

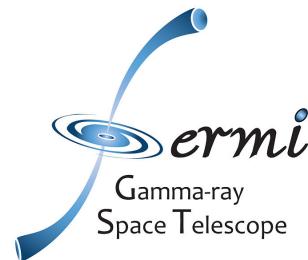


SCIPP

SANTA CRUZ INSTITUTE FOR PARTICLE PHYSICS



UC SANTA CRUZ



Stefano Profumo

UC Santa Cruz

Santa Cruz Institute for Particle Physics

Affiliated Member, Fermi Collaboration

*Terascale: key to the origin of
baryonic and dark matter?*

UCSC - Physics 205

Monday February 13, 2012

- ✓ PhD **Theoretical Particle Physics** (2004)
International School for Advanced Studies (SISSA-ISAS), Trieste, Italy
- ✓ Postdoc, FSU and California Institute of Technology (2005-2007)
Theoretical Astrophysics and Particle Physics
- ✓ Joined **UCSC Physics** Faculty (Assistant Professor, 2007-2011,
Associate Professor, July 2011-)
- ✓ Research funded by Department of Energy (Outstanding Junior
Investigator Award), National Science Foundation, NASA
- ✓ **Fermi**-LAT affiliated scientist; collaborate with **ATLAS** group
- ✓ **SCIPP Deputy Director** for **Theory** (July 2011-)

Particle Theory Group
Banks, Dine, Haber

Profumo's Research Group

Fermi Telescope Group

Atwood, Johnson, Ritz

ATLAS Group

Nielsen, Schumm, Seiden

Postdocs

Draper
Shepherd (will join July 1)

Graduate Students

Kehyias*
Ubaldi**
Linden***
Wainwright***

Cornell
Carlson
Storm (with Jeltema)
Kozaczuk (with Aguirre)
Manning (with Seiden)

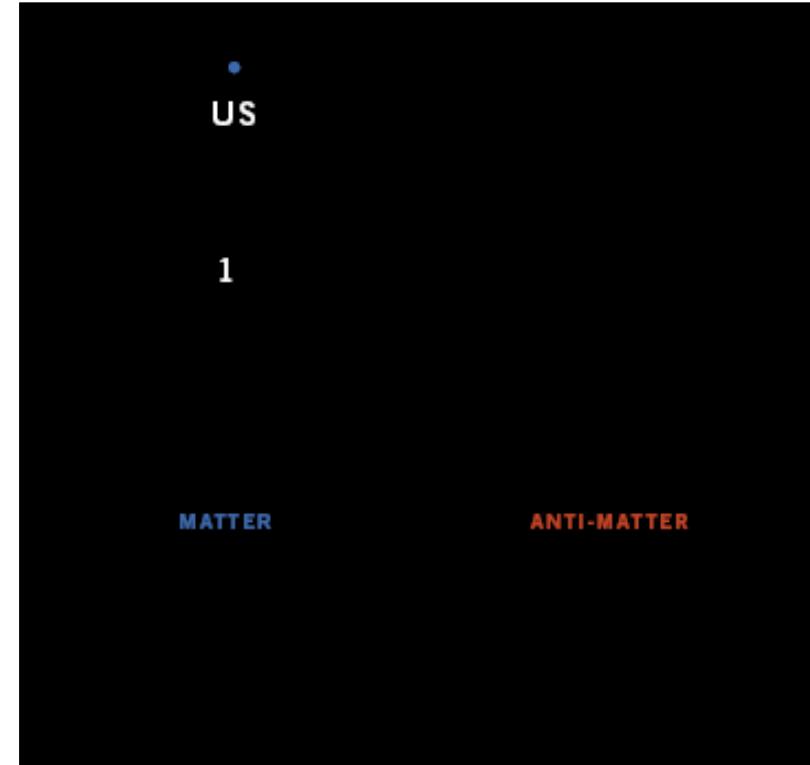
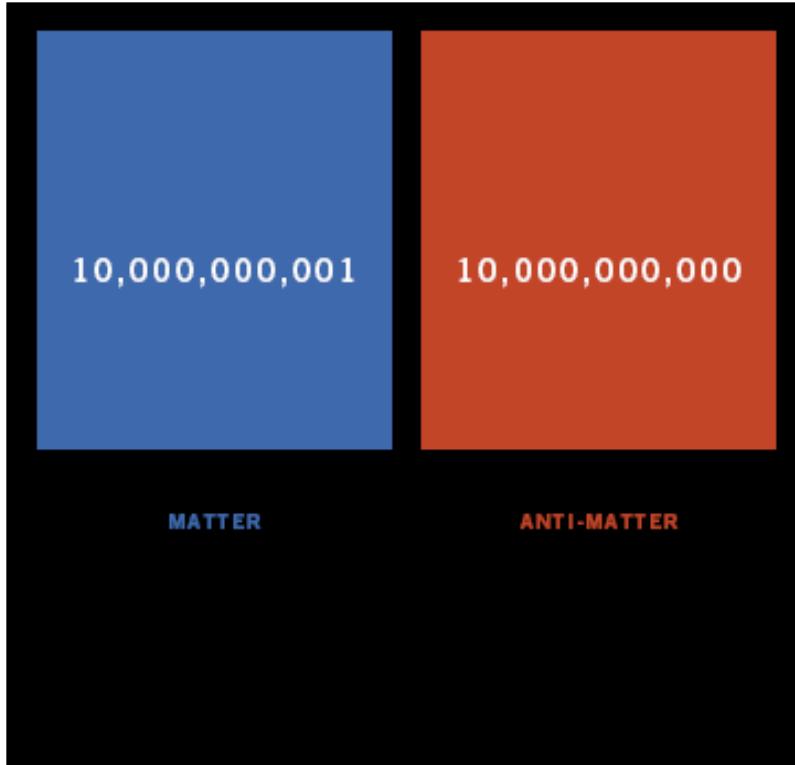
Undergraduate Students

Stopnitzky, Bergeron, Gruzdas

* graduated in 2011, postdoc at University of Tokyo, IPMU

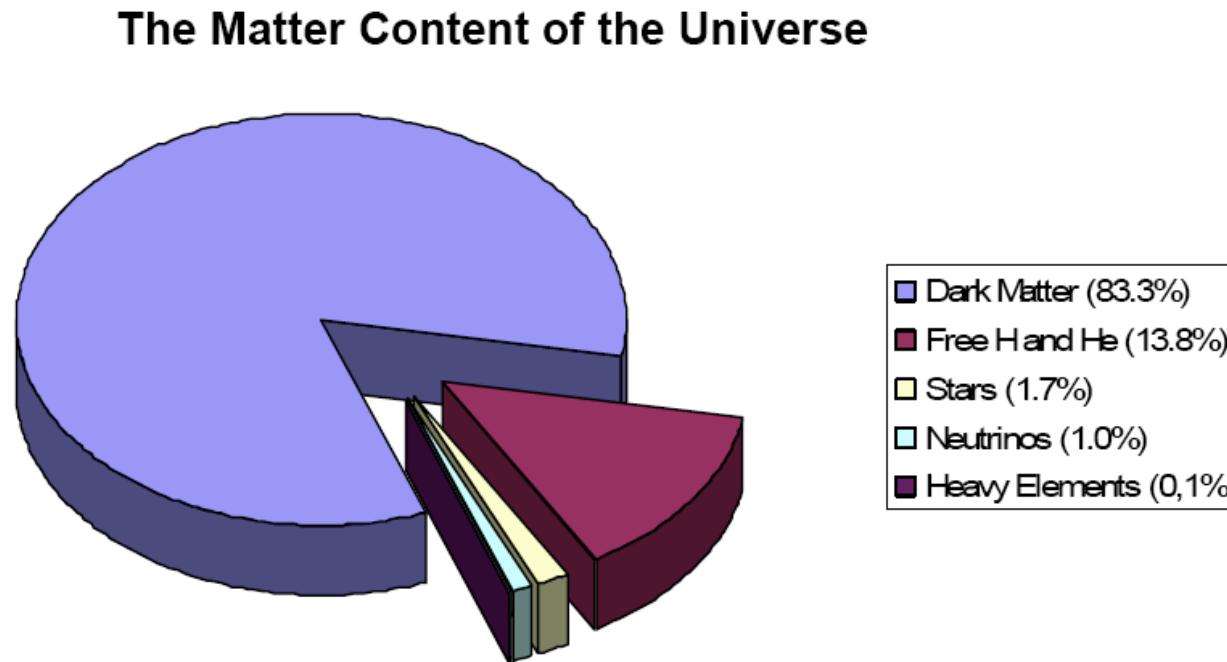
** graduated in 2011, postdoc position at Bonn University

*** will graduate in 2013

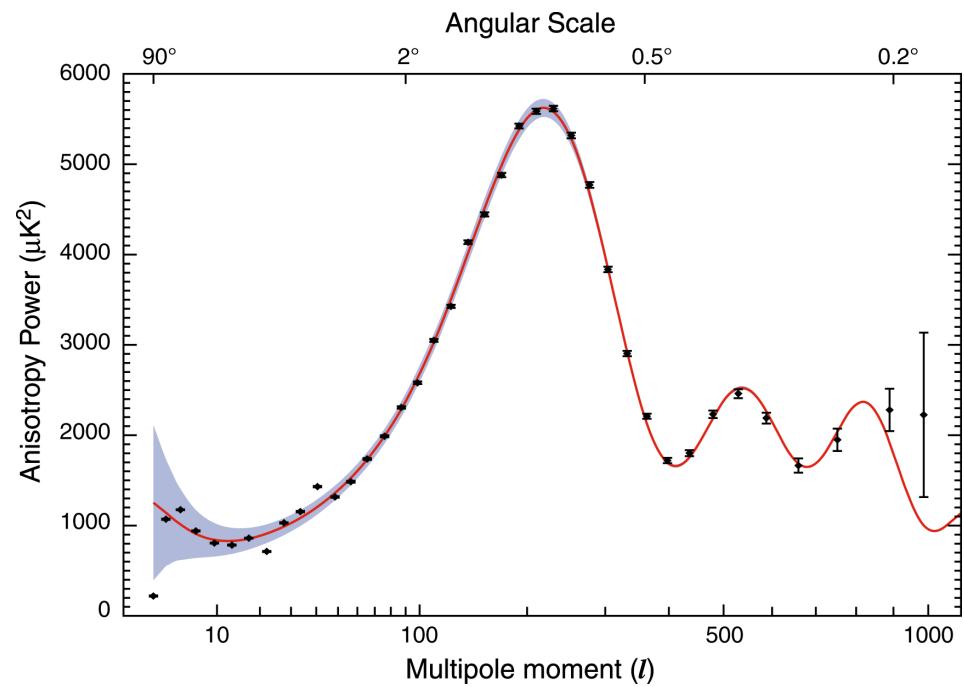
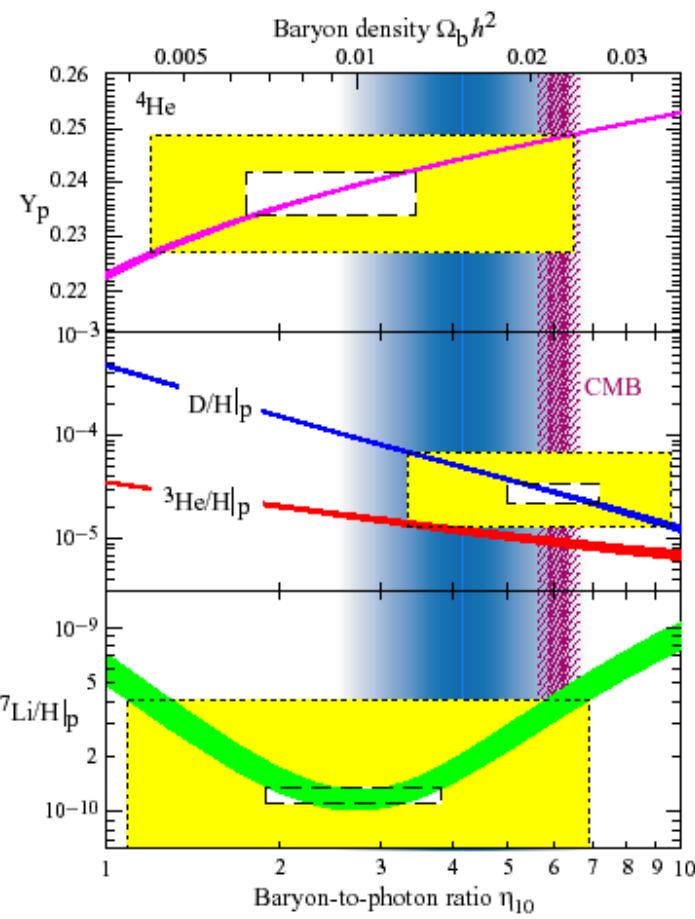


1. What is the origin of the tiny excess of matter over anti-matter?

2. What is the fundamental particle physics nature of Dark Matter?

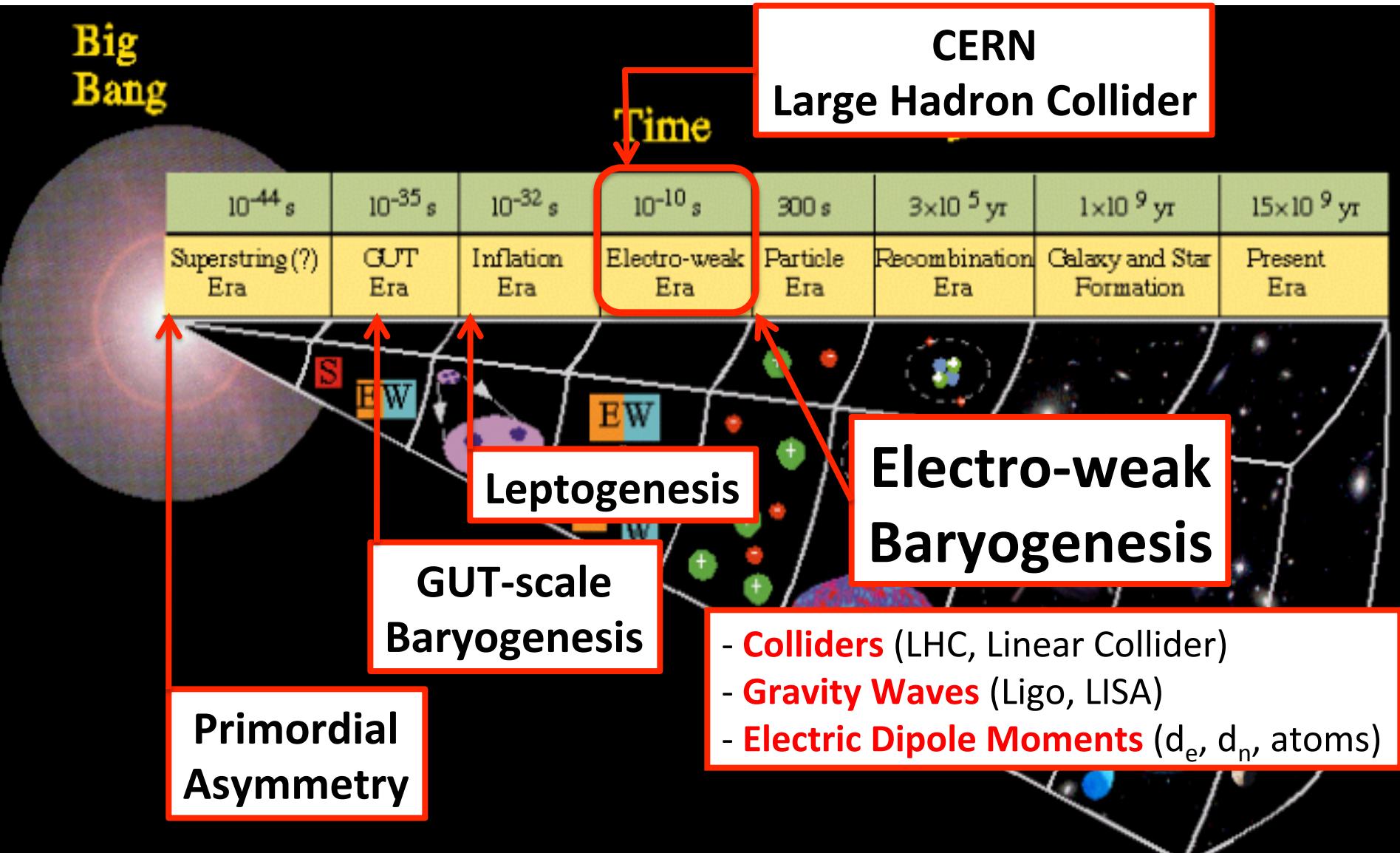


The Matter-Antimatter (Baryon) Asymmetry

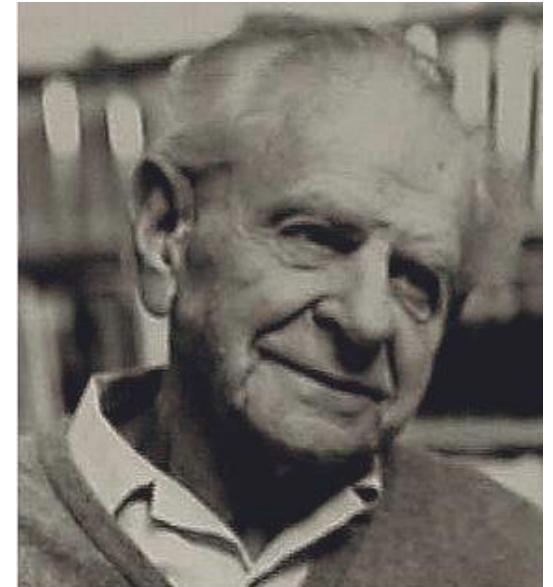


$$\frac{n_{\text{Baryons}} - n_{\text{Antibaryons}}}{n_\gamma} \approx 10^{-10}$$

No “Standard Model” of Baryogenesis!



“In so far as a scientific statement
speaks about **reality**,
it must be **falsifiable**:
And in so far as it is not falsifiable
it does not speak about reality”



Sir Karl Popper (1902-1994)

Karl Popper, “*Logik der Forschung*” (1934)
“*The Logic of Scientific Discovery*”

(Supersymmetric) Electro-Weak Baryogenesis:
a falsifiable theory

Ingredients of Baryogenesis

(1) **B**aryon Number violation

If B is conserved, the present BAU can only reflect asymmetric initial conditions

(2) **C** and **CP** violation

In the absence of a “preference” for matter or antimatter, B -nonconserving interactions will produce baryon and antibaryon excesses at the same rate: no net baryogenesis

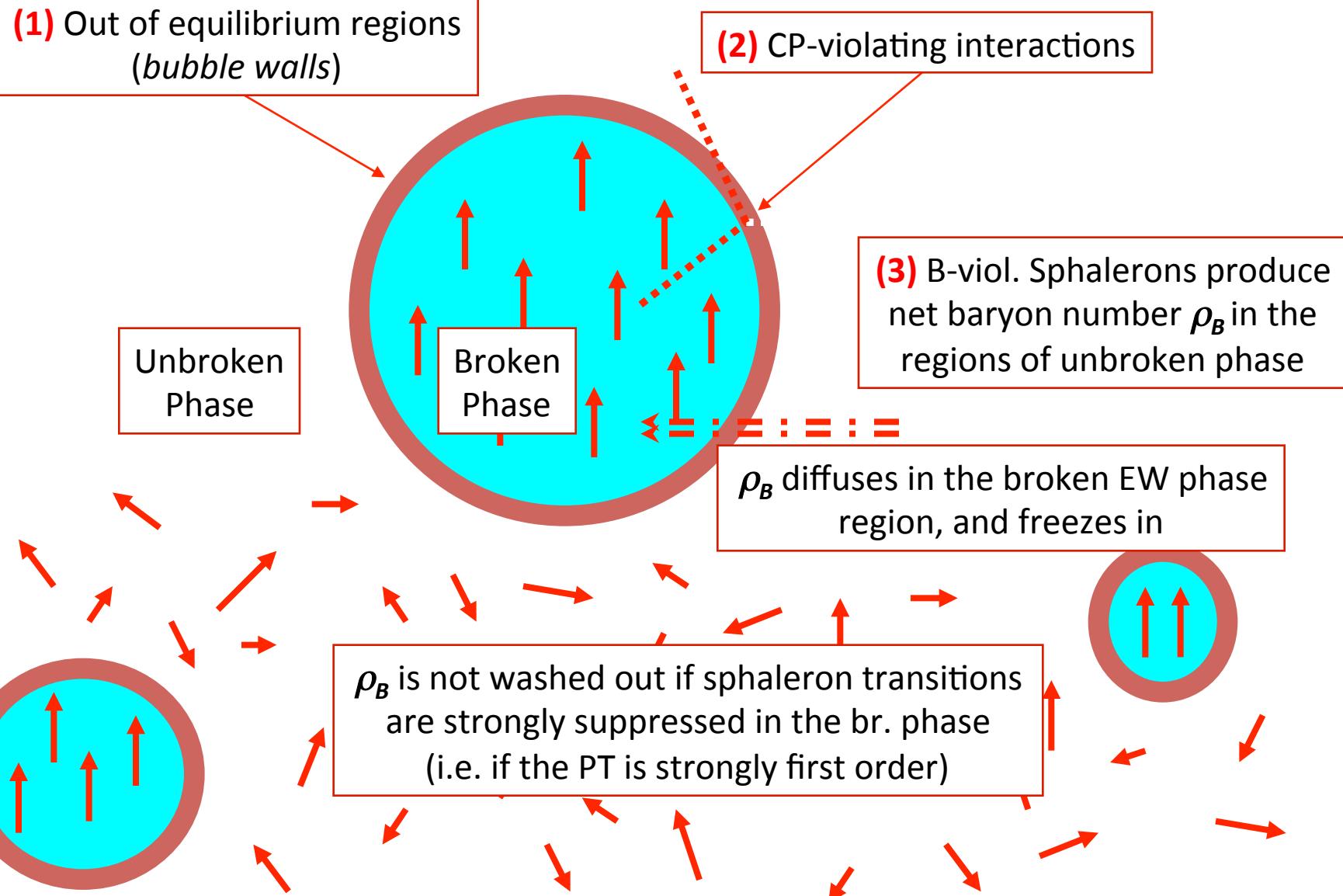
(3) **Out of Equilibrium** conditions

In chemical equilibrium the entropy is maximal when the chemical potential associated with all nonconserved quantum numbers vanishes

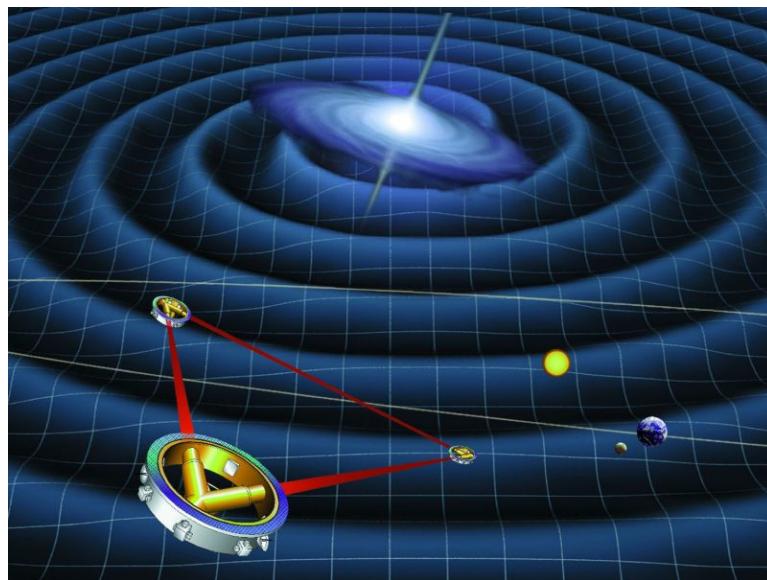
“**Sakharov conditions**”^(*)

^(*)A.D.Sakharov, JETP Letters **5**, 24 (1967)

Electro-Weak Baryogenesis

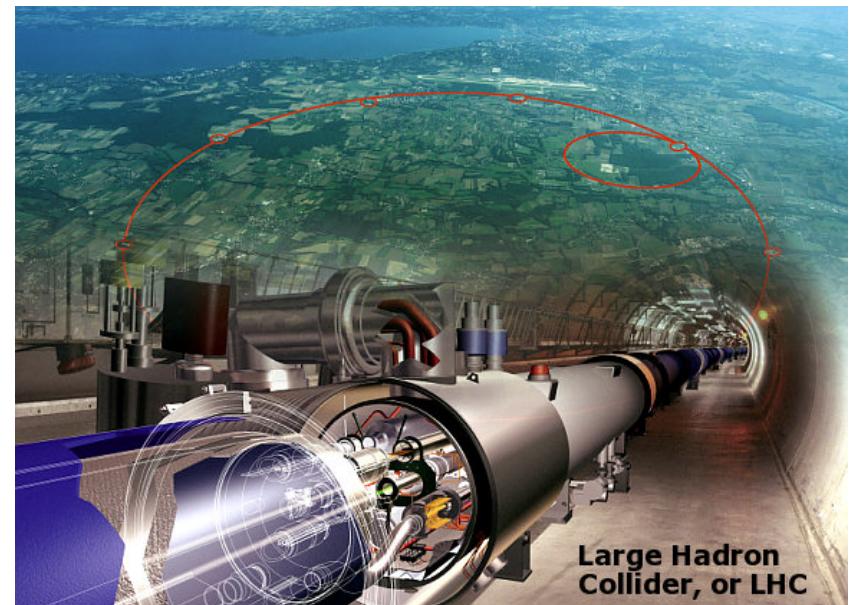
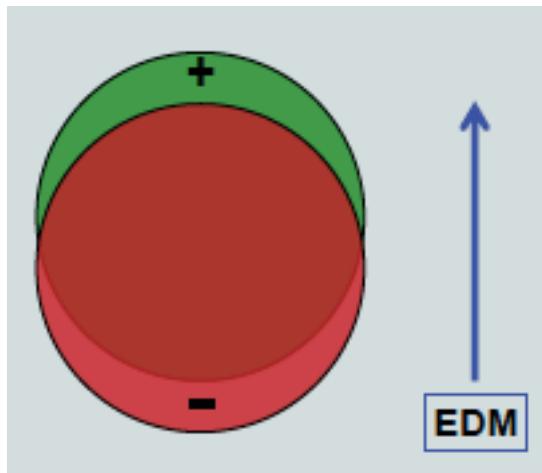


Experimental Tests of Electro-Weak Baryogenesis



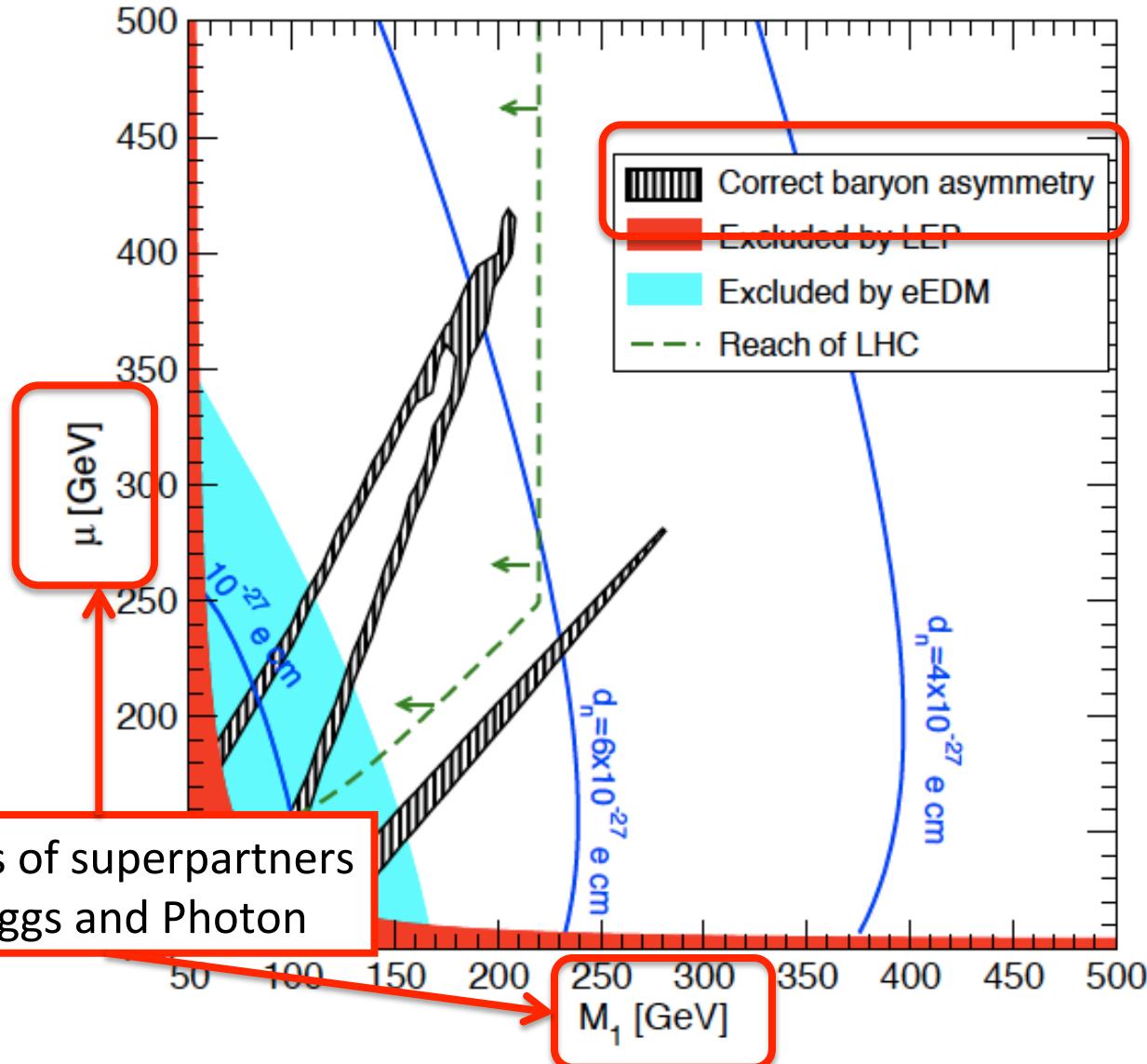
Gravity Waves from
Bubble Collisions

Large
Electric
Dipole
Moments
from CP
violation

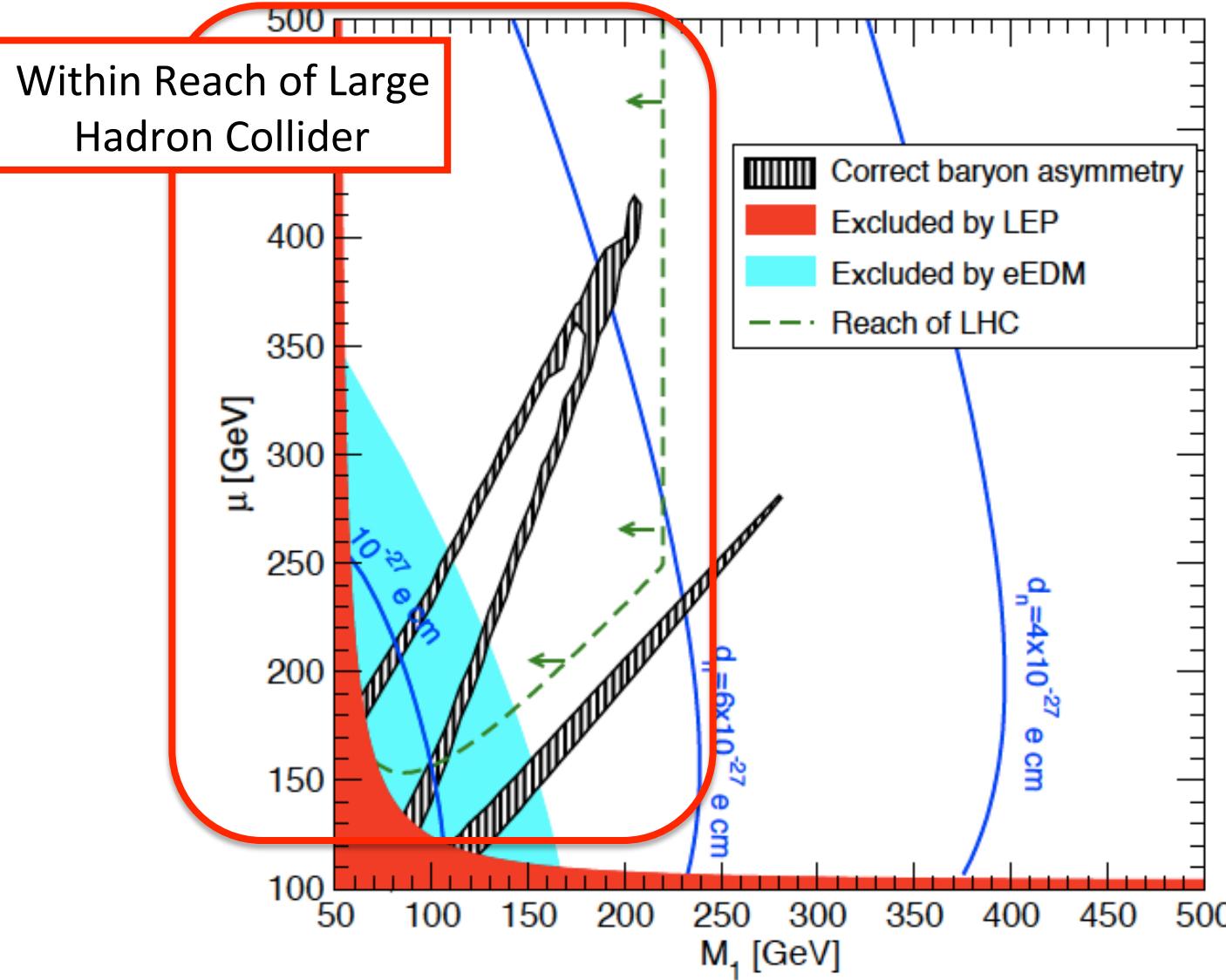


Light superpartners, distinctive
phenomenology at LHC

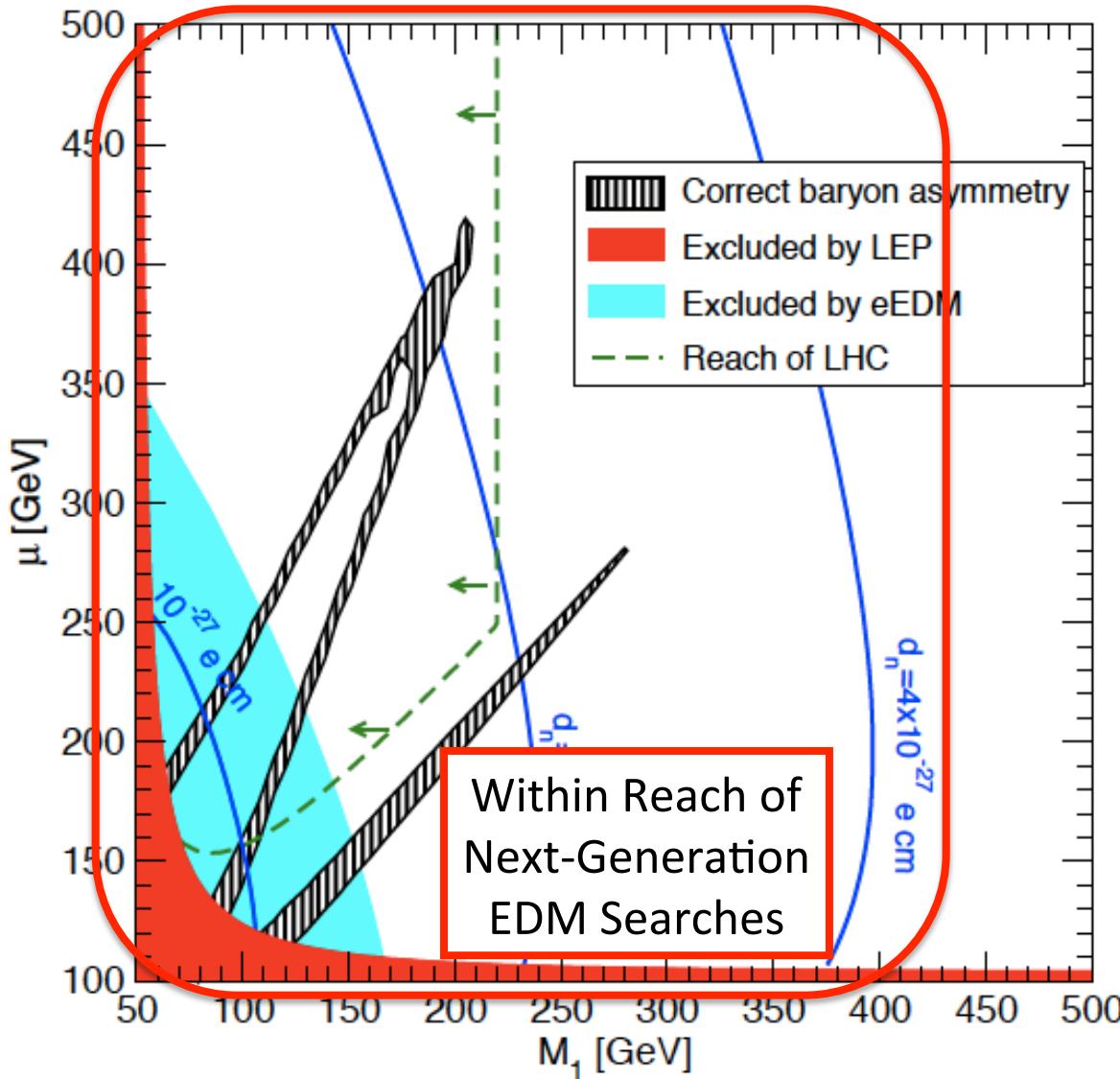
Electro-Weak Baryogenesis



Electro-Weak Baryogenesis

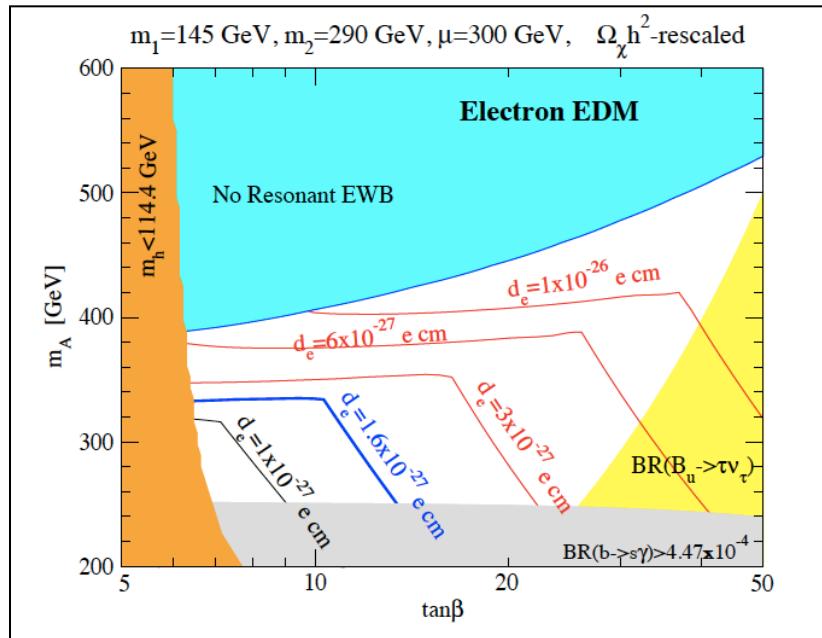


Electro-Weak Baryogenesis

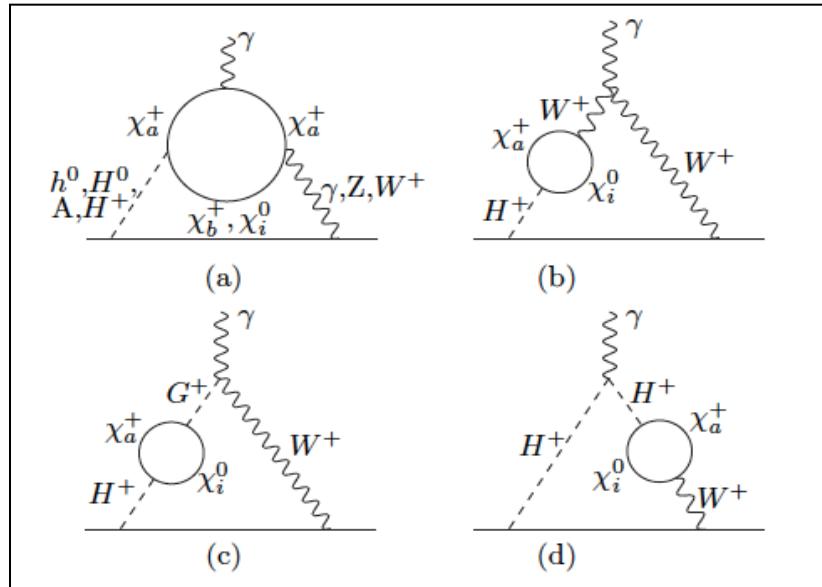


Electro-Weak Baryogenesis: Recent Progress

- Comprehensive Phenomenological Analysis of MSSM Electro-weak Baryogenesis (1)



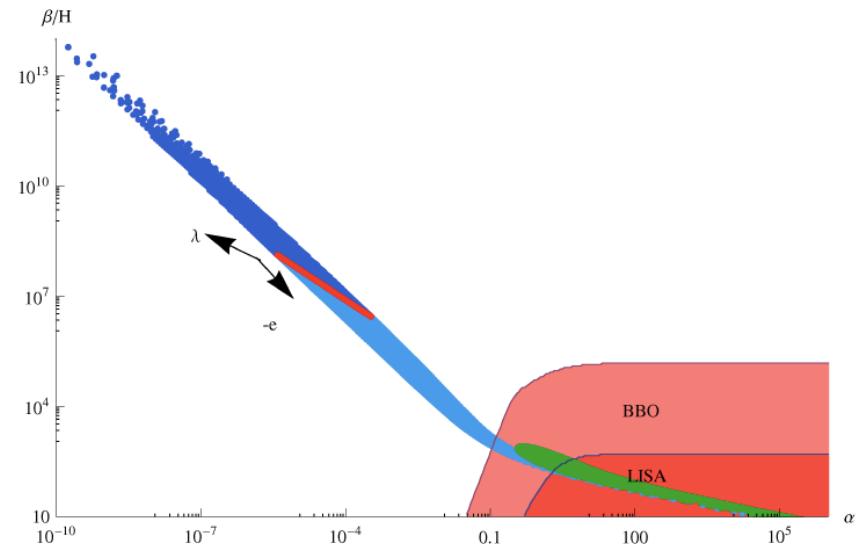
- Complete calculation of electro-weak 2-loop EDM amplitudes (2)
+ publicly available interface to numerical codes (2LEDM, 3)



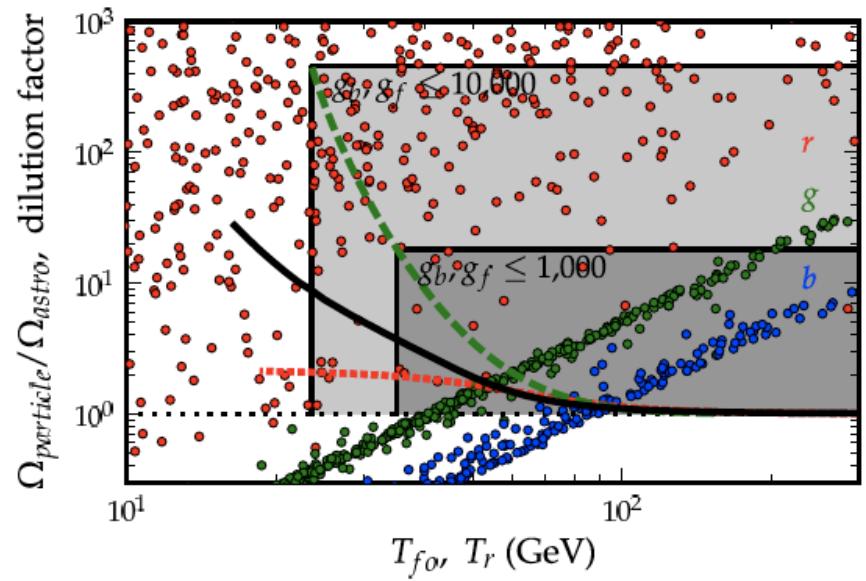
- (1) Cirigliano, Li, Profumo, Ramsey Musolf, JHEP 2010
(2) Li, Profumo, Ramsey-Musolf, PRD 2008, PLB 2009
(3) Li, Profumo, Ramsey-Musolf, JHEP 2010

Electro-Weak Baryogenesis: Recent Progress

- Gravity Waves as potential probes of EWB semi-analytic study (4)



- Impact of strongly first-order EW phase transition on thermal relics (5)

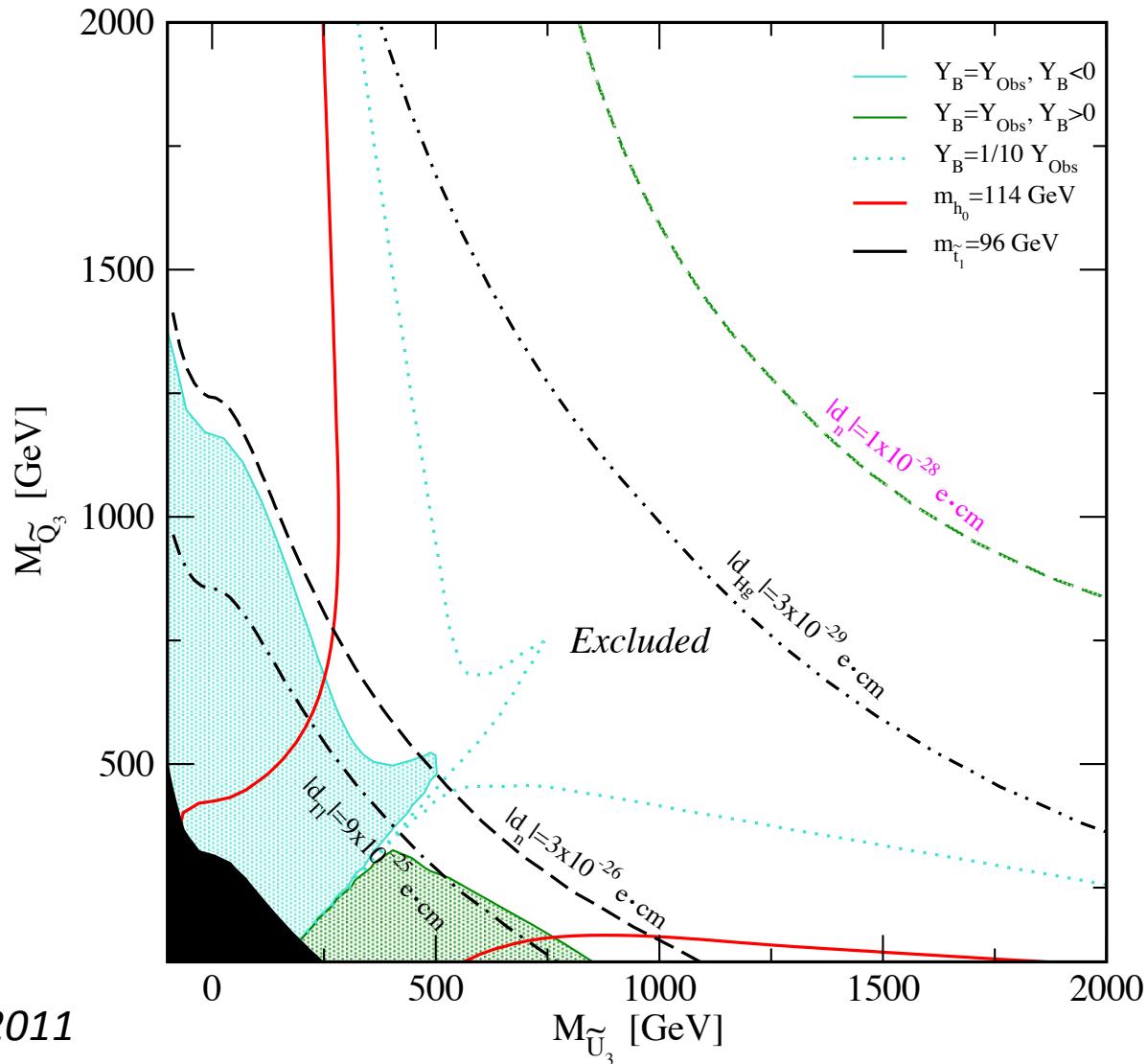


(4) Kehayias and Profumo, JCAP 2010

(5) Wainwright and Profumo, PRD 2009

Baryogenesis: Ongoing Work

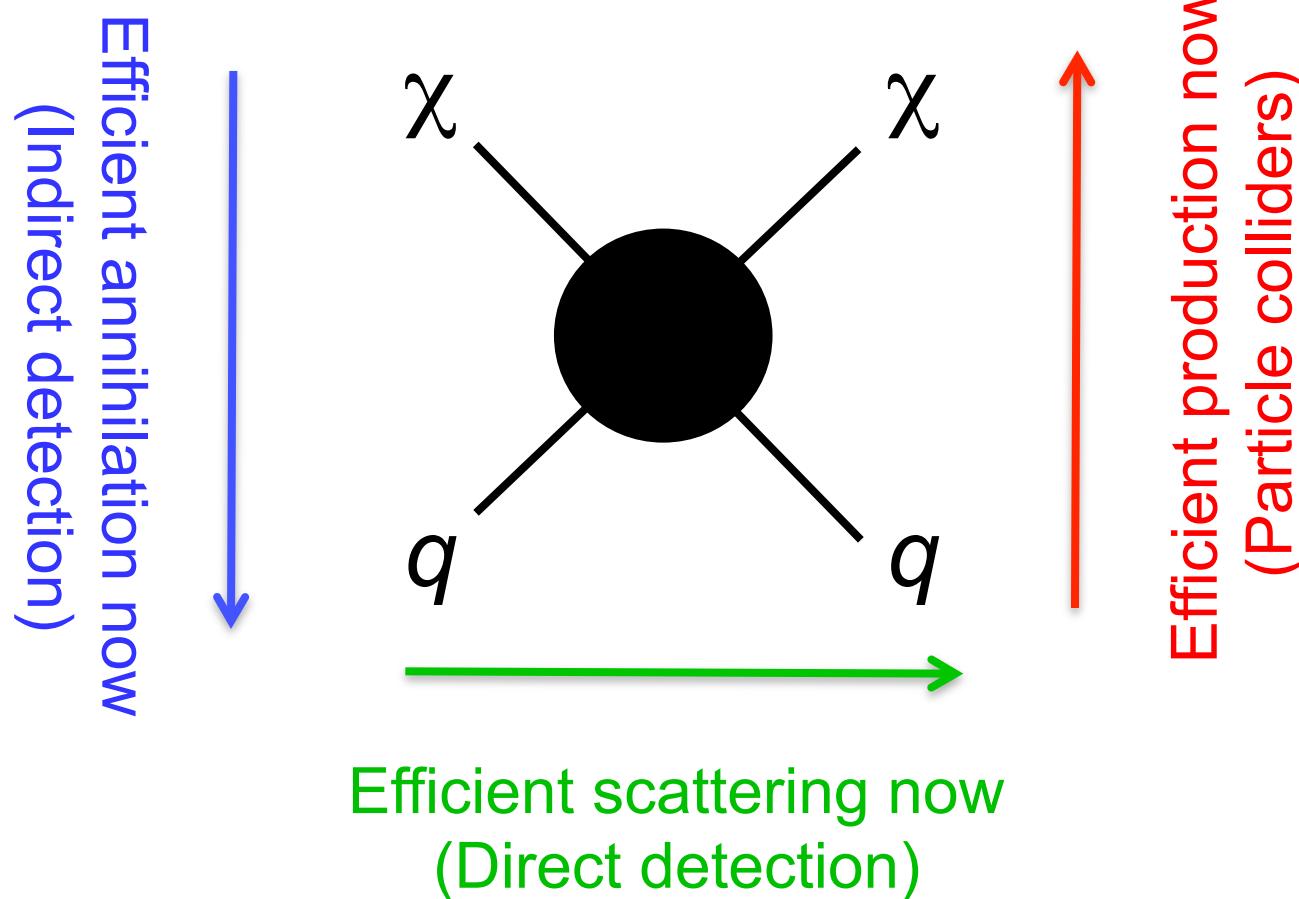
$\tan \beta = 10, A_t = 250 \text{ GeV}, \mu = 1000 \text{ GeV}$



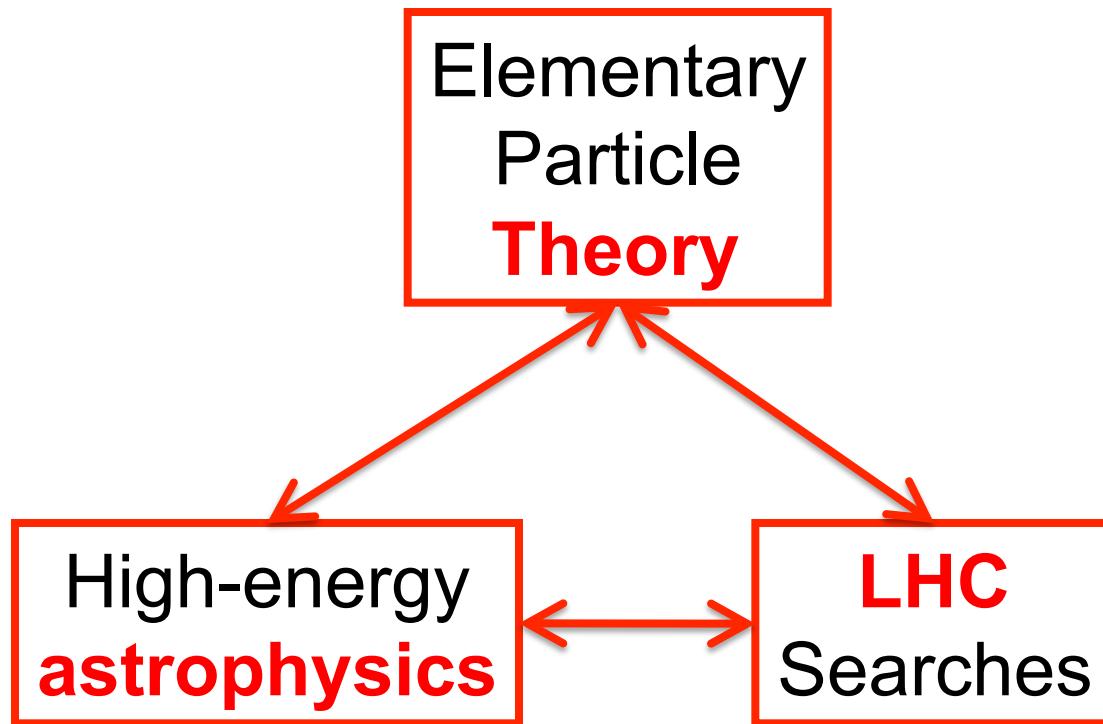
Baryogenesis: Ongoing Work

- New sources for electroweak baryogenesis (staus, sbottoms)
(with Wainwright and Kozaczuk)
- Analysis of issues of gauge invariance in finite-T effective potential
(with Wainwright and Ramsey-Musolf)
- Self-consistent models in MSSM and MSSM+singlet
(with Wainwright and Draper)

Particle Dark Matter: a multi-pronged approach in a Time of Discovery

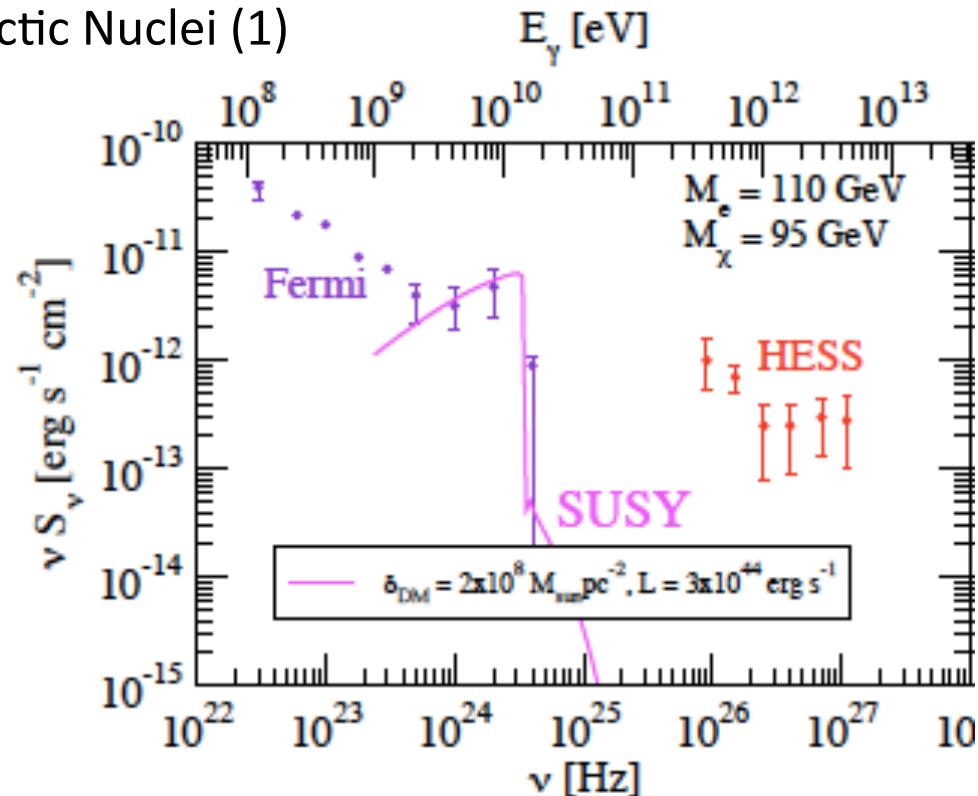


Particle Dark Matter: a comprehensive approach for a cross-disciplinary science



Particle Dark Matter: Theory

- Electron-DM scattering in Active Galactic Nuclei (1)



- Singlet dark matter models – 2-photon decay amplitude, role of meta-stable vacua (2)

- Cosmology of **Saxions** in gauge mediation (3)

- Direct detection in lepton-specific **2 Higgs Doublet models** (4)

- (1) *Ubaldi, Gorchtein and Profumo, PRD*
- (2) *Ubaldi, Wainwright and Profumo, PRD*
- (3) *Dine, Profumo and Ubaldi, in preparation*
- (4) *Boucenna and Profumo, PRD*

Particle Dark Matter: LHC Physics

- Searches for **2UED** with LHC;
ATLAS code development
(FeynRules -> Madgraph) (1)

- Vertex {WH11i, 1}, {WH11i, 2}, {Wmu11i, 3}, {W

$$-\frac{3}{4}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_3,\mu_4}-\frac{3}{4}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_3,\mu_4}$$

- Vertex {Wmu11i, 1}, {Wmu11i, 2}, {Wmu11i, 3}, {W

$$\begin{aligned}&-\frac{9}{2}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_1,\mu_4}\eta_{\mu_2,\mu_3}+\frac{9}{4}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_1,\mu_3}\eta_{\mu_2,\mu_4}\\&\frac{9}{4}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_1,\mu_3}\eta_{\mu_2,\mu_4}-\frac{9}{2}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_1,\mu_3}\eta_{\mu_2,\mu_4}\\&\frac{9}{4}ig_w^2\delta_{k_1,k_4}\delta_{k_2,k_3}\eta_{\mu_1,\mu_2}\eta_{\mu_3,\mu_4}+\frac{9}{4}ig_w^2\delta_{k_1,k_3}\delta_{k_2,k_4}\eta_{\mu_1,\mu_2}\eta_{\mu_3,\mu_4}\end{aligned}$$

- **Multi-top** (4x, 6x, 8x) LHC searches (2)

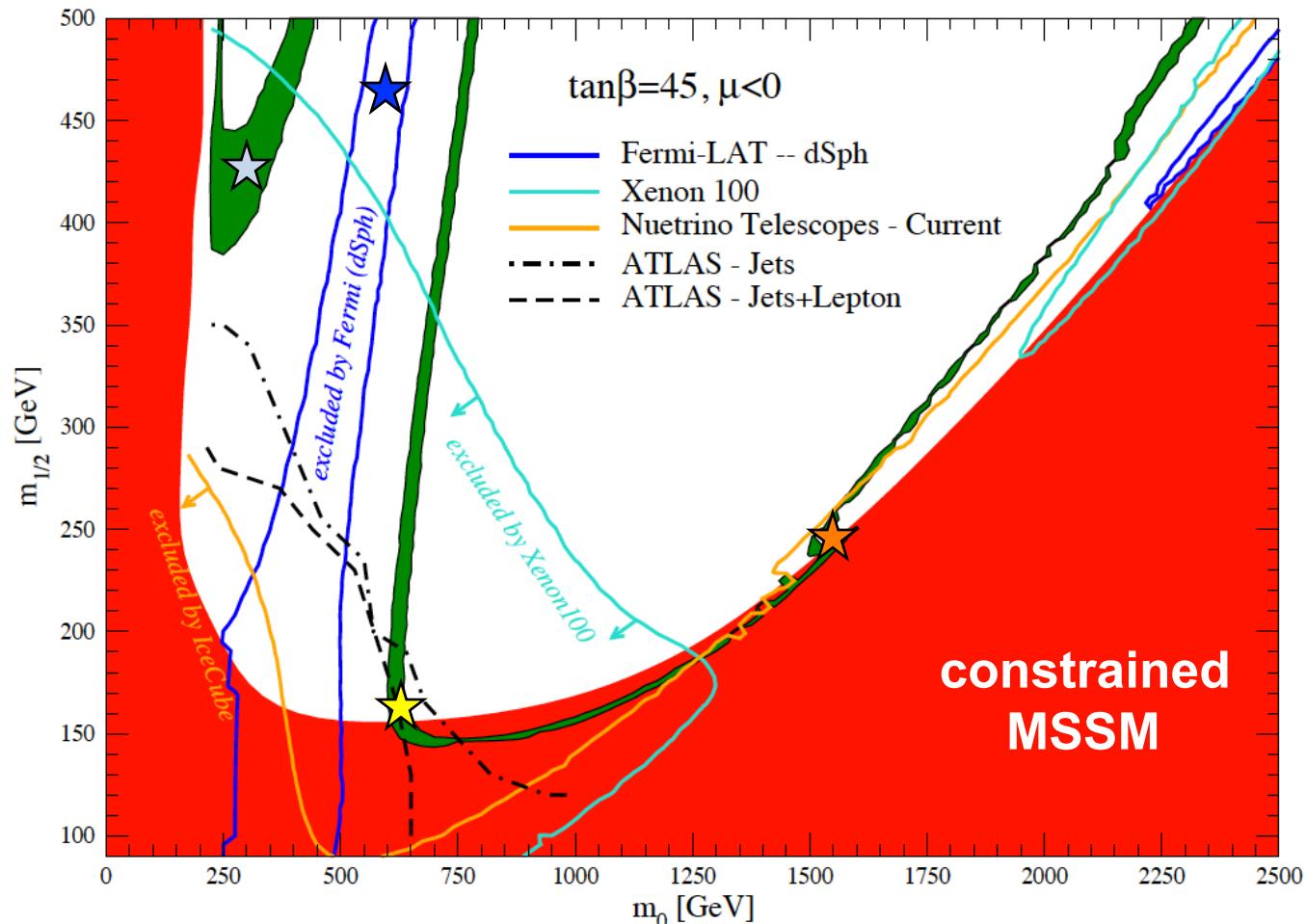
- Discriminating **axinos** and **gravitinos** at LHC (3)

(1) Manning and Profumo, in progress

(2) Manning, Profumo and Seiden, in progress

(3) Lee, Profumo and Ubaldi, in progress

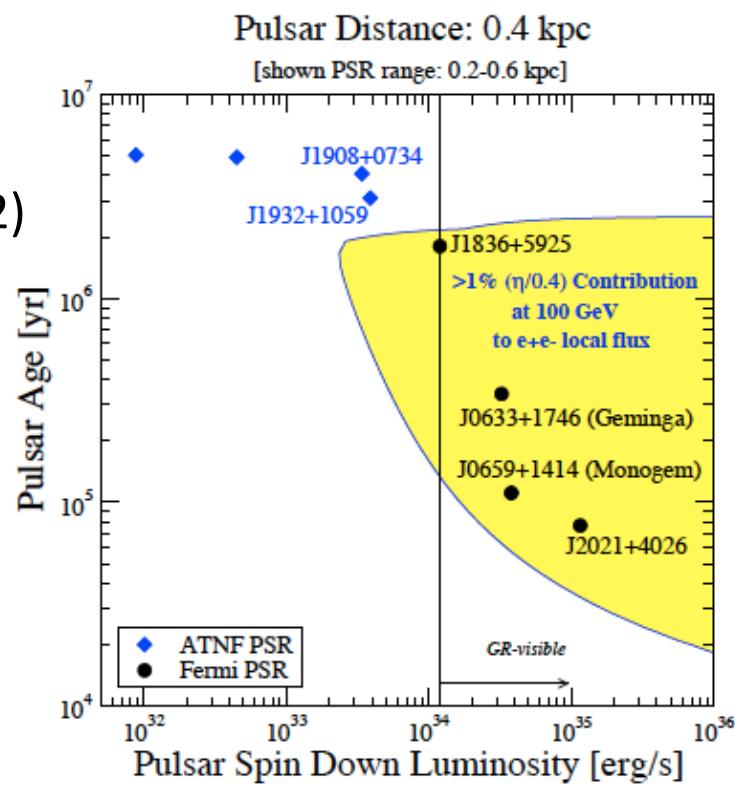
- ★ LHC is probing cosmologically interesting regions
- ☆ Only probed by Direct Detection
- ★ Only probed by Neutrino Telescopes
- ★ Only probed by Gamma-Ray Telescopes



Particle Dark Matter: Astrophysical Backgrounds

The field does not seem to need theory model building as much as **reliable assessment of signals**, and hence of the relevant **astrophysical backgrounds**

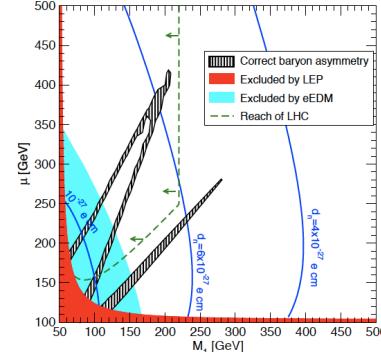
- **Pulsars** as sources of high-energy positrons – connection to Fermi gamma-ray pulsars (1)
- **Milli-second pulsars** and diffuse gamma rays (2)
- Systematic effects in signals from diffuse emissions (“**Fermi Haze**”)(3)
- Hadronic models for **cluster radio halos** (4) – relevant for **Cosmology !!!**



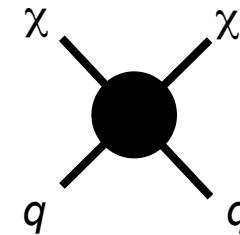
- (1) Gendelov, Profumo and Dormody, JCAP 2010
- (2) Siegal-Gaskins, Profumo et al, MNRAS 2011
- (3) Linden and Profumo, APJL 2011
- (4) Jeltema and Profumo, Astroph. J. 2010

Summary

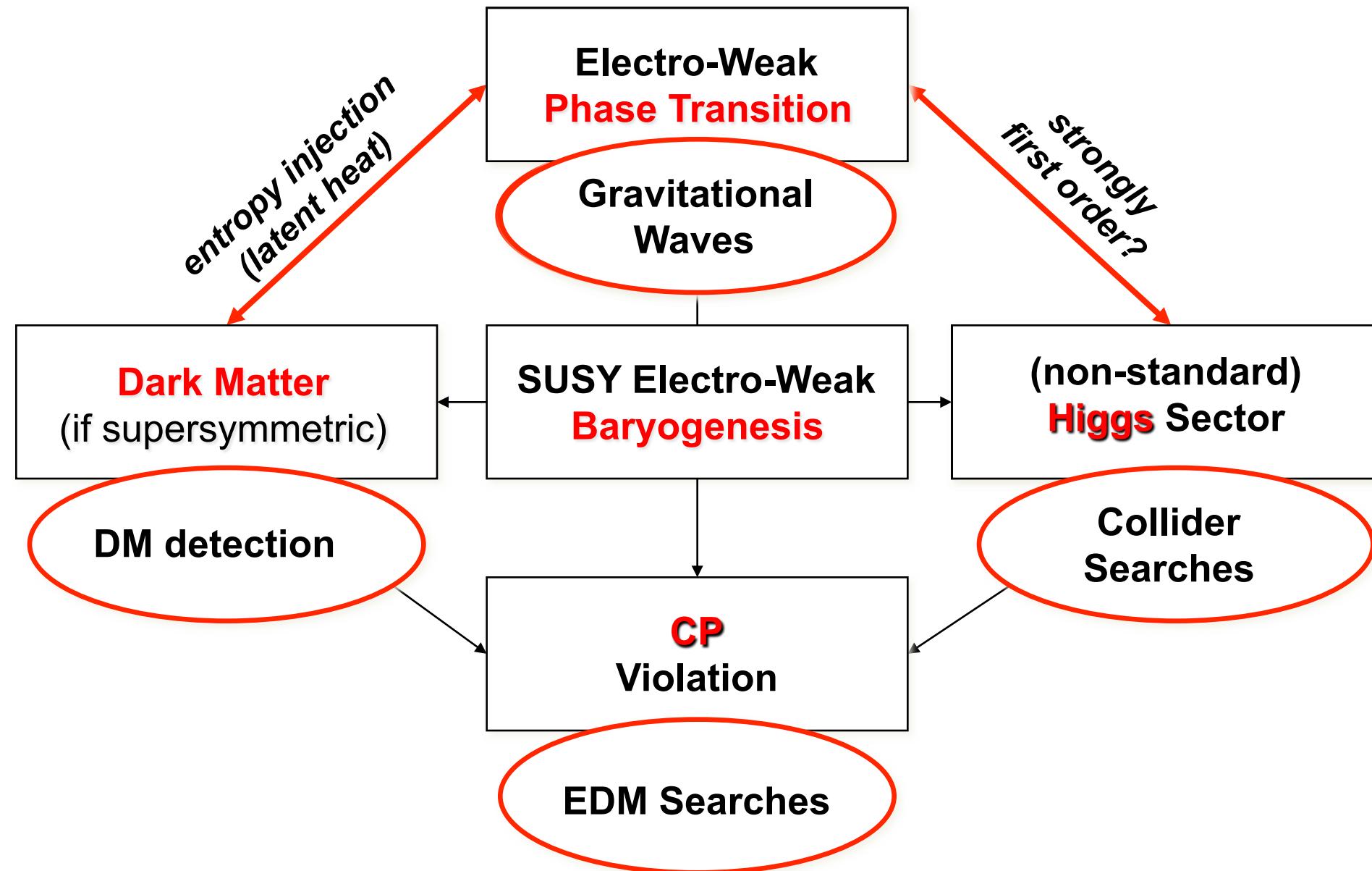
Testable theories for
the origin of matter



A cross-disciplinary approach
in the hunt for dark matter



Electro-Weak Baryogenesis: Probes



Ingredients of Baryogenesis

- **B**aryon Number violation

If B is conserved, the present BAU can only reflect asymmetric initial conditions

- **C** and **CP** violation

In the absence of a “preference” for matter or antimatter, B -nonconserving interactions will produce baryon and antibaryon excesses at the same rate: no net baryogenesis

- **Out of Equilibrium** conditions

In chemical equilibrium the entropy is maximal when the chemical potential associated with all nonconserved quantum numbers vanishes

“**Sakharov conditions**”^(*)

Electro-Weak Baryogenesis

The Electro-Weak Phase Transition fulfills all
3 Sakharov requirements^(*) (**Electro-weak Baryogenesis**)

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Electro-Weak Baryogenesis

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- ✓ **B** violation: Weak **Sphaleron** Transitions

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Electro-Weak Baryogenesis

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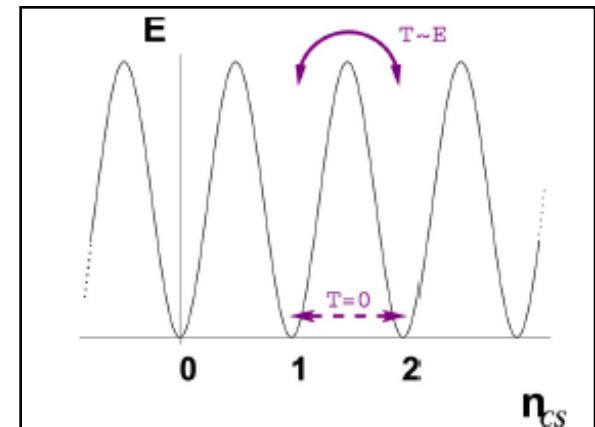
→ ✓ **B** violation: Weak **Sphaleron** Transitions

- Classically, baryonic and leptonic currents are **conserved** in the EW theory
- Quantum corrections produce **anomalous transitions** between non-degenerate SU(2) field configurations vacua that **violate B+L** (but preserve B-L)
- **B-violation rate is unsuppressed at $T > T_c$, and is exponentially suppressed at $T < T_c$**

$$\Gamma_{sph} \propto \alpha_W T^4$$

$$\Gamma_{sph} \propto \exp[-E_{sph}(T)/T]$$

$$E_{sph}(T) \propto \langle \phi \rangle(T)$$



Different vacua: $\Delta(B+L) = \Delta n_{CS}$

Electro-Weak Baryogenesis

The Electro-Weak Phase Transition fulfills all
3 Sakharov requirements^(*) (**Electro-weak Baryogenesis**)

- ✓ **B** violation: Weak **Sphaleron** Transitions
- ✓ **CP** violation: **CKM** (or new *CP-phases*)

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Electro-Weak Baryogenesis

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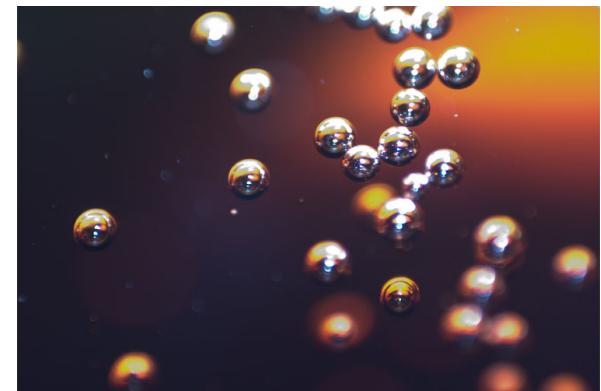
- ✓ **B** violation: Weak **Sphaleron** Transitions
- ✓ **CP** violation: **CKM** (or new *CP-phases*)
- ✓ **Out of Equilibrium**: **Bubble Walls** of broken EW phase

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Electro-Weak Baryogenesis

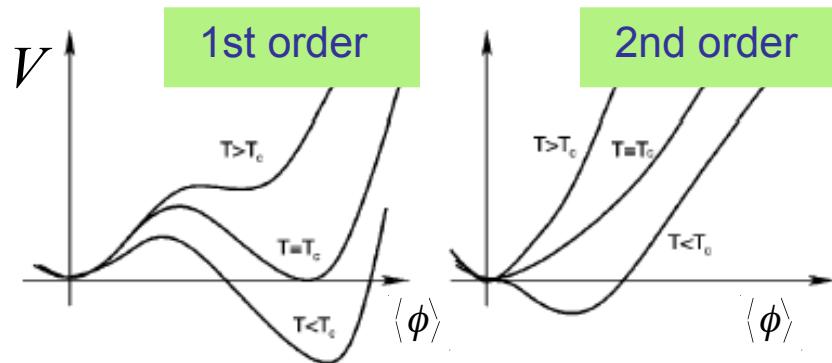
The Electro-Weak Phase Transition fulfills all
3 Sakharov requirements^(*) (**Electro-weak Baryogenesis**)

→ ✓ **B** violation: Weak **Sphaleron** Transitions



→ ✓ **CP** violation: **CKM** (or new **CP-phases**)

→ ✓ **Out of Equilibrium**: **Bubble Walls** of broken EW phase



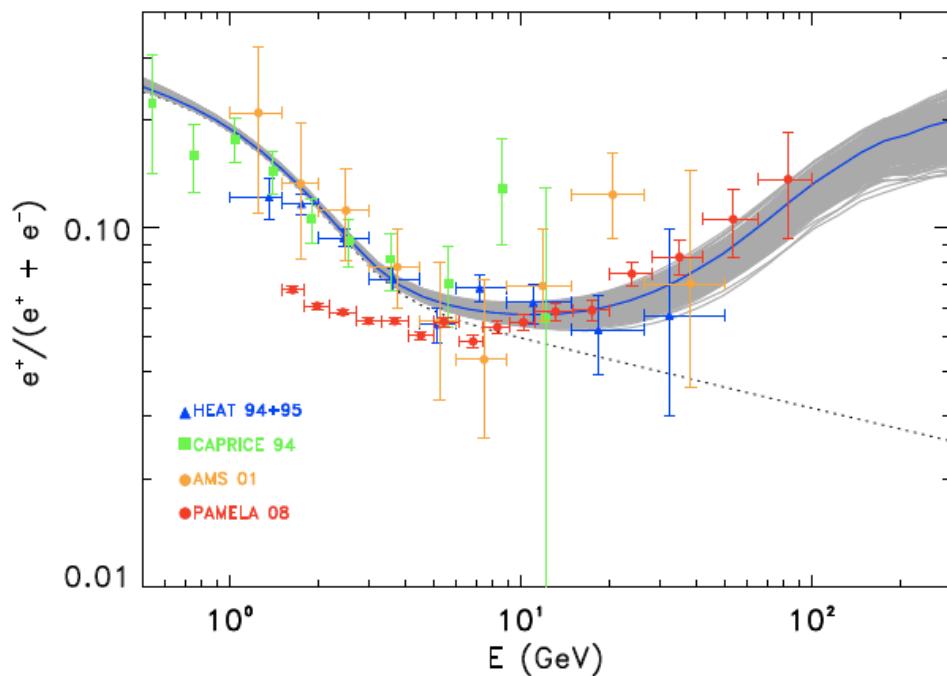
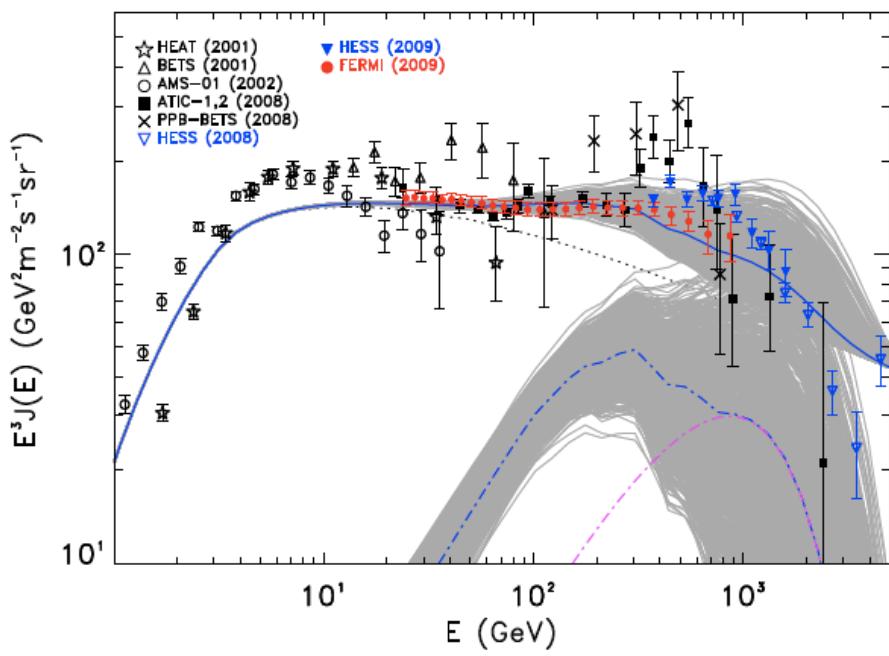
- If the EWPT is first order (cubic term!), it proceeds through Bubble nucleation
- The (expanding) Bubble Walls are out of thermal equilibrium

^(*)V.A.Kuzmin, V.A.Rubakov and M.E.Shaposhnikov, Phys.Lett. **B197**, 49 (1989)

Particle Dark Matter: Indirect Detection and Theory

■ Cosmic Rays

- Interpretation of Fermi electron/positron data (1)
- Fermi pulsars vs excess positrons (2)
- Dark Matter vs excess positrons (3,4,5)



(1) Grasso, Profumo, Strong et al, Astropart.Phys. 2009

(2) Gendevlev, Profumo and Dormody, JCAP 2010

(3) Jeltema and Profumo, JCAP 2009

(4) Brun, Delahayie, Diemand, Profumo, PRD 2009

(5) Cyr-Racine, Profumo and Sigurdson, PRD 2009

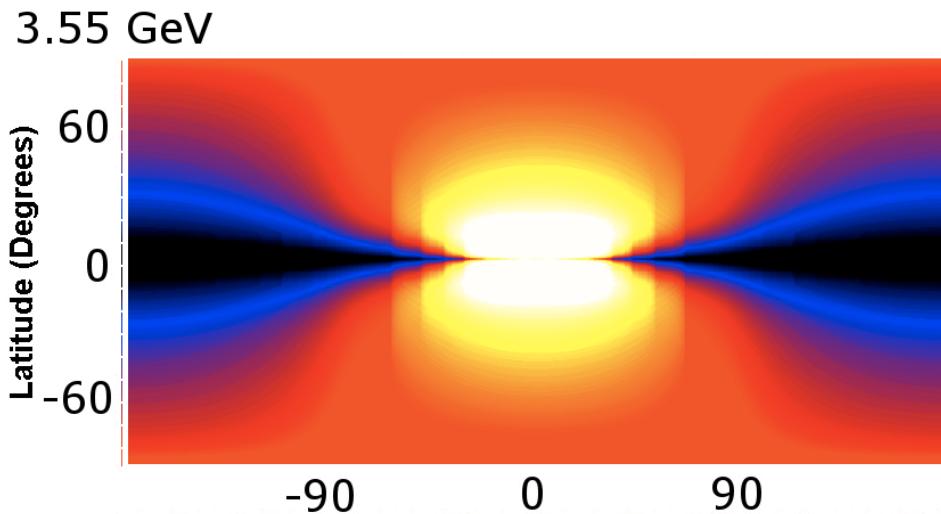
Particle Dark Matter: Indirect Detection and Theory

■ Diffuse Emissions

- WMAP Haze with consistent CR models (1)
- A Fermi haze? Systematic effects (2)
- Extragalactic Inverse Compton (3)

■ Theory/Multi-disciplinary

- Multi-component Dark Matter models:
direct, indirect and collider searches (4)



(1) Linden, Profumo and Anderson, PRD 2010 sub.

(2) Linden and Profumo, Astroph. J. Lett. 2010

(3) Jeltema and Profumo, JCAP 2009

(4) Profumo, Sigurdson and Ubaldi, PRD 2009

