-1994)

Karl Popper, “*Logik der Forschung*” (1934)
“*The Logic of Scientific Discovery*”

***(Supersymmetric) Electro-Weak Baryogenesis:
a falsifiable theory***

Ingredients of Baryogenesis

(1) **B**aryon Number violation

If B is conserved, the present BAU can only reflect asymmetric initial conditions

(2) **C** and **CP** violation

In the absence of a “preference” for matter or antimatter, B-nonconserving interactions will produce baryon and antibaryon excesses at the same rate: no net baryogenesis

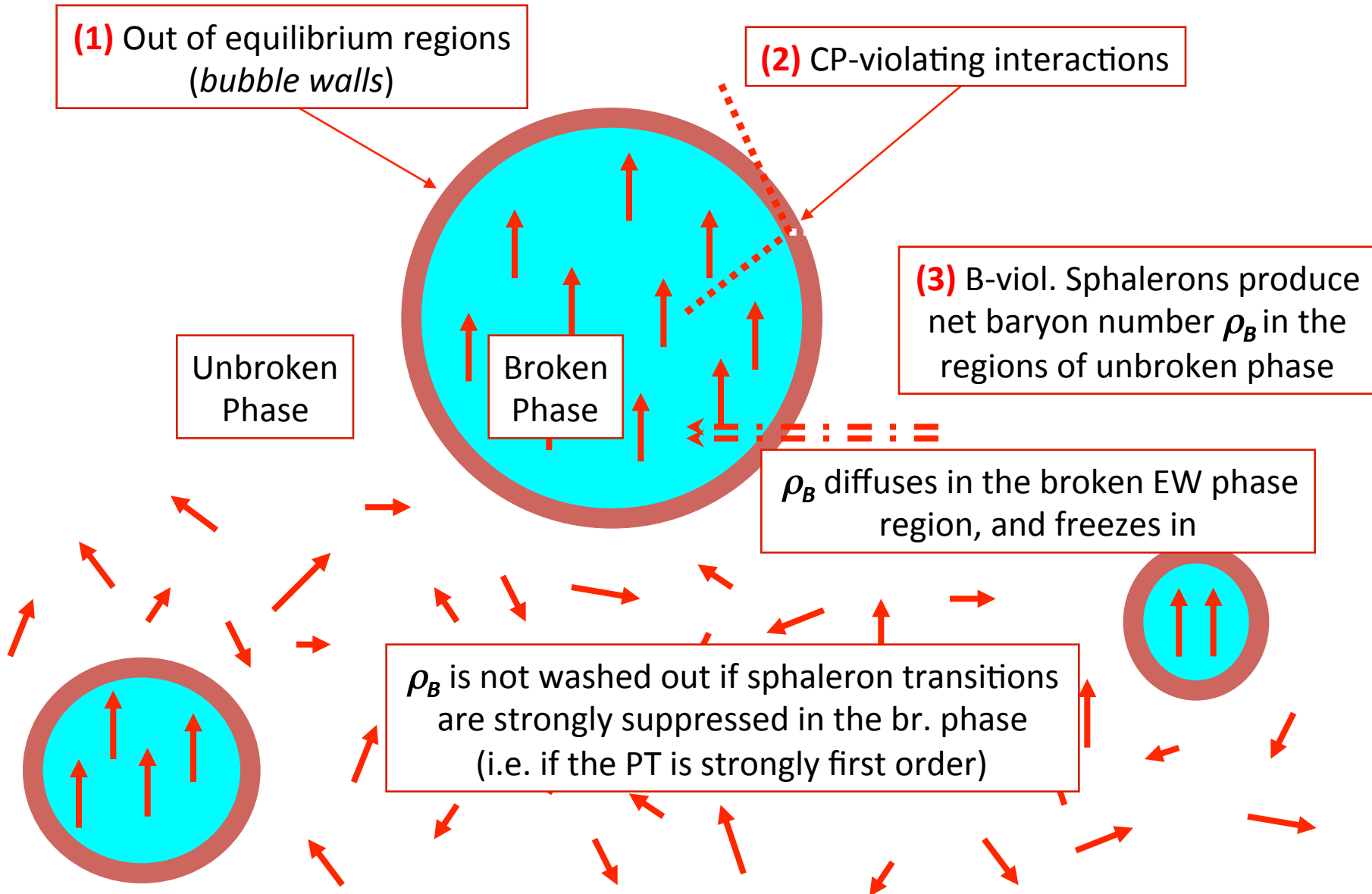
(3) **Out of Equilibrium** conditions

In chemical equilibrium the entropy is maximal when the chemical potential associated with all nonconserved quantum numbers vanishes

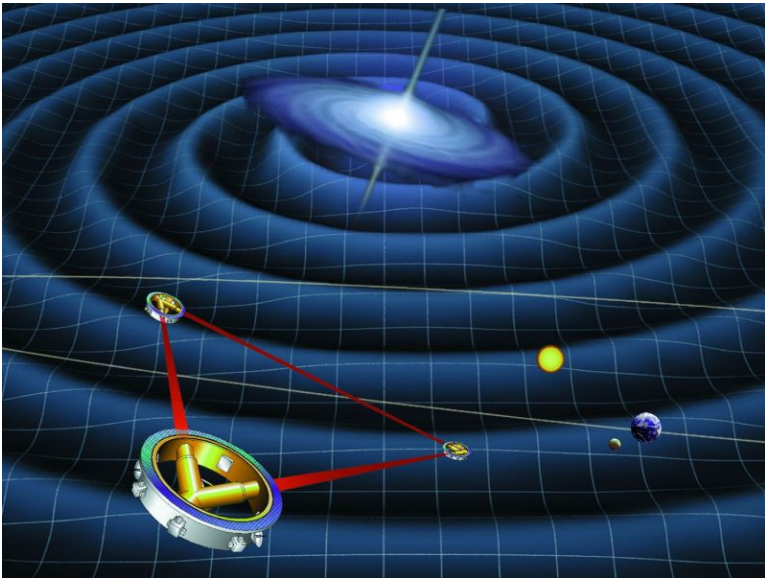
“**Sakharov conditions**”^(*)

^(*)A.D.Sakharov, JETP Letters **5**, 24 (1967)

Electro-Weak Baryogenesis

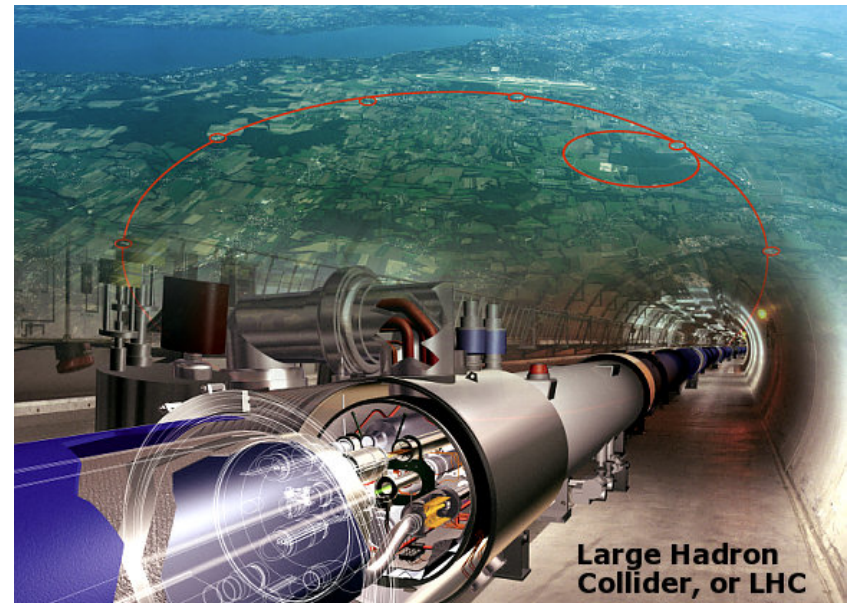
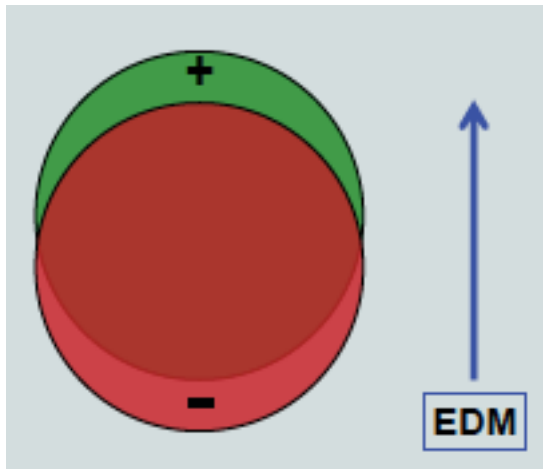


Experimental Tests of Electro-Weak Baryogenesis



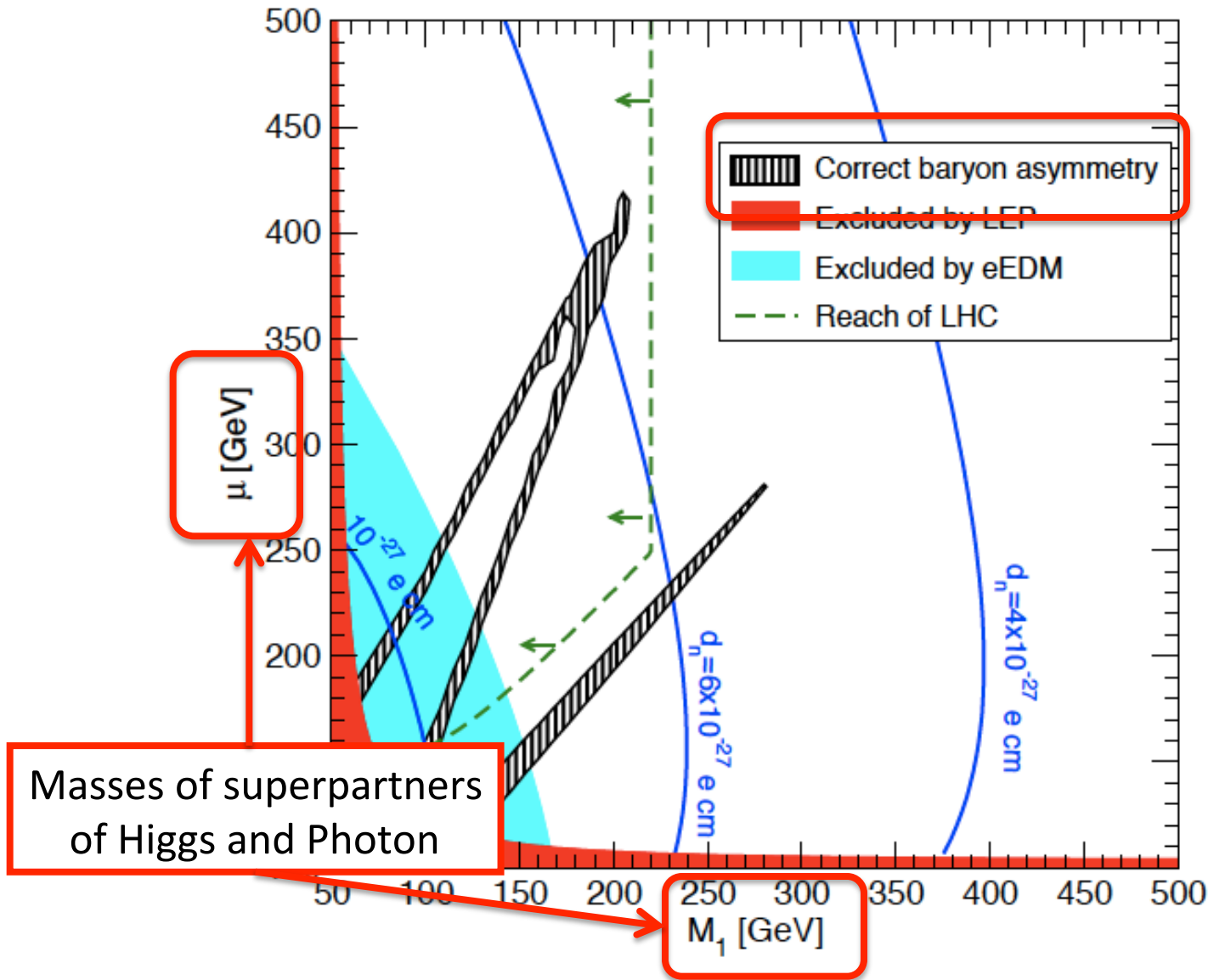
Gravity Waves from
Bubble Collisions

Large
Electric
Dipole
Moments
from CP
violation

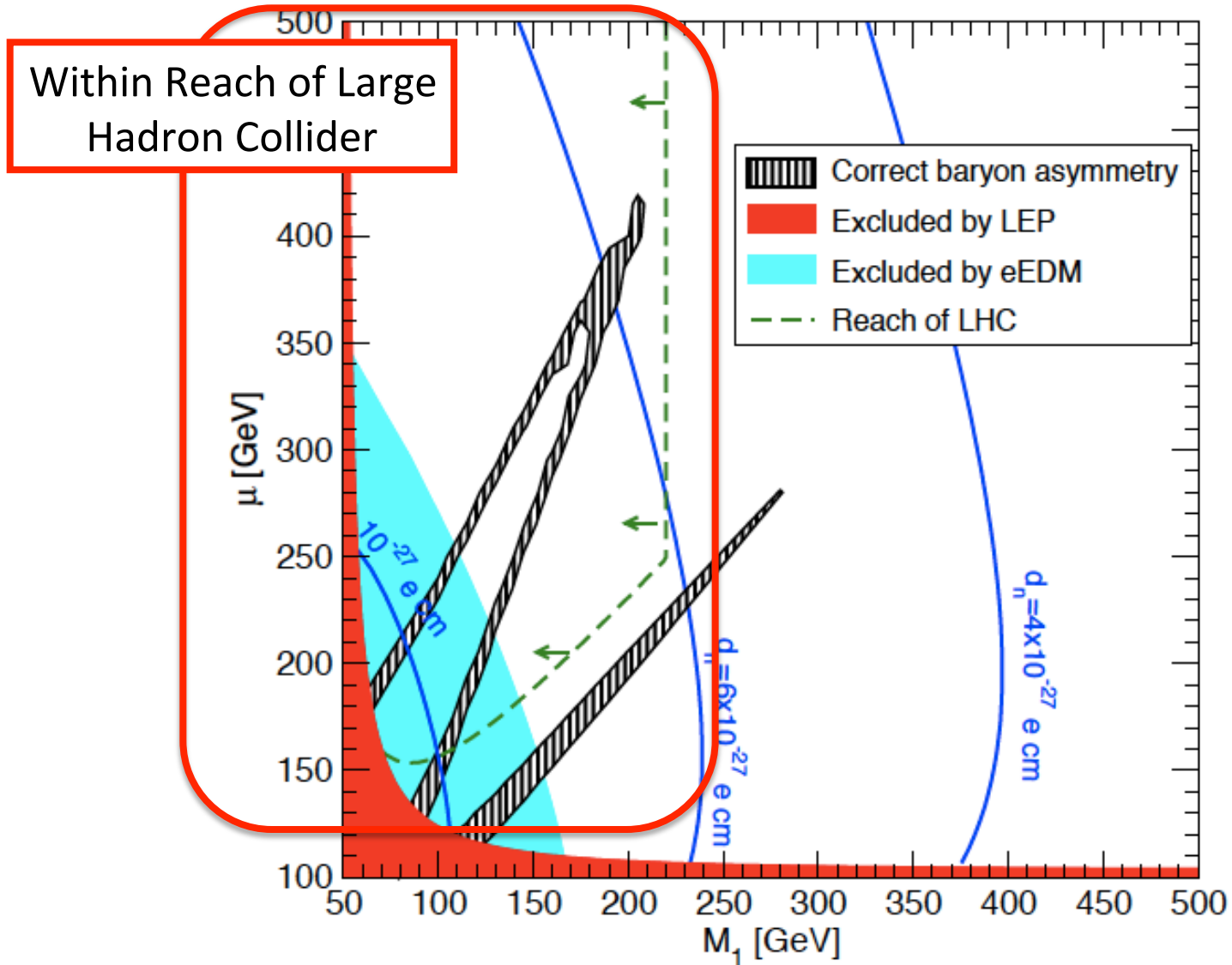


Light superpartners, distinctive
phenomenology at LHC

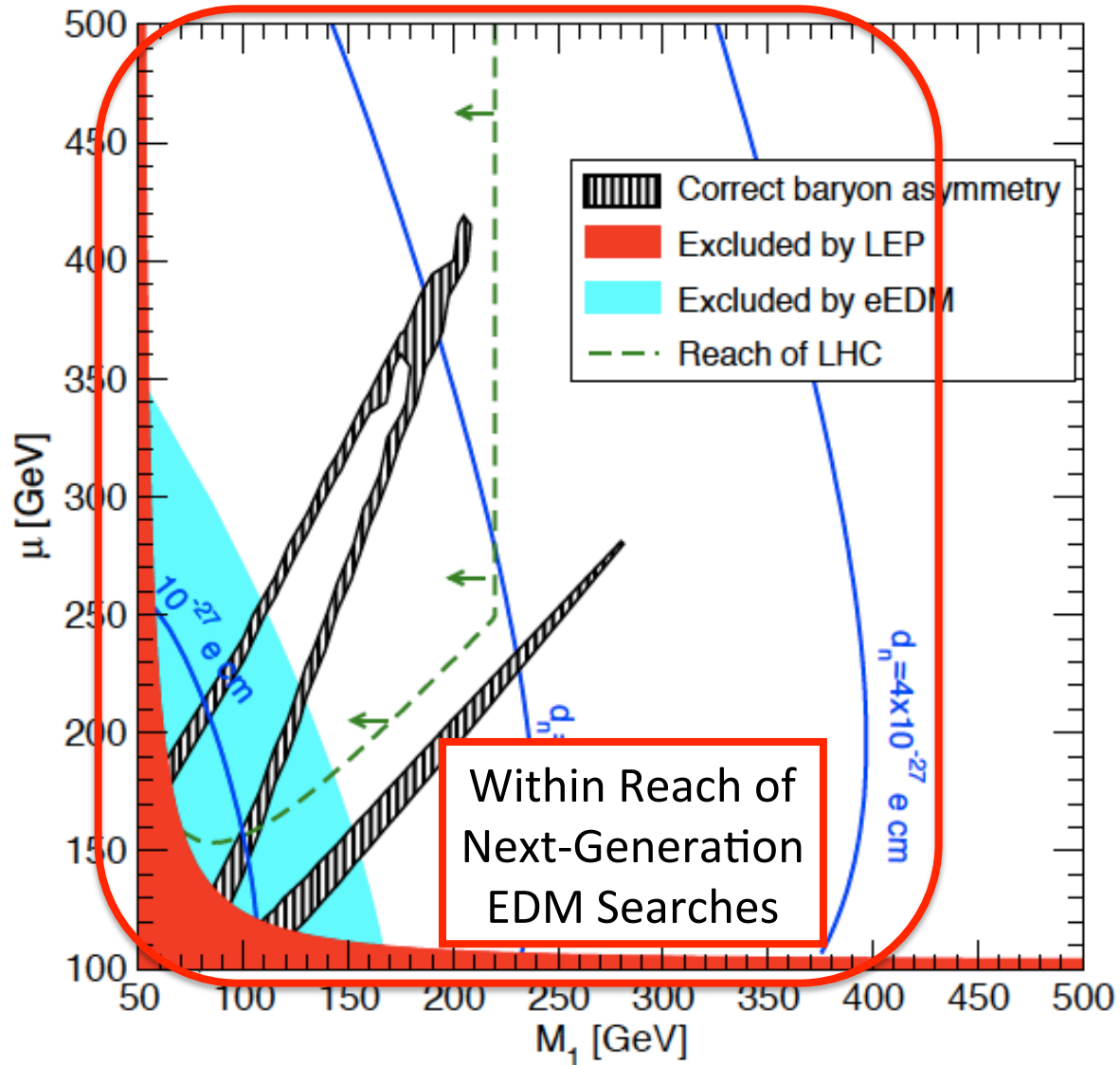
Electro-Weak Baryogenesis



Electro-Weak Baryogenesis

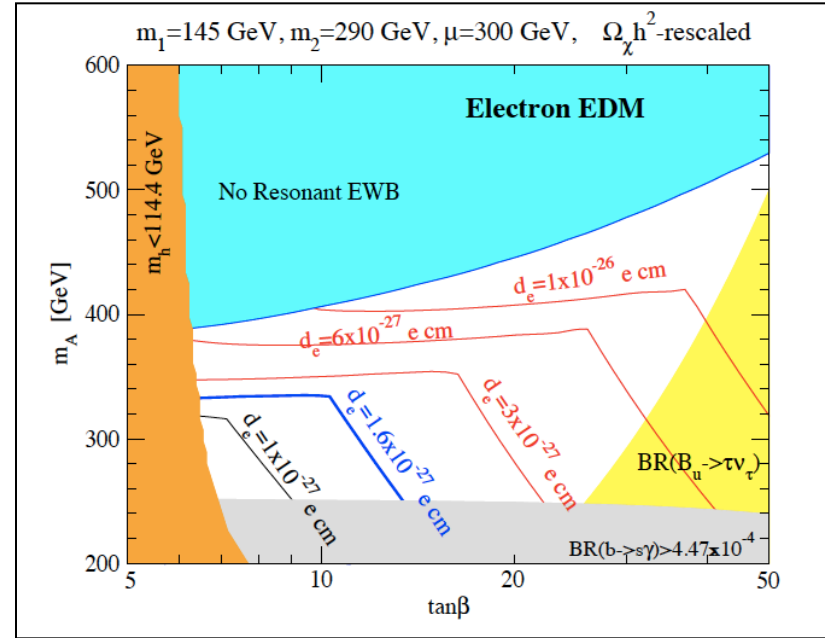


Electro-Weak Baryogenesis

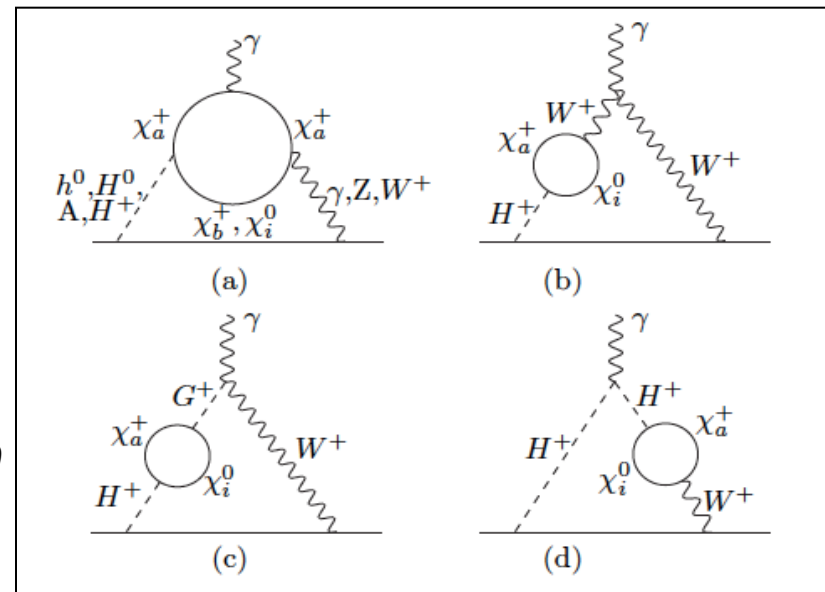


Electro-Weak Baryogenesis: Recent Progress

- Comprehensive Phenomenological Analysis of MSSM Electro-weak Baryogenesis (1)



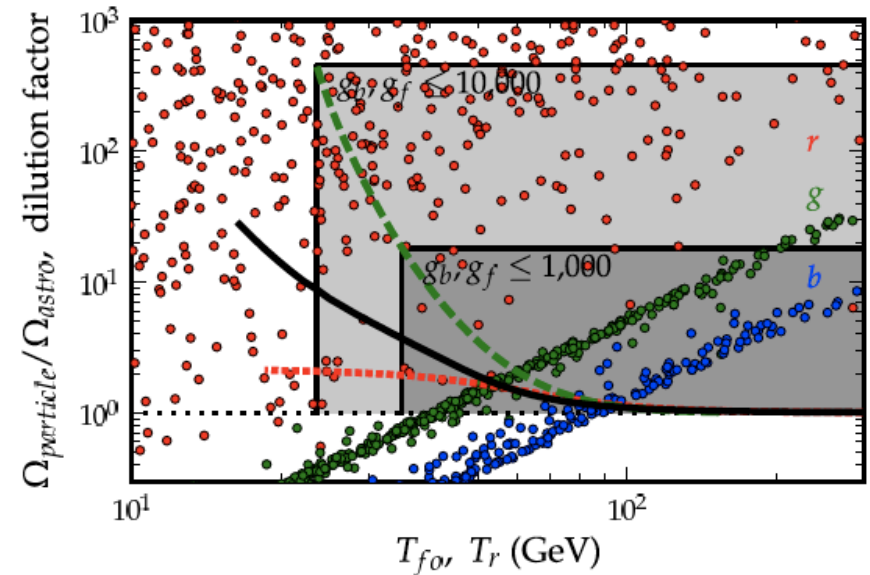
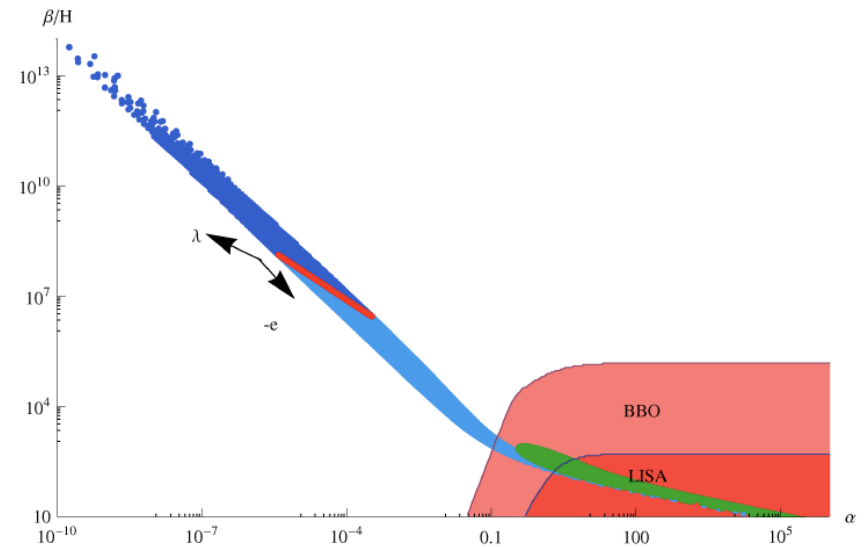
- Complete calculation of electro-weak 2-loop EDM amplitudes (2) + publicly available interface to numerical codes (2LEDM, 3)



- (1) Cirigliano, Li, Profumo, Ramsey-Musolf, JHEP 2010
 (2) Li, Profumo, Ramsey-Musolf, PRD 2008, PLB 2009
 (3) Li, Profumo, Ramsey-Musolf, JHEP 2010

Electro-Weak Baryogenesis: Recent Progress

- Gravity Waves as potential probes of EWB semi-analytic study (4)
- Impact of strongly first-order EW phase transition on thermal relics (5)

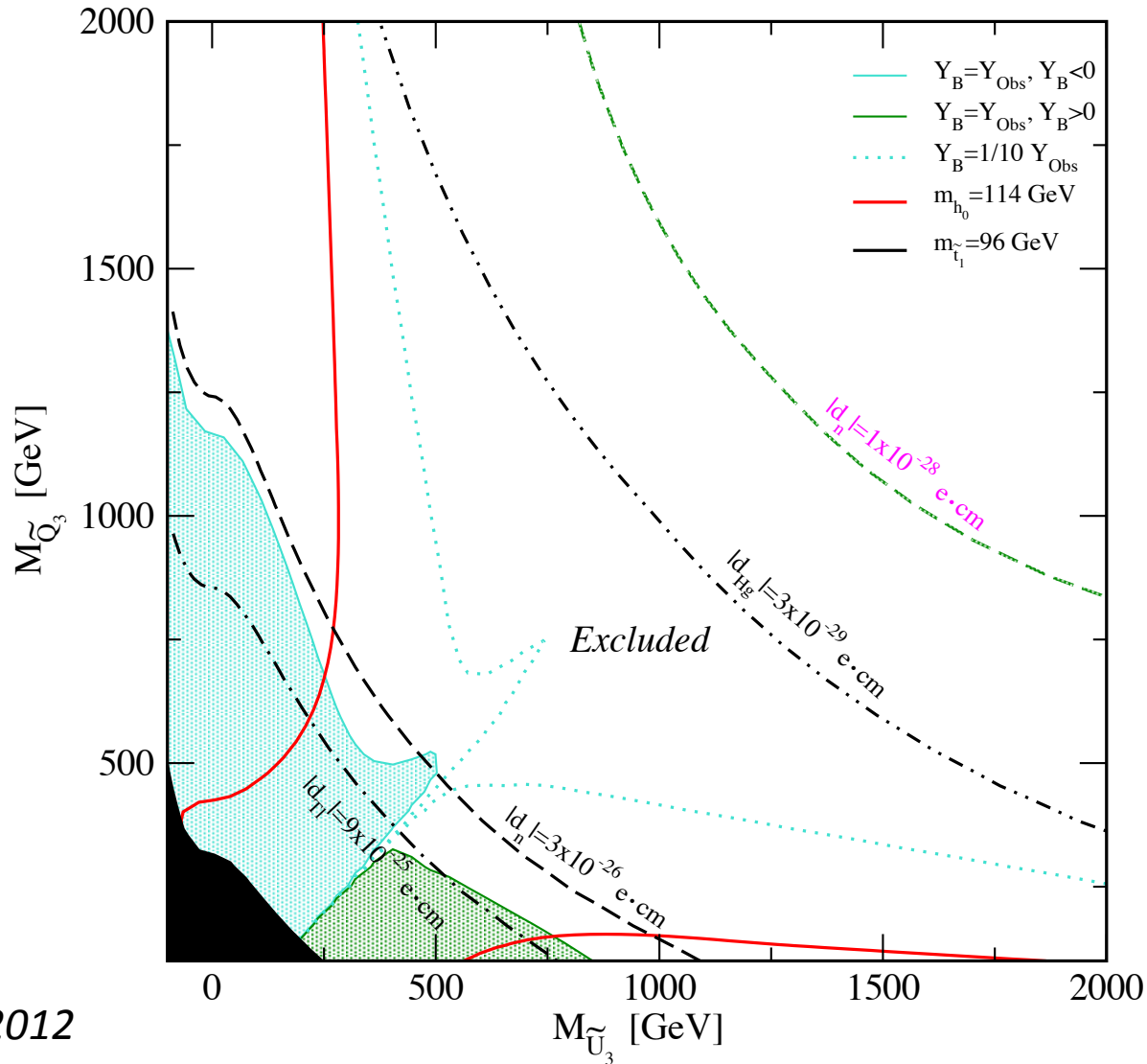


(4) Kehayias and Profumo, JCAP 2010

(5) Wainwright and Profumo, PRD 2009

Baryogenesis: Ongoing Work

$\tan \beta=10, A_t=250 \text{ GeV}, \mu=1000 \text{ GeV}$

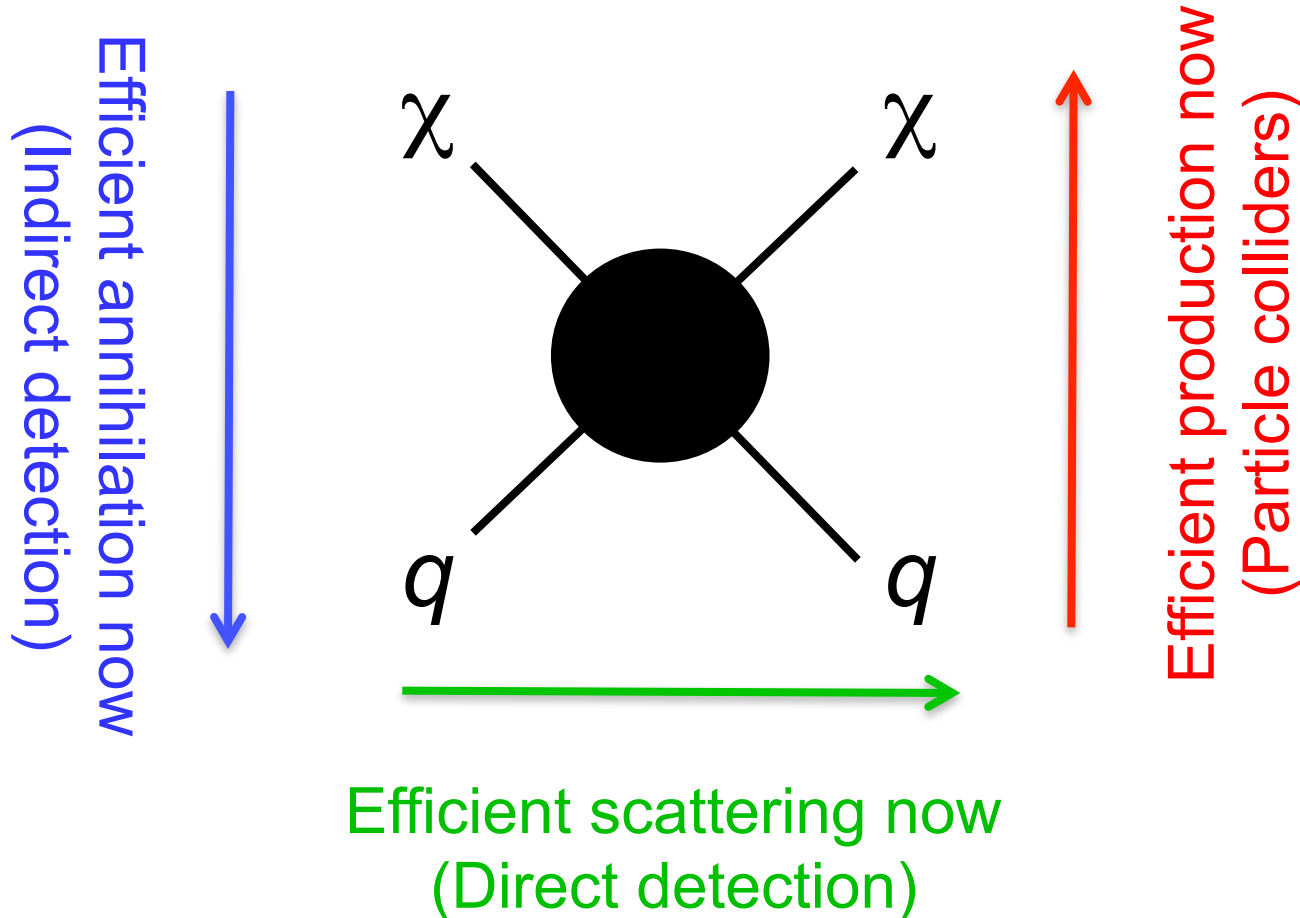


Kozaczuk,
Wainwright
and Profumo, 2012

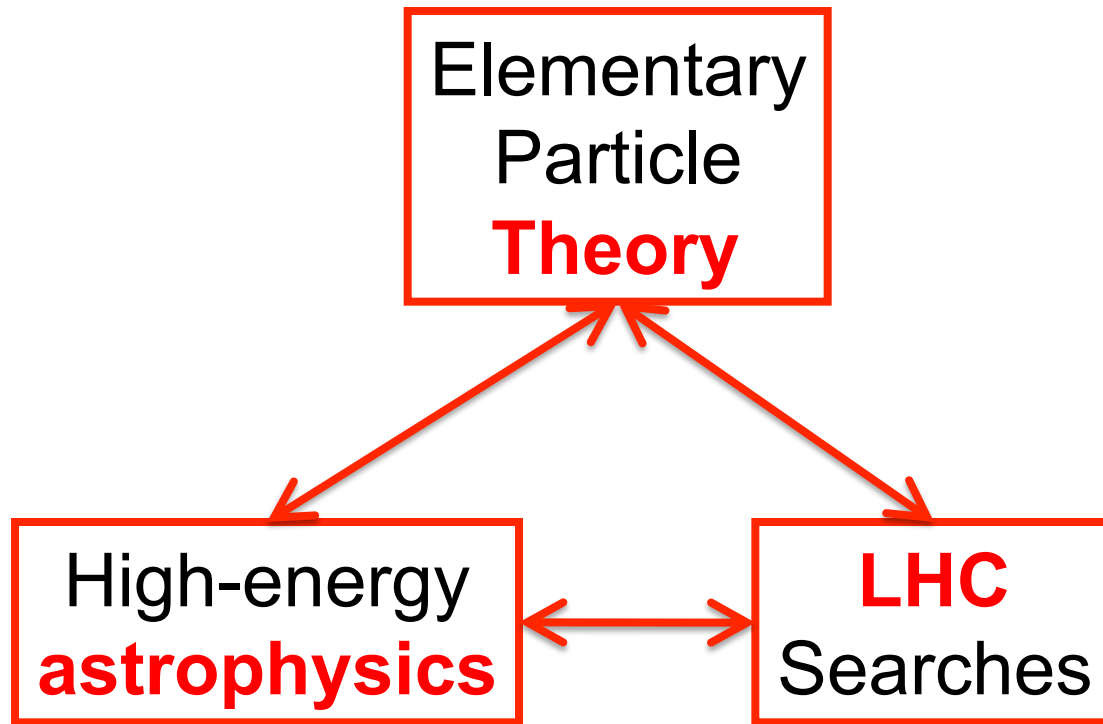
Baryogenesis: Ongoing Work

- New sources for electroweak baryogenesis (staus, sbottoms)
*(with **Wainwright** and **Kozaczuk**, 2012)*
- Analysis of issues of gauge invariance in finite-T effective potential
*(with **Wainwright** and Ramsey-Musolf, 2012)*
- Self-consistent models for Higgs, 130 GeV line, thermal DM in MSSM+singlet *(with **Wainwright** and **Kozaczuk**, 2013)*

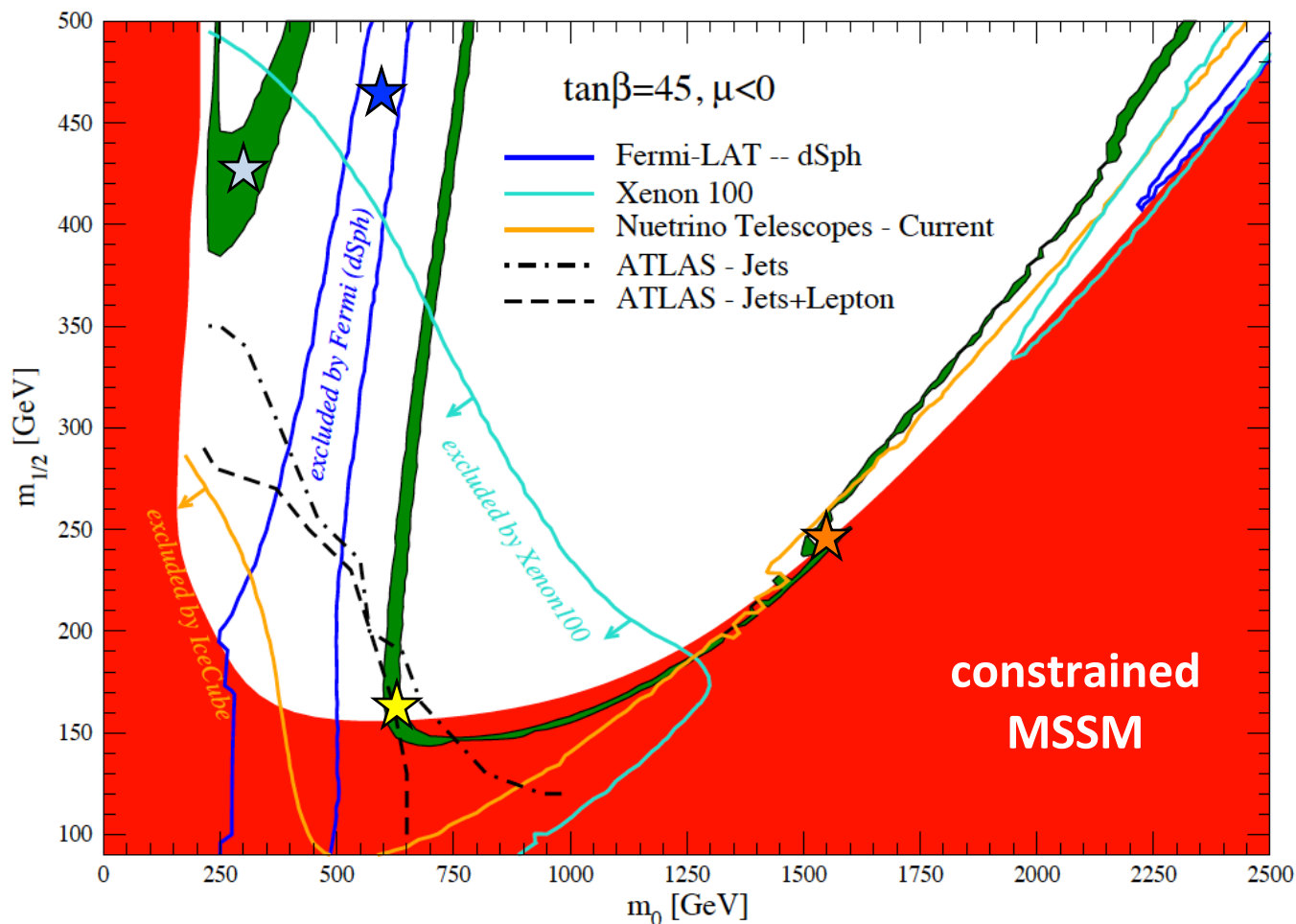
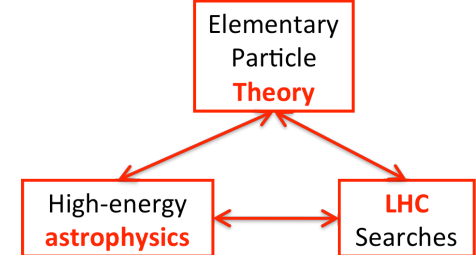
Particle Dark Matter: a multi-pronged approach in a **Time of Discovery**



Particle Dark Matter: a comprehensive approach for a cross-disciplinary science



- ★ LHC is probing cosmologically interesting regions
- ★ Only probed by Direct Detection
- ★ Only probed by Neutrino Telescopes
- ★ Only probed by Gamma-Ray Telescopes



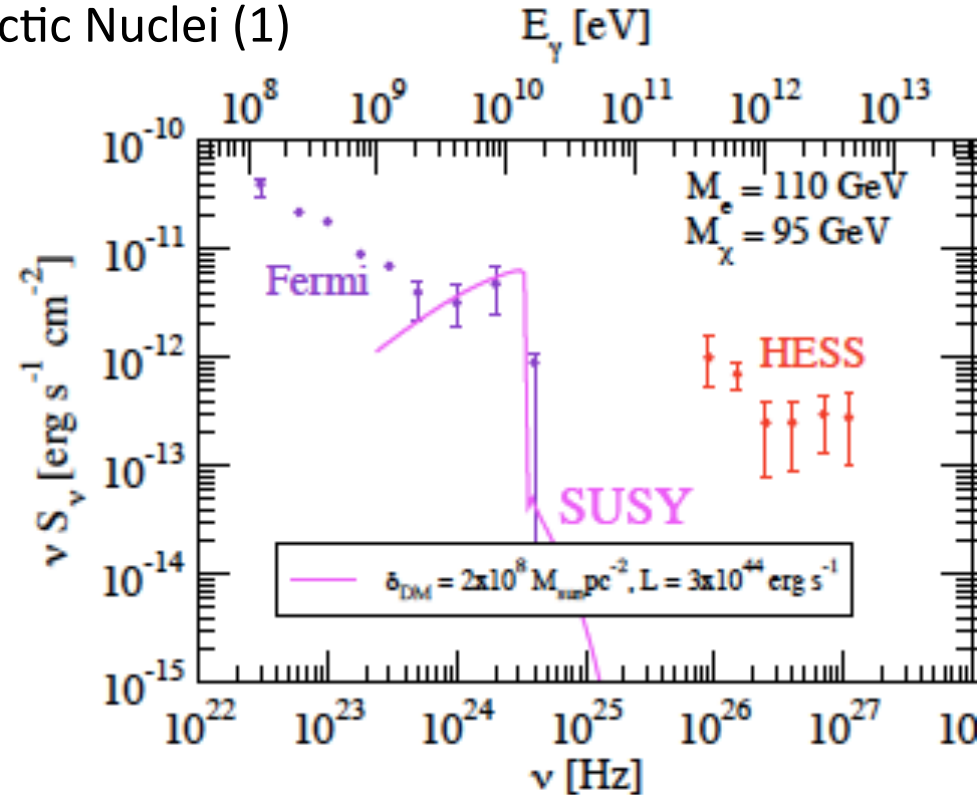
Particle Dark Matter: Theory

- **Electron-DM scattering** in Active Galactic Nuclei (1)

- Singlet dark matter models – **2-photon decay** amplitude, role of meta-stable vacua (2)

- Cosmology of **Saxions** in gauge mediation (3)

- Direct detection in lepton-specific **2 Higgs Doublet models** (4)



- (1) Ubaldi, Gorchteyn and Profumo, PRD
- (2) Ubaldi, Waiwright and Profumo, PRD
- (3) Dine, Profumo and Ubaldi, in preparation
- (4) Boucenna, Cornell and Profumo

Particle Dark Matter: LHC Physics

- Searches for **2UED** with LHC; ATLAS code development (**FeynRules -> Madgraph**) (1)

- Vertex $\{WH11i, 1\}$, $\{WH11i, 2\}$, $\{Wmu11i, 3\}$, $\{W$

$$-\frac{3}{4}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_3,\mu_4} - \frac{3}{4}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_3,\mu_4}$$

- Vertex $\{Wmu11i, 1\}$, $\{Wmu11i, 2\}$, $\{Wmu11i, 3\}$, $\{$

$$-\frac{9}{2}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_1,\mu_4}\eta_{\mu_2,\mu_3} + \frac{9}{4}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_1,\mu_3}\eta_{\mu_2,\mu_4} - \frac{9}{4}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_1,\mu_3}\eta_{\mu_2,\mu_4} - \frac{9}{2}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_1,\mu_3}\eta_{\mu_2,\mu_4} + \frac{9}{4}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_1,\mu_2}\eta_{\mu_3,\mu_4} + \frac{9}{4}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_1,\mu_2}\eta_{\mu_3,\mu_4}$$

- **Multi-top** (4x, 6x, 8x) LHC searches (2)

- Discriminating **axinos** and **gravitinos** at LHC (3)

(1) Manning and Profumo, in progress

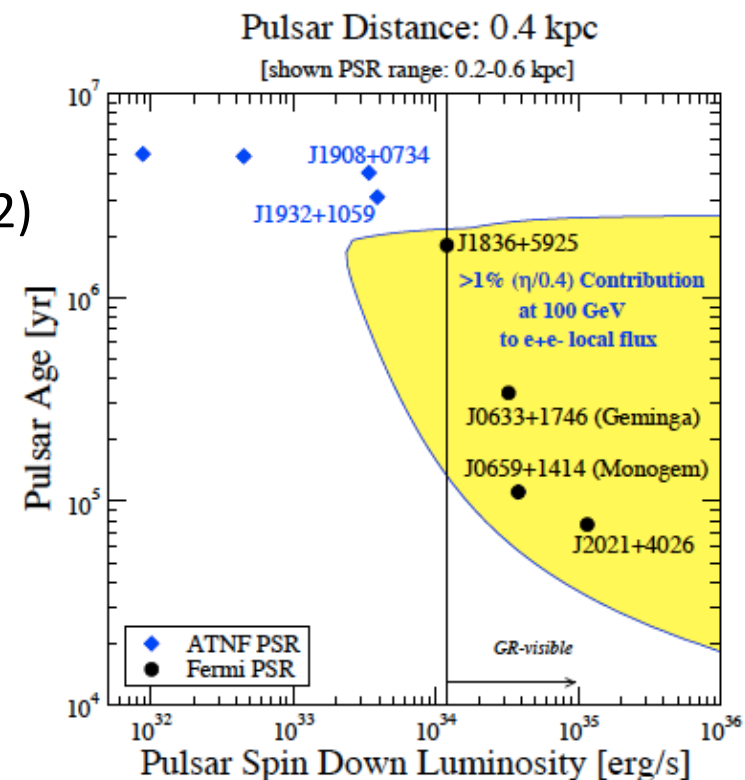
(2) Manning, Profumo and Seiden, in progress

(3) Lee, Profumo and Ubaldi, in progress

Particle Dark Matter: Astrophysical Backgrounds

The field does not seem to need theory model building as much as **reliable assessment of signals**, and hence of the relevant **astrophysical backgrounds**

- **Pulsars** as sources of high-energy positrons – connection to Fermi gamma-ray pulsars (1)
- **Milli-second pulsars** and diffuse gamma rays (2)
- Systematic effects in signals from diffuse emissions (“**Fermi Haze**”)(3)
- Hadronic models for **cluster radio halos** (4) – relevant for **Cosmology** !!!



(1) Gendeleev, Profumo and Dormody, JCAP 2010

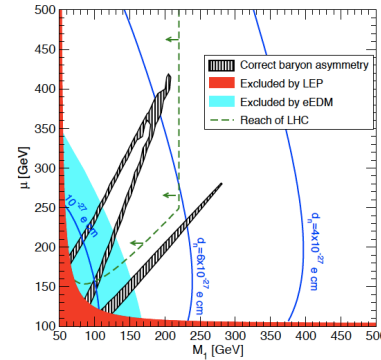
(2) Siegal-Gaskins, Profumo et al, MNRAS 2011

(3) Linden and Profumo, APJL 2011

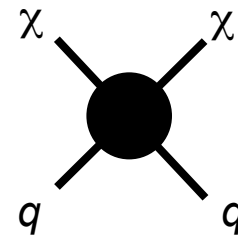
(4) Jeltama and Profumo, Astroph. J. 2010

Summary

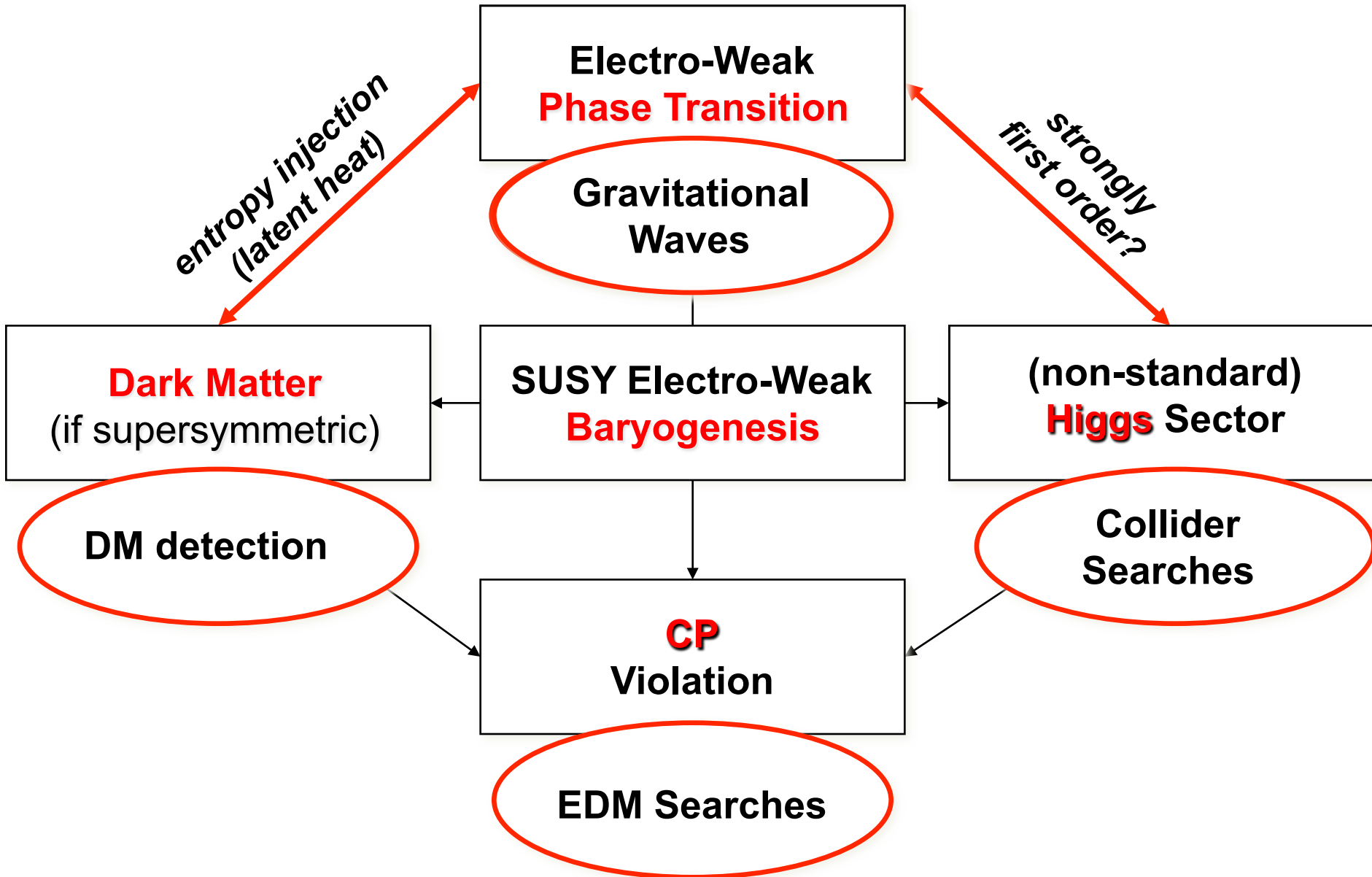
Testable theories for
the **origin of matter**



A **cross-disciplinary approach**
in the hunt for **dark matter**



Electro-Weak Baryogenesis: Probes



Ingredients of Baryogenesis

- **B**aryon Number violation

If B is conserved, the present BAU can only reflect asymmetric initial conditions

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“**Sakharov conditions**”^(*)

^(*)A.D.Sakharov, JETP Letters **5**, 24 (1967)

Electro-Weak Baryogenesis

The Electro-Weak Phase Transition fulfills all
3 Sakharov requirements^(*) (**Electro-weak Baryogenesis**)

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Electro-Weak Baryogenesis

The Electro-Weak Phase Transition fulfills all
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✓ **B** violation: Weak **Sphaleron** Transitions

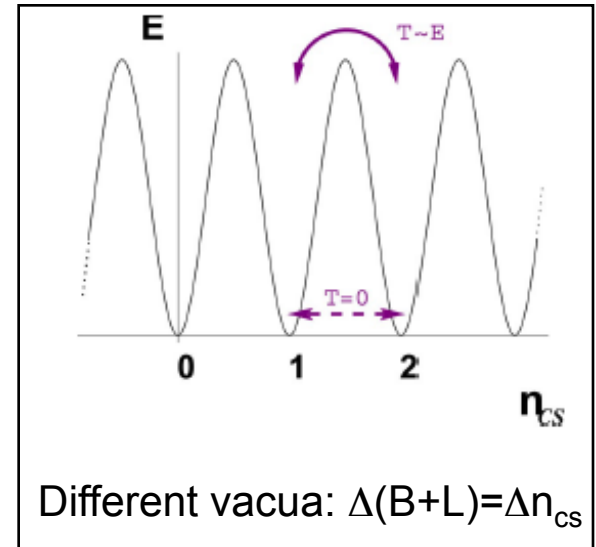
^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Electro-Weak Baryogenesis

The Electro-Weak Phase Transition fulfills all 3 Sakharov requirements^(*) (**Electro-weak Baryogenesis**)

✓ **B** violation: Weak **Sphaleron** Transitions

- Classically, baryonic and leptonic currents are **conserved** in the EW theory
- Quantum corrections produce **anomalous transitions** between non-degenerate $SU(2)$ field configurations vacua that **violate $B+L$** (but preserve $B-L$)
- **B-violation** rate is **unsuppressed** at $T > T_c$, and is **exponentially suppressed** at $T < T_c$



$$\Gamma_{sph} \propto \alpha_W T^4$$

$$\Gamma_{sph} \propto \exp[-E_{sph}(T)/T]$$

$$E_{sph}(T) \propto \langle \phi \rangle(T)$$

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

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^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

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✓ **Out of Equilibrium**: **Bubble Walls** of broken EW phase

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

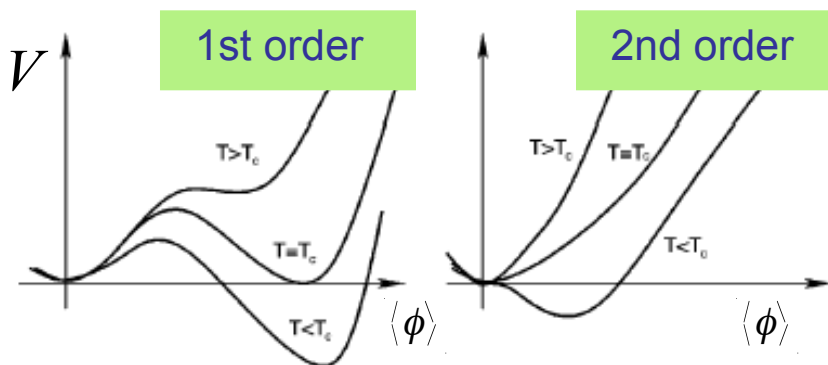
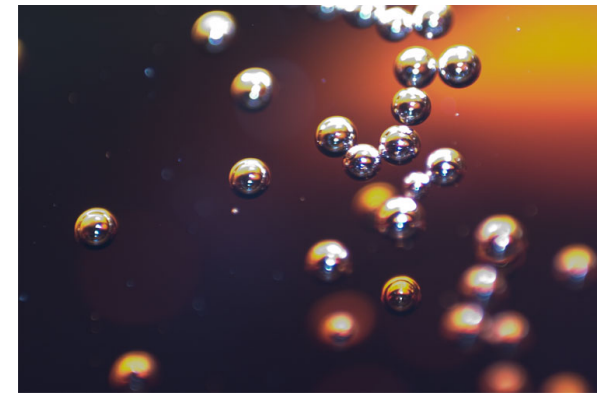
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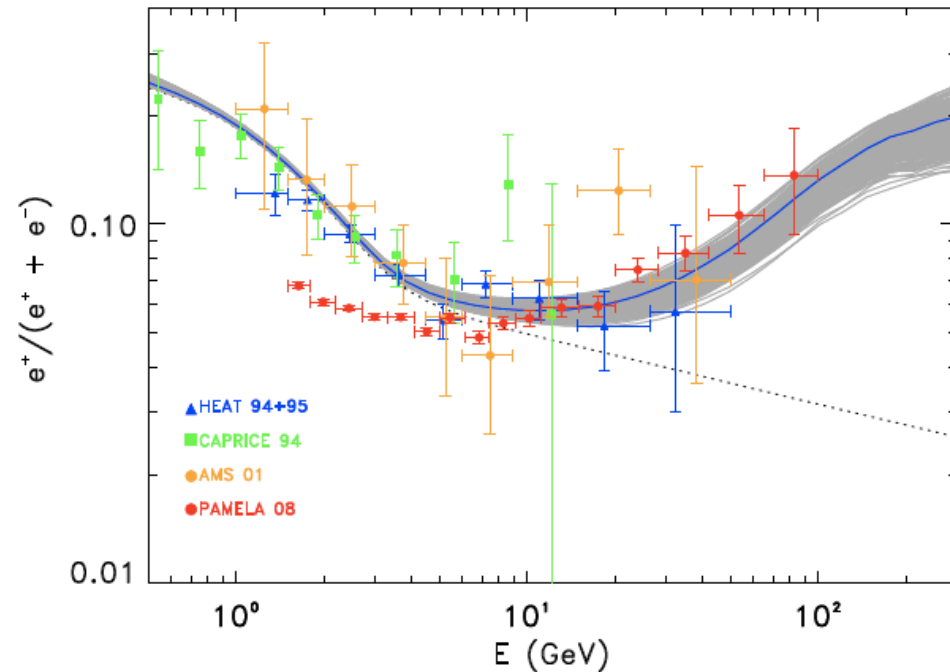
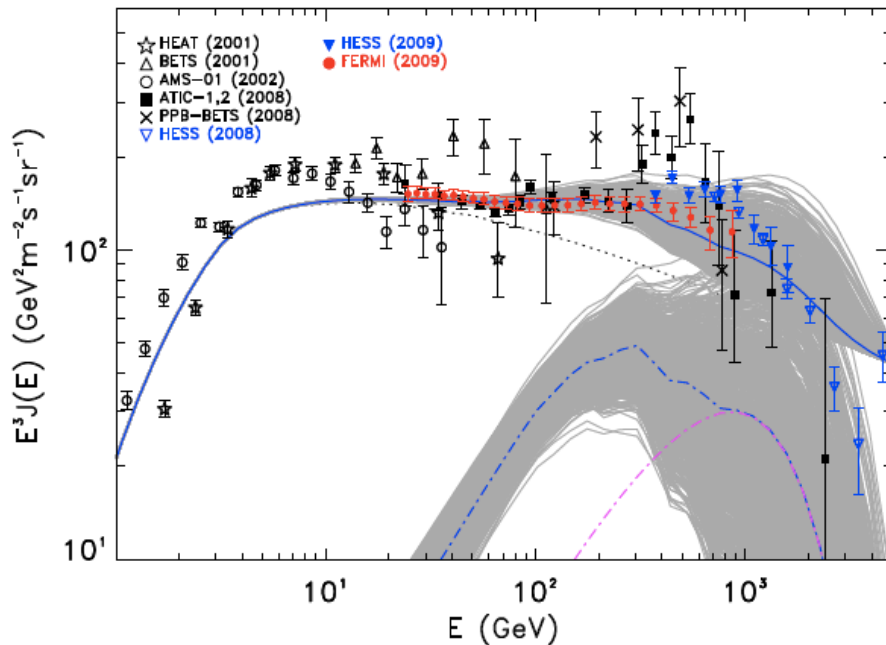
- If the EWPT is first order (cubic term!), it proceeds through Bubble nucleation
- The (expanding) Bubble Walls are out of thermal equilibrium

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Particle Dark Matter: Indirect Detection and Theory

■ Cosmic Rays

- Interpretation of Fermi electron/positron data (1)
- Fermi pulsars vs excess positrons (2)
- Dark Matter vs excess positrons (3,4,5)



(1) Grasso, Profumo, Strong et al, *Astropart.Phys.* 2009

(2) Gendeleev, Profumo and Dormody, *JCAP* 2010

(3) Jeltema and Profumo, *JCAP* 2009

(4) Brun, Delahayie, Diemand, Profumo, *PRD* 2009

(5) Cyr-Racine, Profumo and Sigurdson, *PRD* 2009

Particle Dark Matter: Indirect Detection and Theory

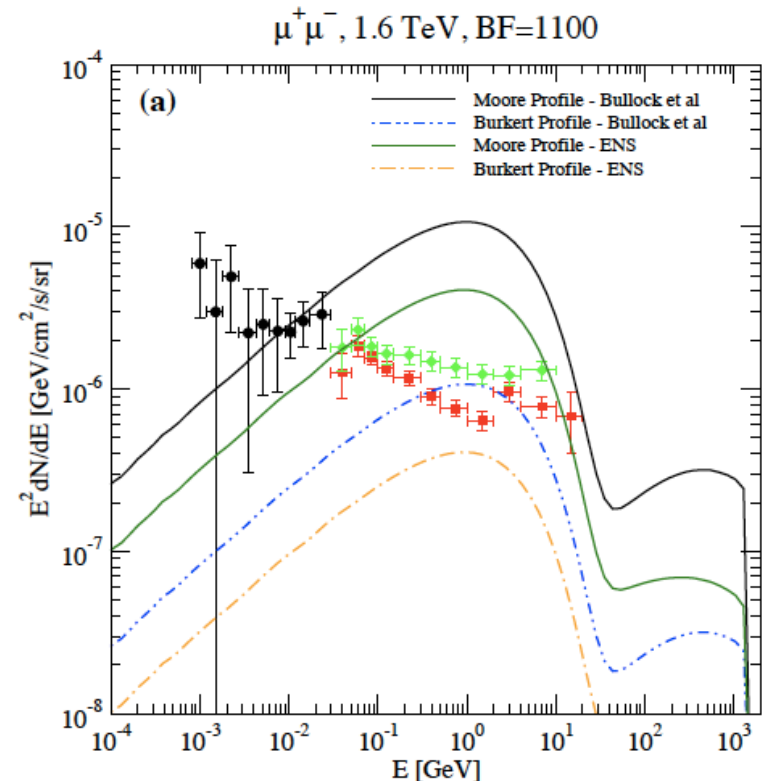
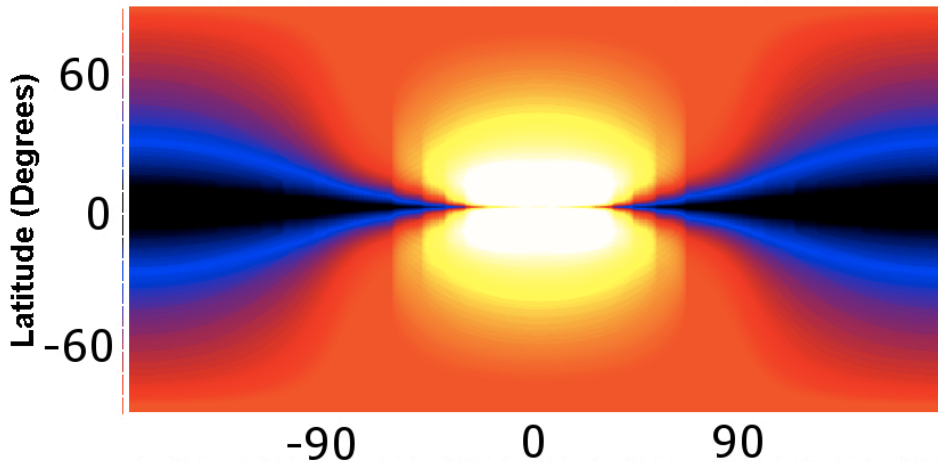
Diffuse Emissions

- WMAP Haze with consistent CR models (1)
- A Fermi haze? Systematic effects (2)
- Extragalactic Inverse Compton (3)

Theory/Multi-disciplinary

- Multi-component Dark Matter models: direct, indirect and collider searches (4)

3.55 GeV



(1) Linden, Profumo and Anderson, PRD 2010 sub.

(2) Linden and Profumo, Astroph. J. Lett. 2010

(3) Jeltama and Profumo, JCAP 2009

(4) Profumo, Sigurdson and Ubaldi, PRD 2009