FLATIRON INSTITUTE

Center for Computational Astrophysics

Colloquium 10 September 2021

New Challenges in Cosmology, Galaxy Formation, and Planet Habitability

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11 December 2020

New Challenges in Cosmology

The Hubble tension between early universe and local measurements of H_0 can be resolved by a brief episode of dark energy at redshift z ~ 3500. New N-body simulations have shown that this Early Dark Energy scenario predicts earlier structure formation, e.g. ~ 50% more clusters than ΛCDM at redshift z ~ 1.

Galaxy Formation

Galaxies were long thought to start as disks, but HST images show that most galaxies instead start prolate (pickle shaped). Galaxy simulations can explain this as a consequence of the filamentary nature of the ΛCDM dark matter distribution. But comparisons between simulations and observations using novel machine learning methods reveal other potential challenges, including massive star-forming clumps seen in many high-redshift galaxies.

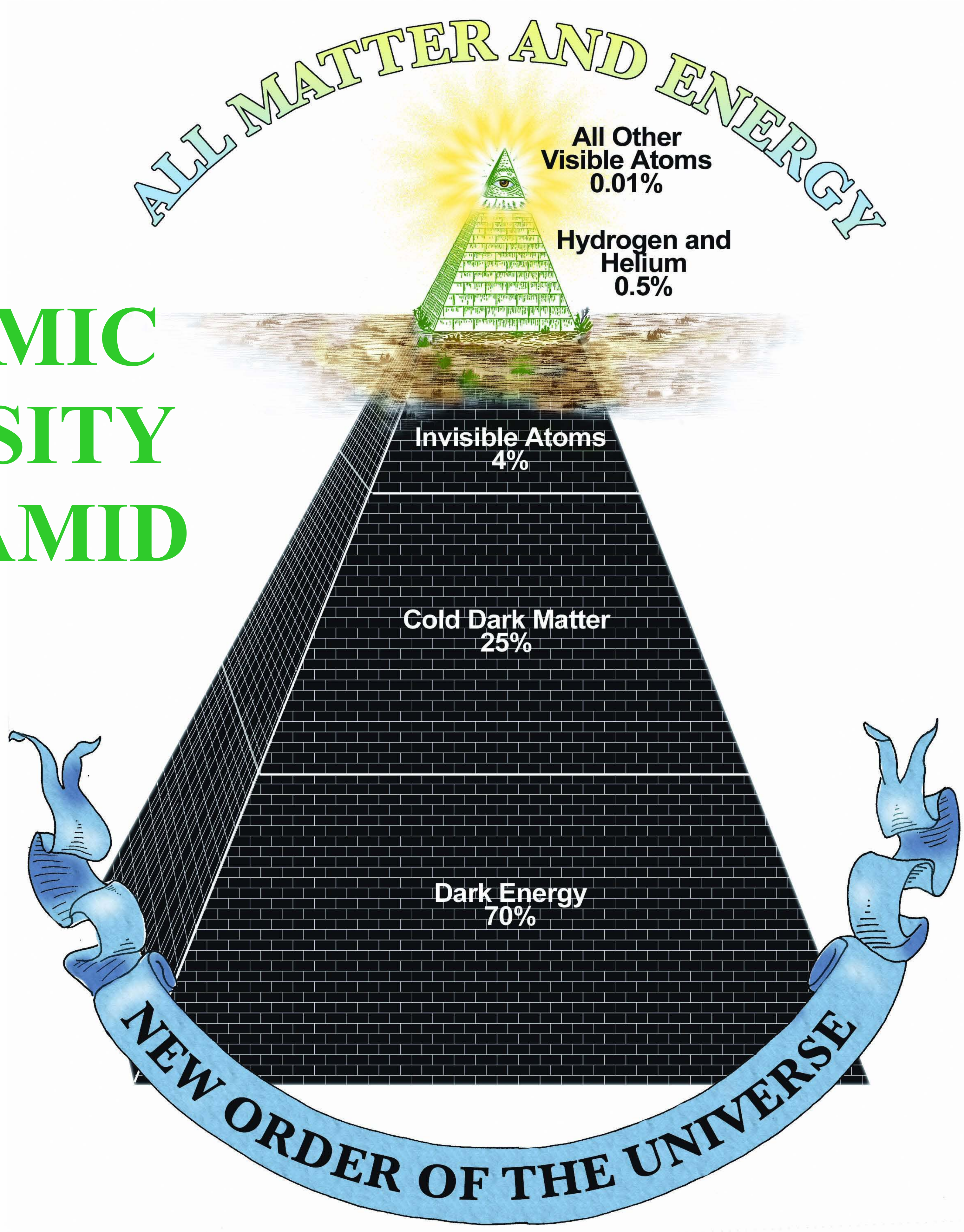
Planet Habitability

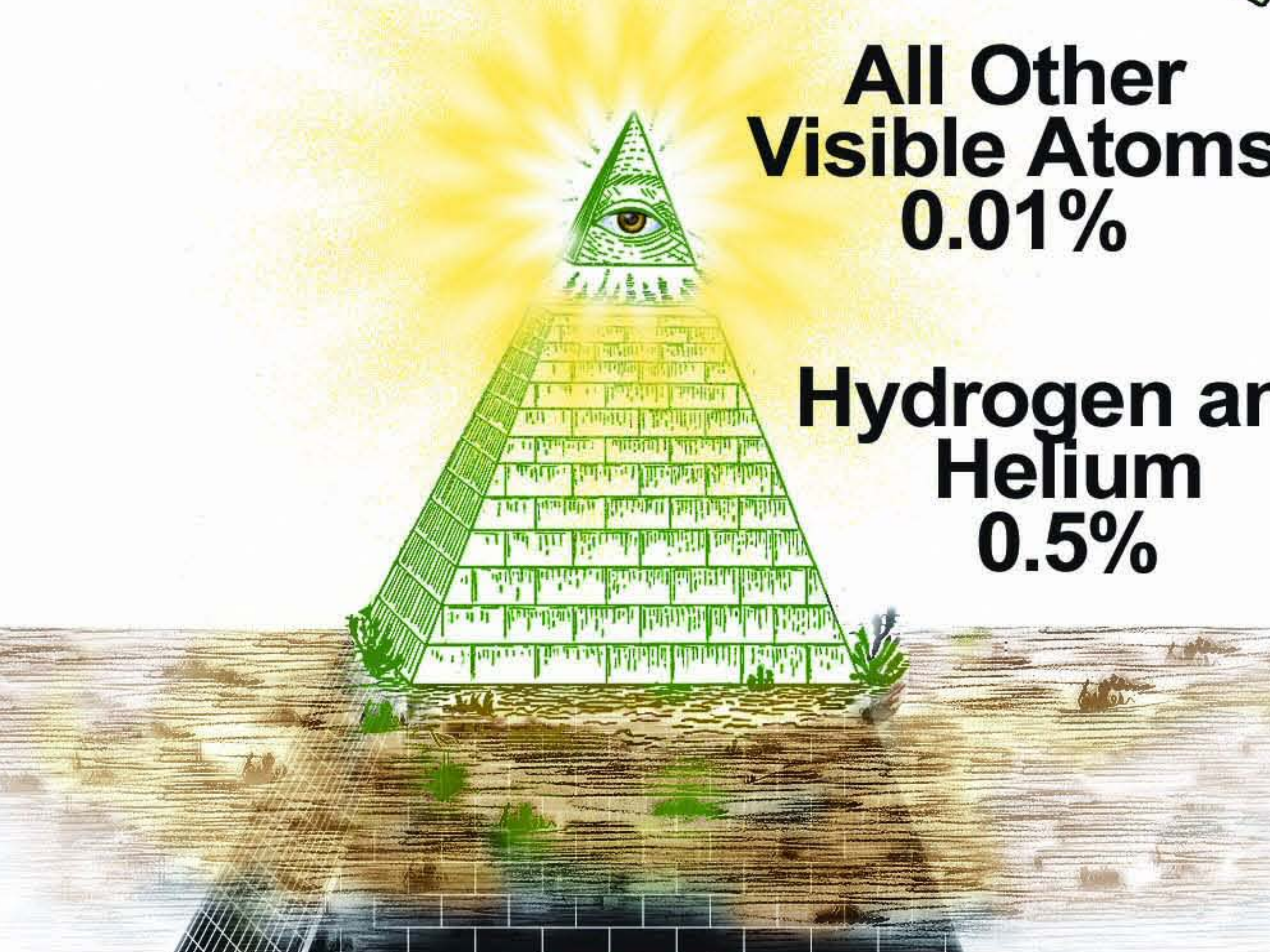
Earth may be a radioactively Goldilocks planet, with just the right amount of radiogenic heating by Th and U for a magnetic field and plate tectonics, both of which may be necessary for the evolution of complex life.

This picture is beautiful but misleading, since it only shows about 0.5% of the cosmic density.

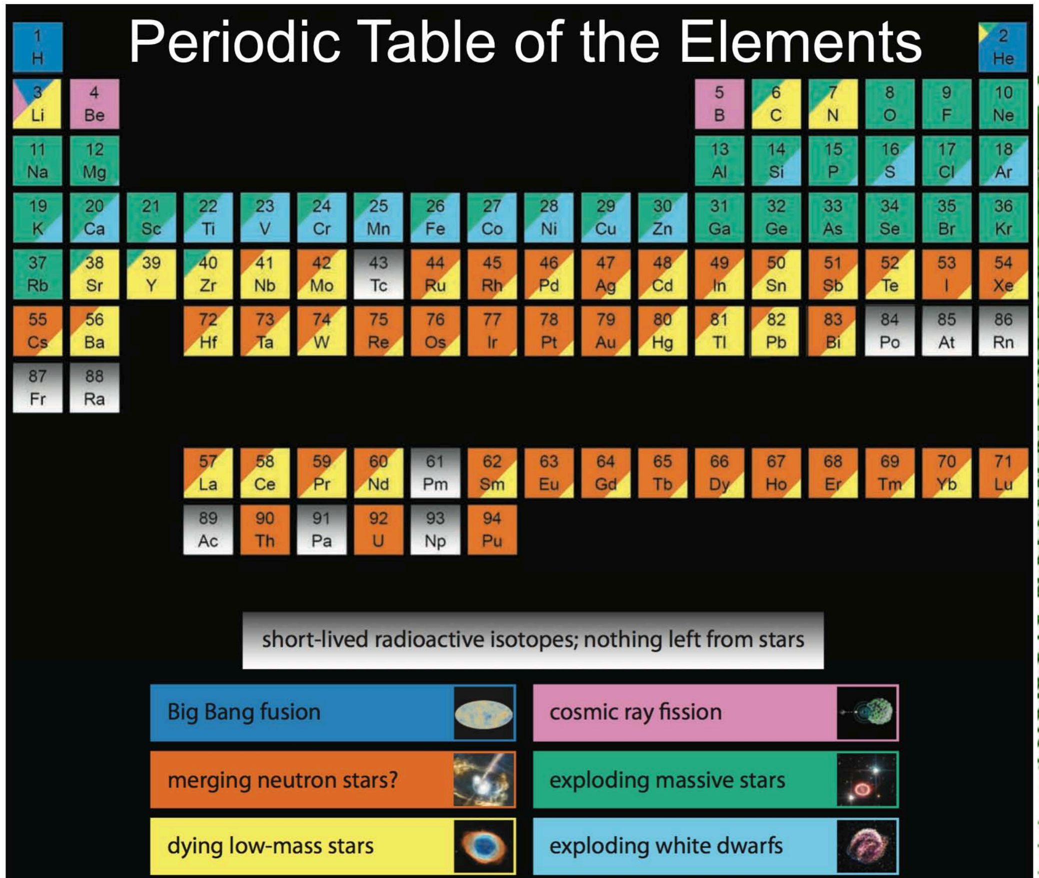
The other 99.5% of the universe is dark.

COSMIC DENSITY PYRAMID



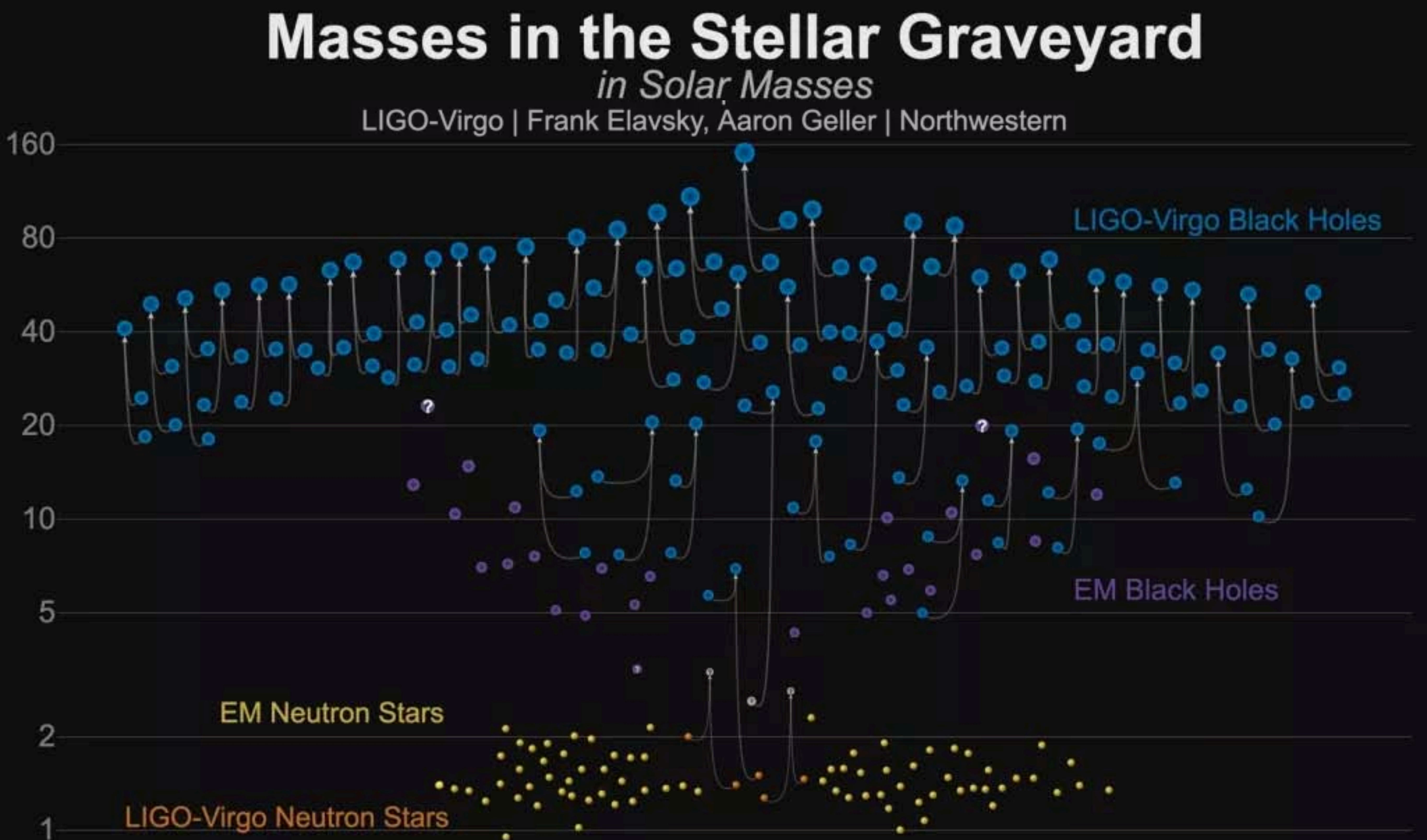


All Other Visible Atoms
0.01%
Hydrogen ar
Helium
0.5%



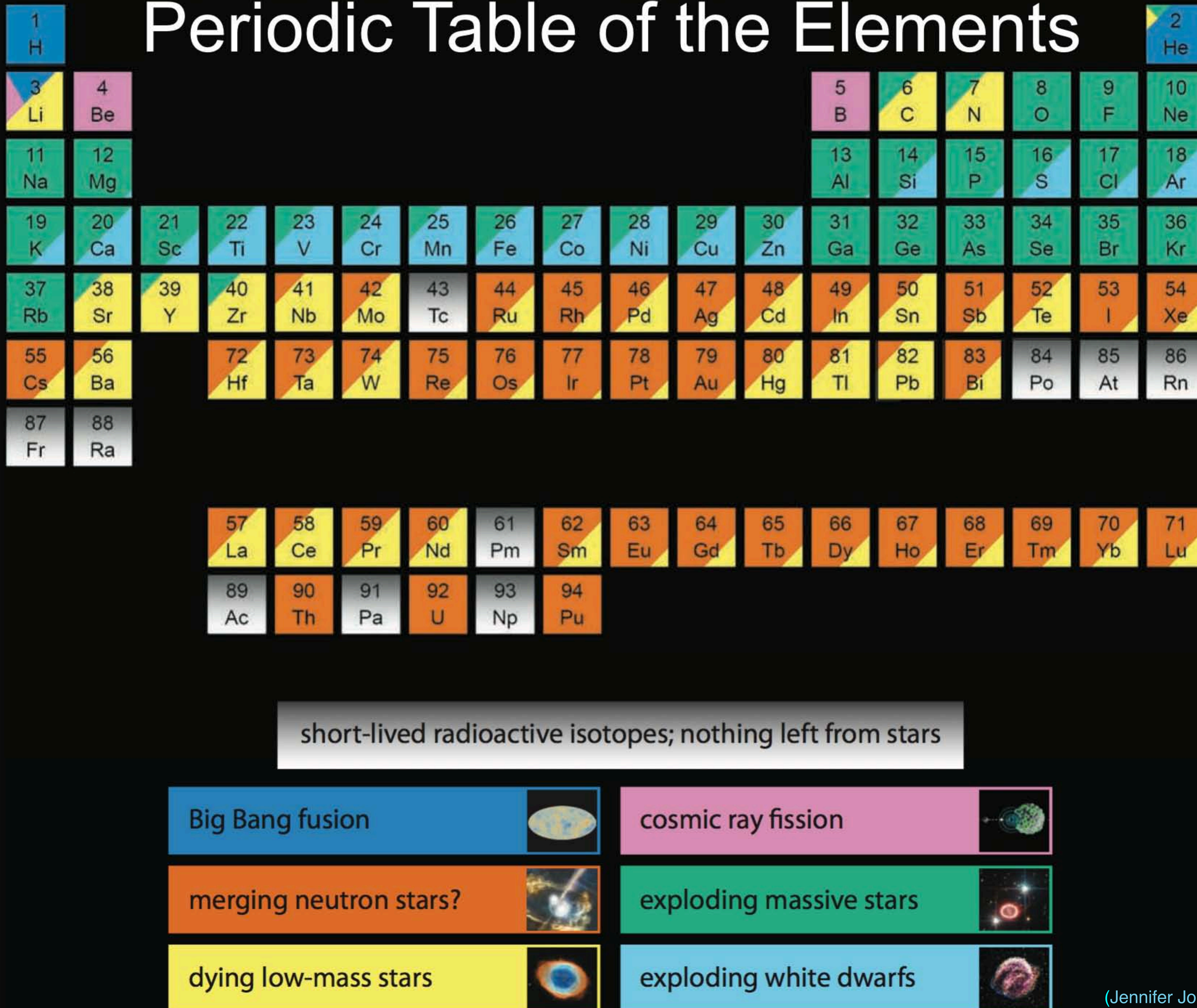
stardust

first
stars



In August 2017 LIGO and VIRGO announced the discovery of gravity waves from a neutron star merger. Such events probably generate most of the r-process elements like europium, gold, thorium, and uranium.

Periodic Table of the Elements



(Jennifer Johnson)

Dark Matter Ships

on a

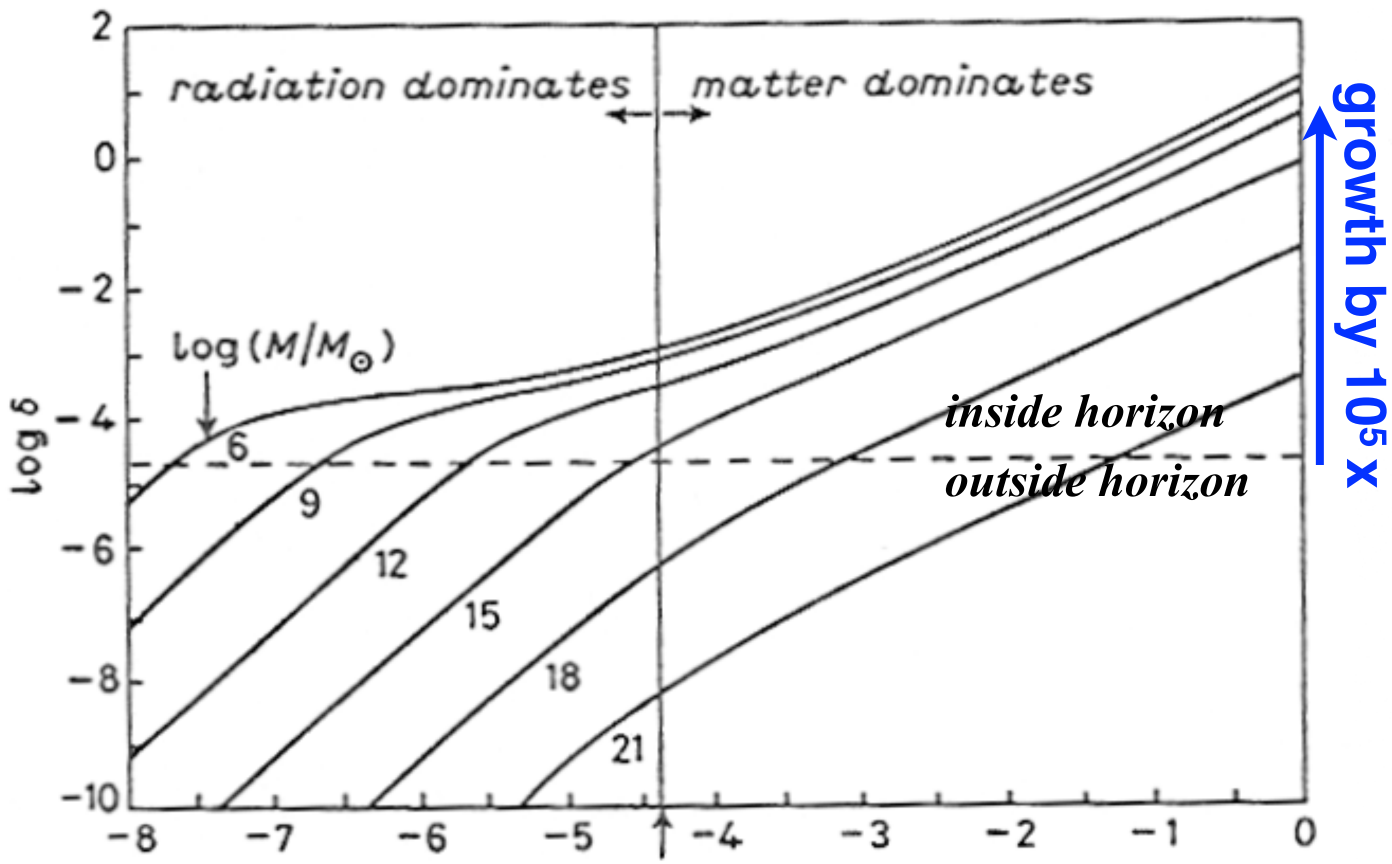
Dark Energy Ocean

Matter and Energy Content of the Universe

ΛCDM

Double Dark Theory

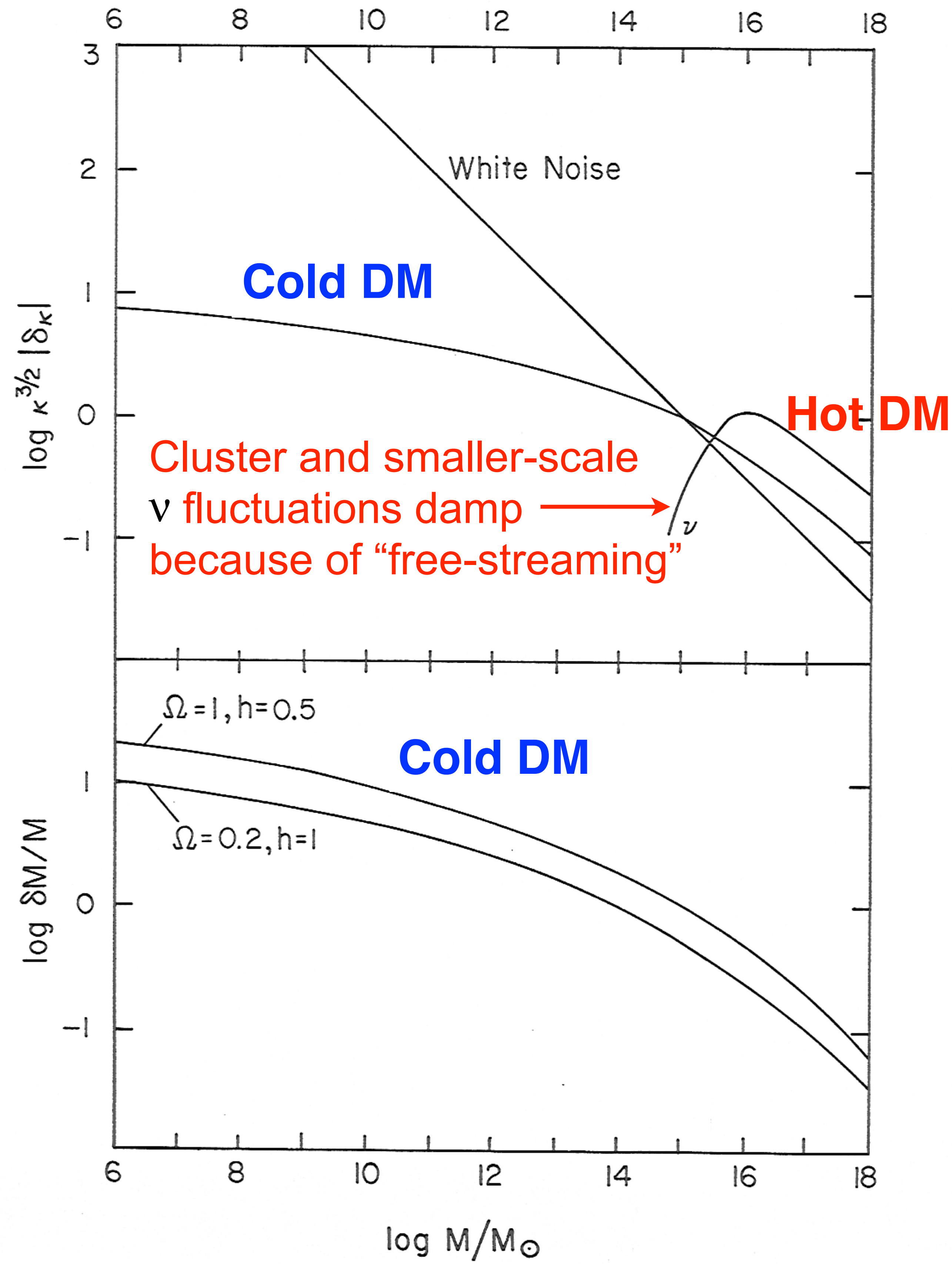
CDM Structure Formation: Linear Theory



Scale factor $a = 1/(1+z)$

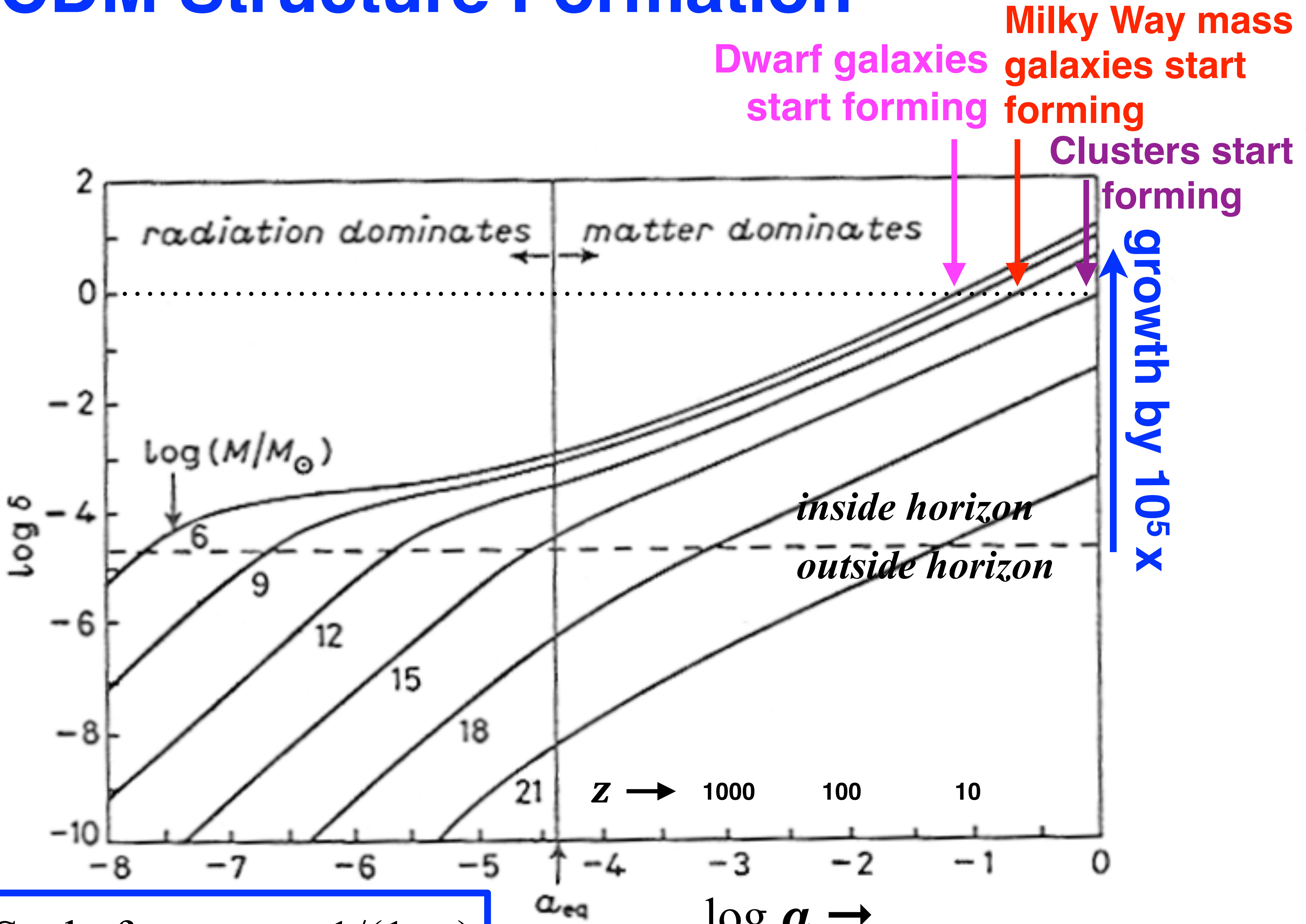
CDM fluctuations that enter the horizon during the radiation dominated era, with masses less than about 10^{15} M⊙, grow only ∝ log a, because they are not in the gravitationally dominant component. But matter fluctuations that enter the horizon in the matter-dominated era grow ∝ a. This explains the characteristic shape of the CDM fluctuation spectrum, with $\delta(k) \propto k^{-n/2-2} \log k$ and $\delta_{\text{primordial}} = k^n$

Primack & Blumenthal 1983,
Primack Varenna Lectures 1984



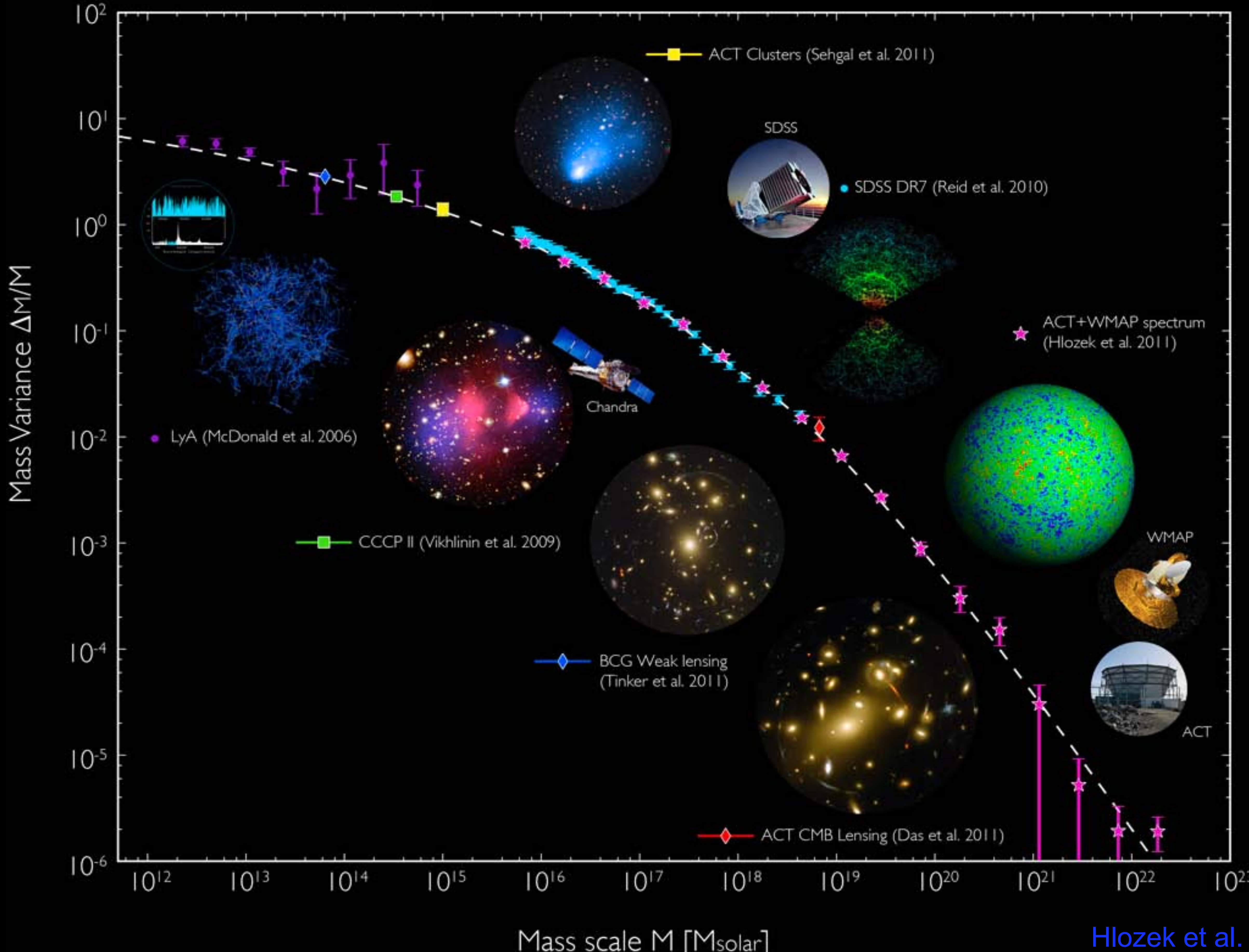
Blumenthal, Faber, Primack, & Rees 1984

CDM Structure Formation



Scale factor $a = 1/(1+z)$

Matter Distribution Agrees with Double Dark Theory!

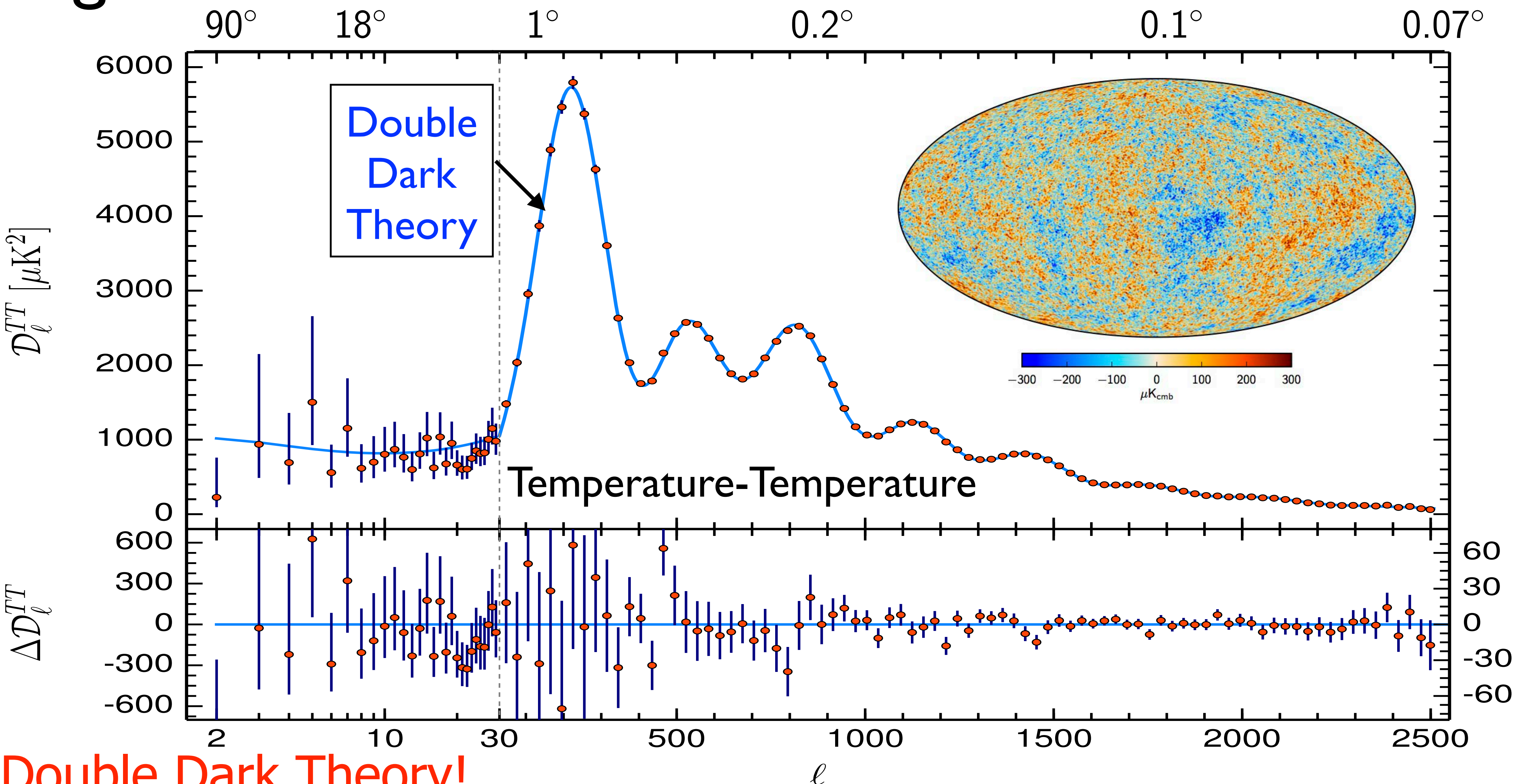


Hlozek et al. 2012

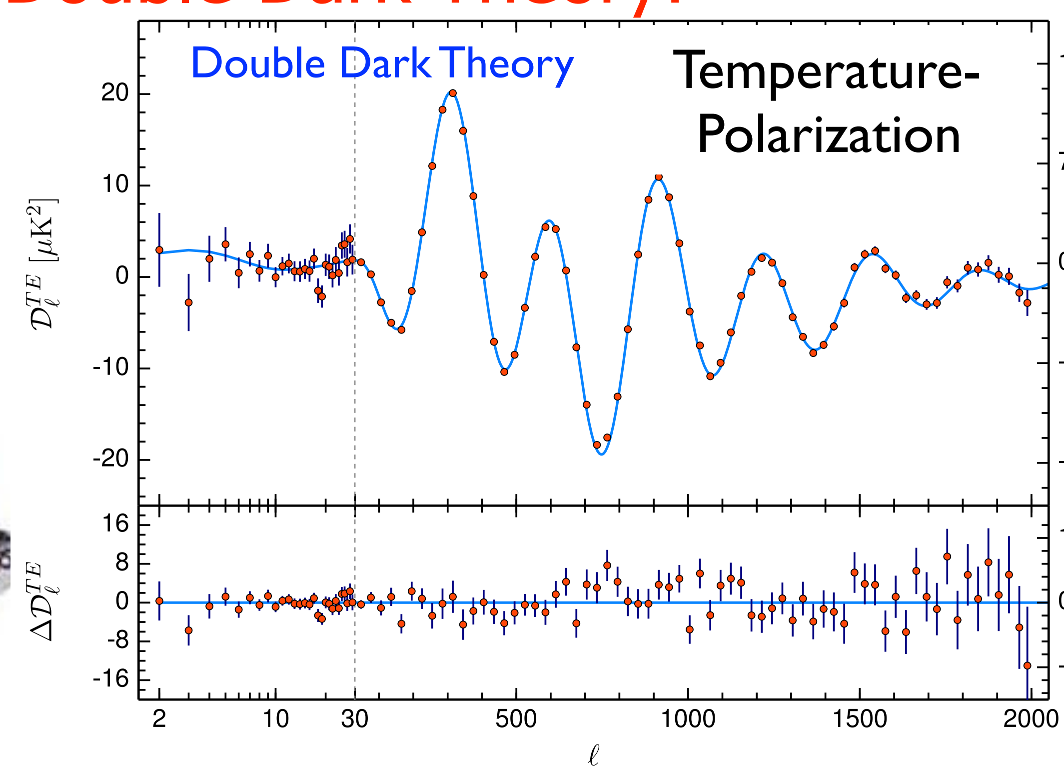
Cosmic Background Radiation

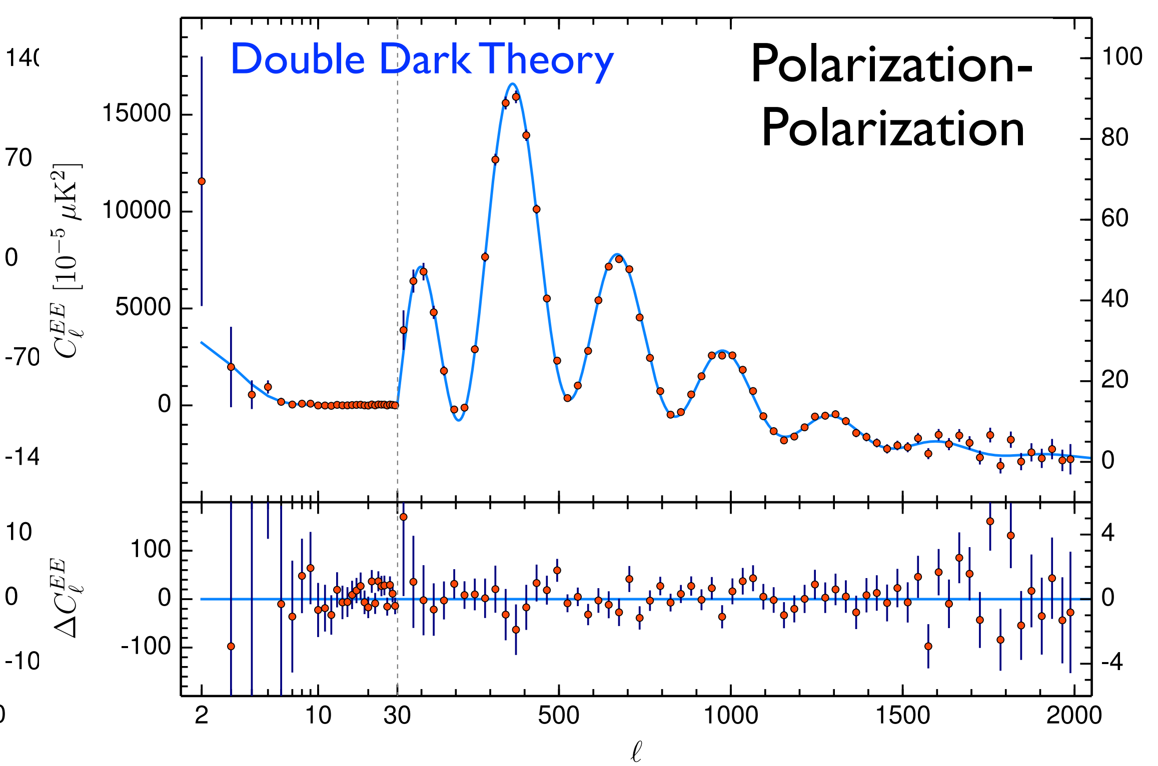
European Space Agency PLANCK Satellite Data

Released September 24, 2019



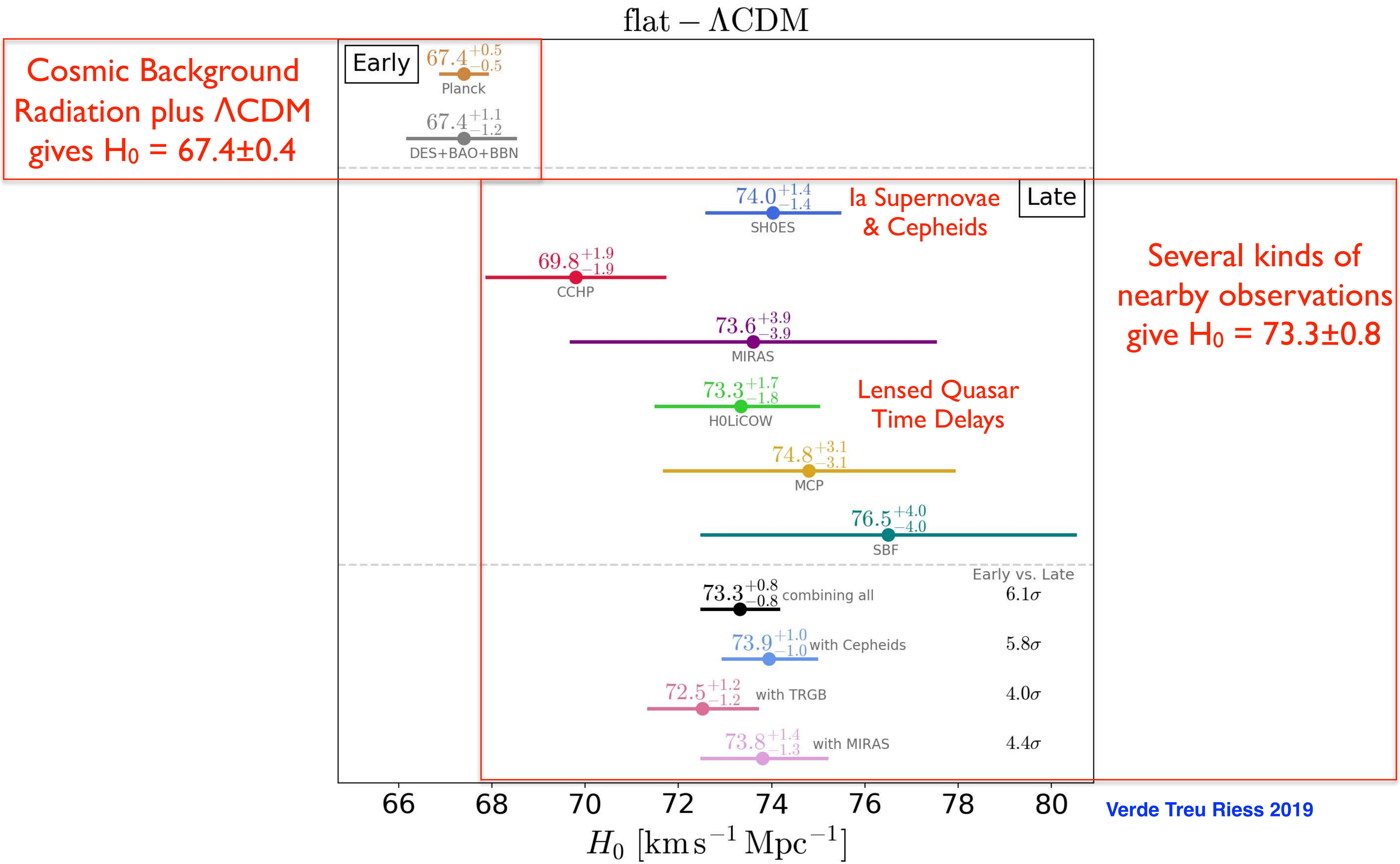
Agrees with Double Dark Theory!





The Hubble parameter H_0 is the expansion rate of the universe today.

A possibly serious difficulty for ΛCDM is the **Hubble parameter tension:**



"Early Dark Energy," a brief period of ≲10% extra dark energy at z ~ 3500, could resolve this

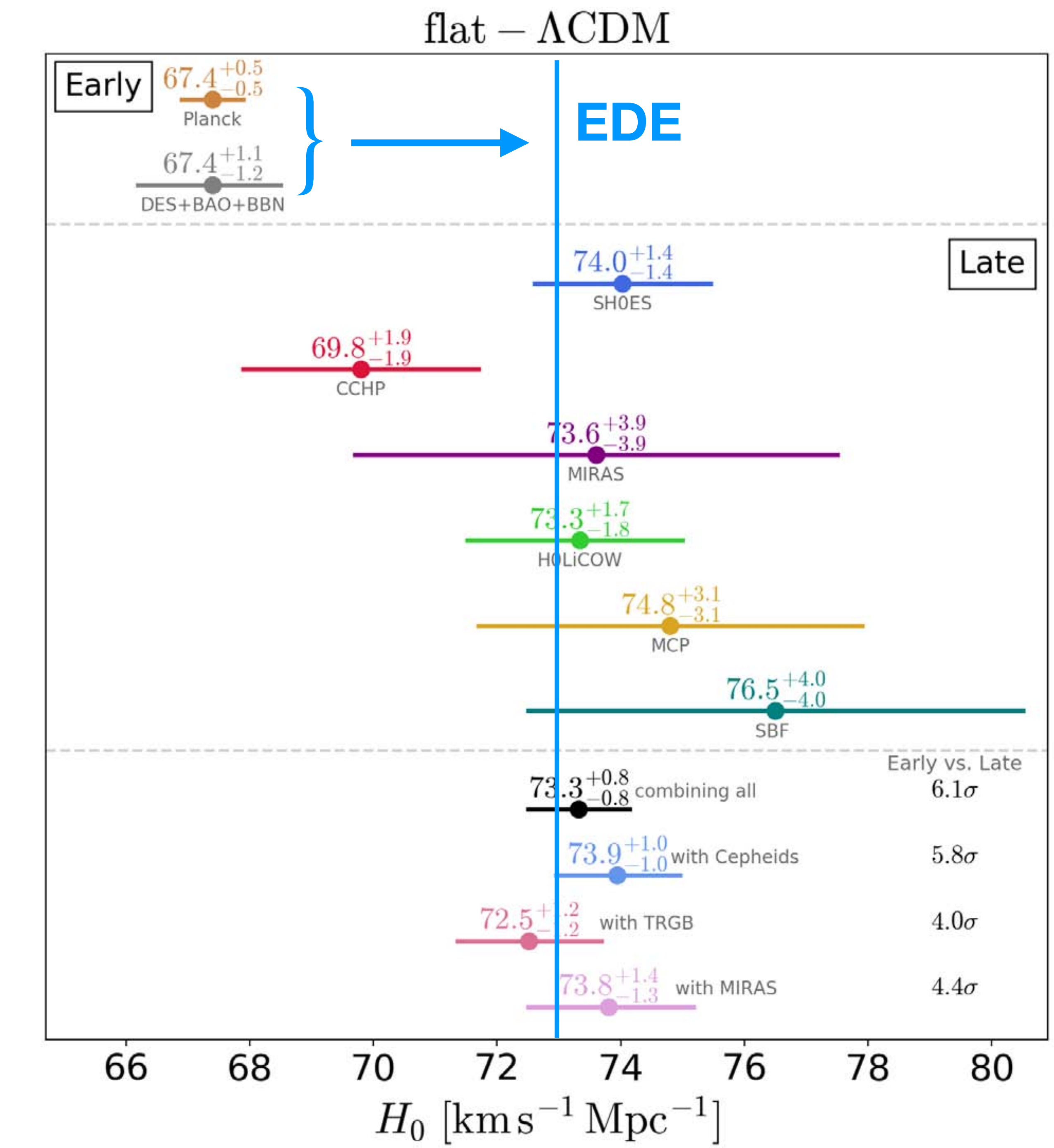
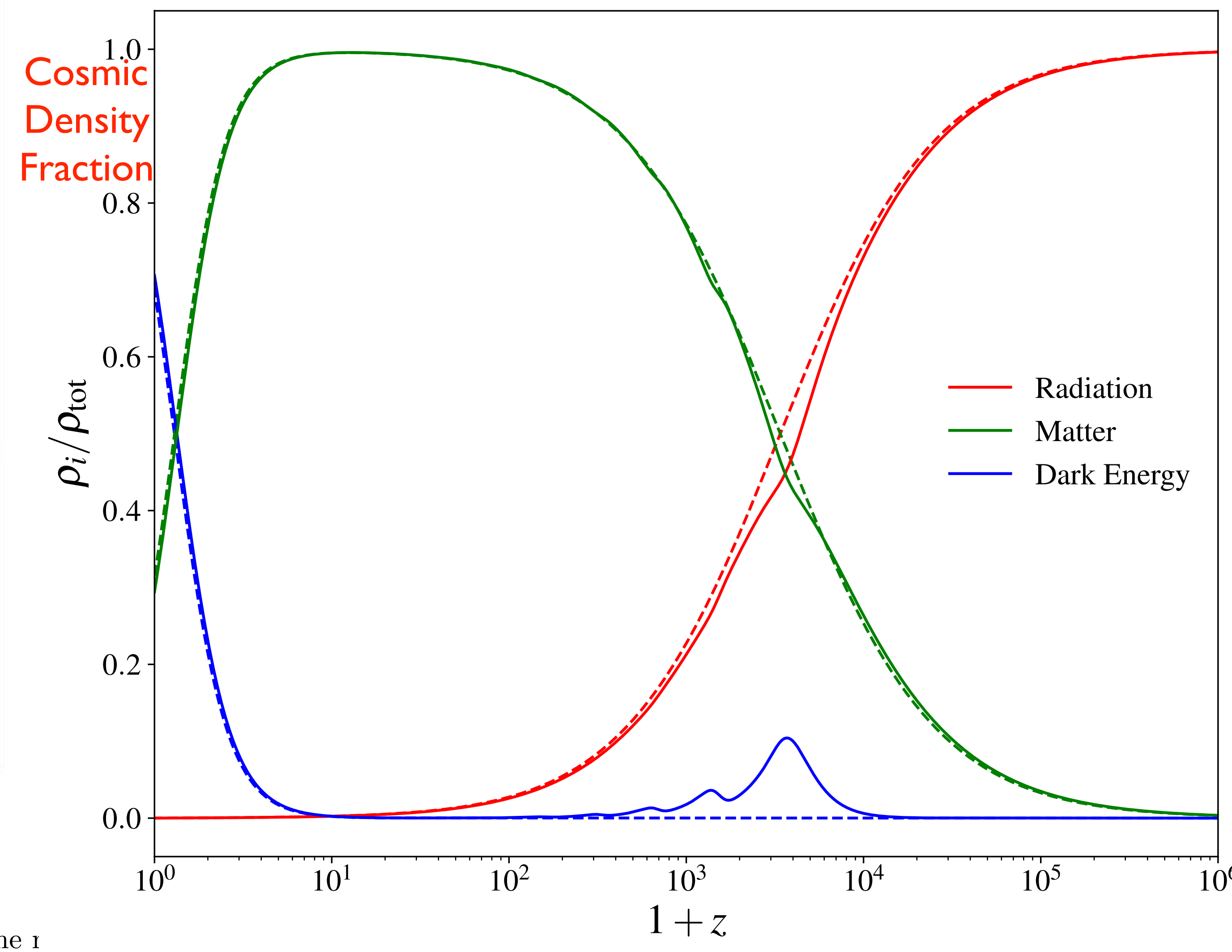


Figure 1. Compilation of Hubble Constant predictions and measurements taken from the recent literature and presented or discussed at the meeting. Two independent predictions based on early-Universe data (Planck Collaboration et al. 2018; Abbott et al. 2018) are shown at the top left (more utilizing other CMB experiments have been presented with similar findings), while the middle panel shows late Universe measurements. The bottom panel shows combinations of the late-Universe measurements and lists the tension with the early-Universe predictions. We stress that the three variants of the local distance ladder method (SHOES=Cepheids; CCHP=TRGB; MIRAS) share some Ia calibrators and cannot be considered as statistically independent. Likewise the SBF method is calibrated based on Cepheids or TRGB and thus it cannot be considered as fully independent of the local distance ladder method. Thus the "combining all" value should be taken for illustration only, since its derivation neglects covariance between the data. The three combinations based on Cepheids, TRGB, Miras are based on statistically independent datasets and therefore the significance of their discrepancy with the early universe prediction is correct - even though of course separating the probes gives up some precision. A fair summary is that the difference is more than 4 σ, less than 6 σ, while robust to exclusion of any one method, team or source. Figure courtesy of Vivien Bonvin.

Verde, Treu, Riess 2019

A brief episode of Early Dark Energy about ~ 50,000 years after the Big Bang modifies the ΛCDM extrapolation of H_0 and resolves the Hubble tension.



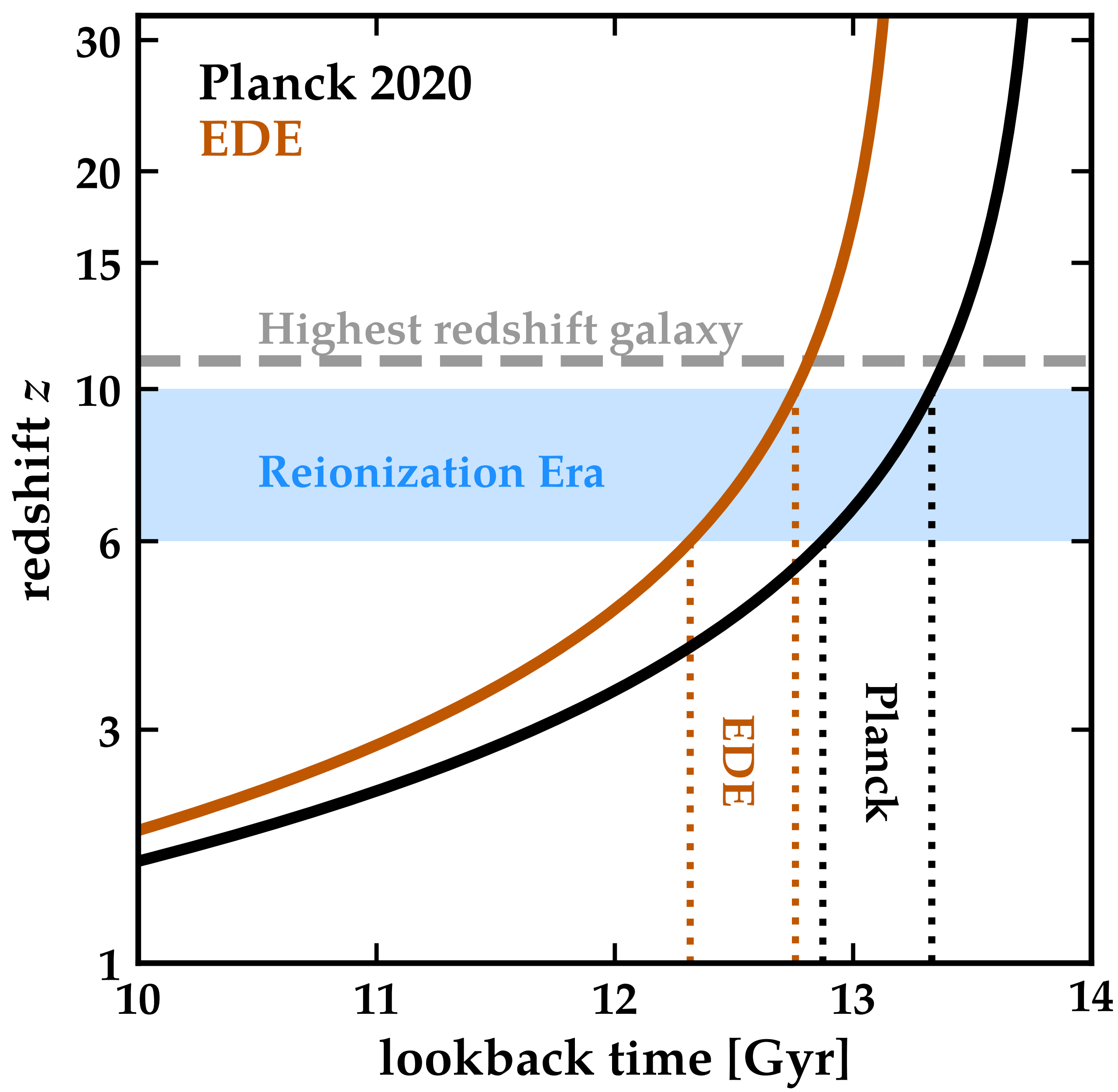
Solid curves represent our ΛCDM+EDE model, and dashed curves are standard ΛCDM with the Planck parameters. Our N-body simulations show that structure forms earlier than in standard ΛCDM, but the present-day universe is very similar.

Klypin, Poulin, Prada, Primack, et al. MNRAS 2021

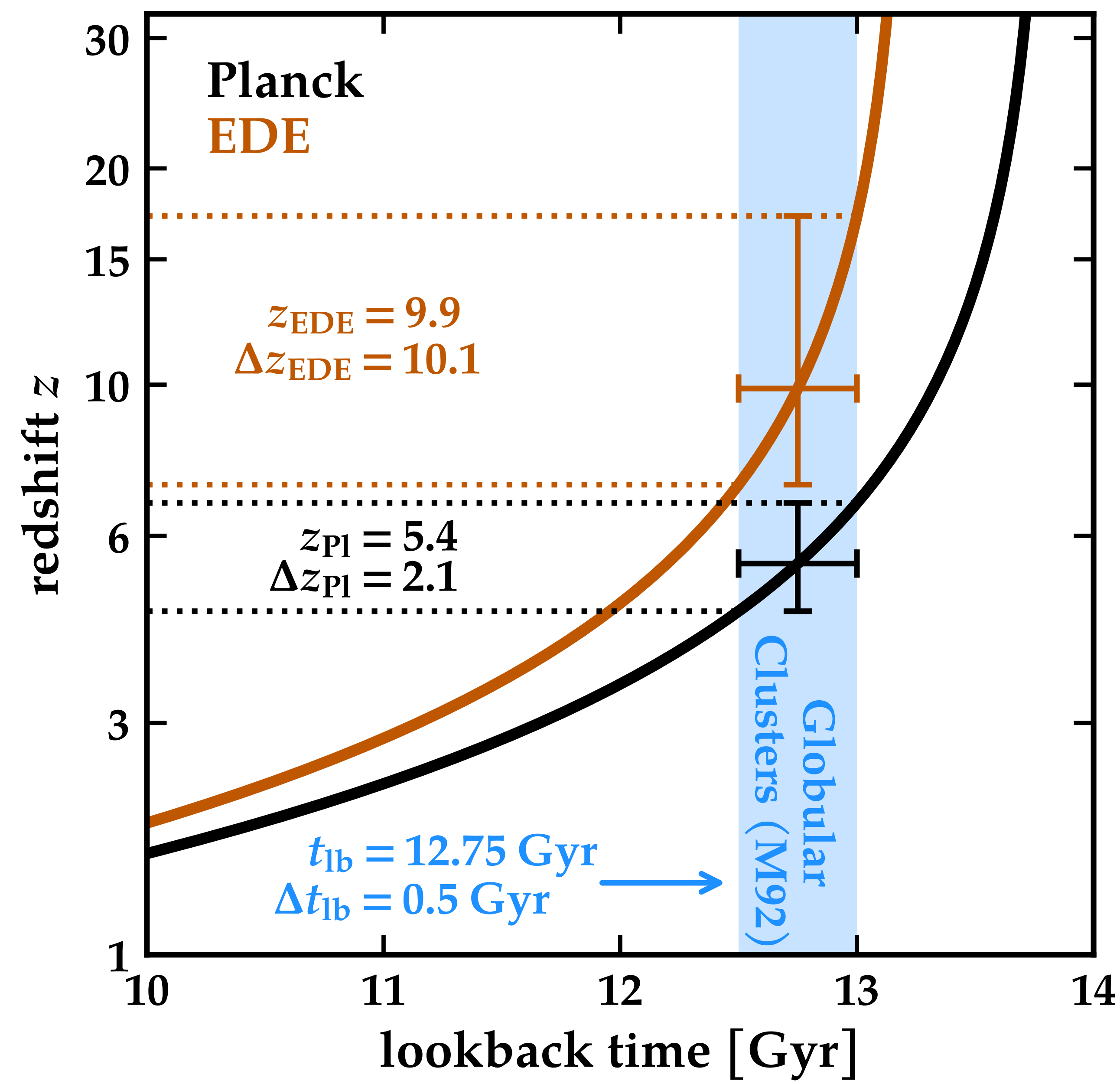
Early Dark Energy ==> age of the Universe t_0 ≈13.2 Gyr rather than Planck ΛCDM's 13.8 Gyr.

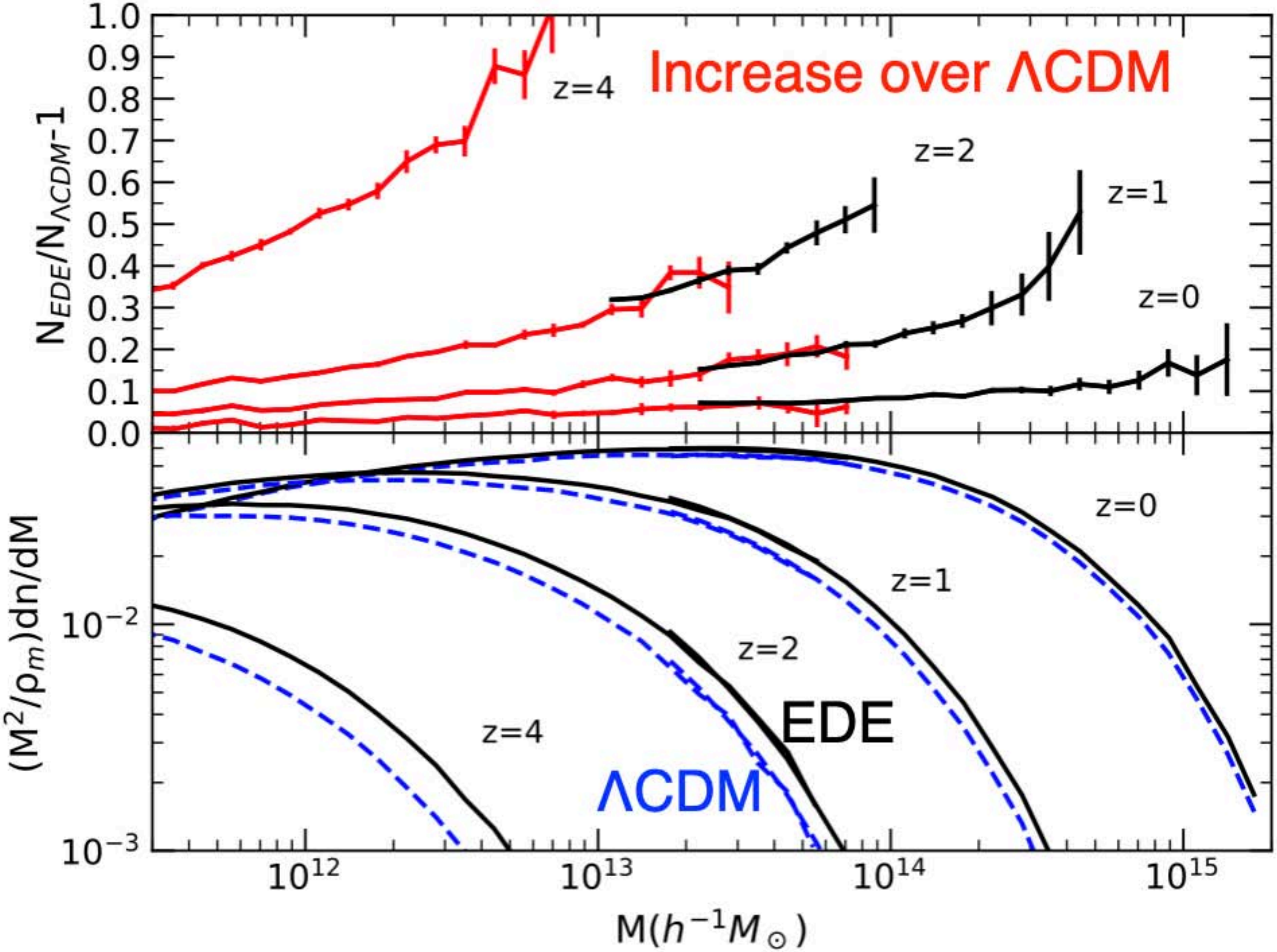
2021MNRAS.505.2764B by Michael Boylan-Kolchin and Dan Weisz shows that

The Reionization Era at z ≈ 6 - 10 corresponds to different cosmic ages for Planck ΛCDM and EDE:



Formation of >12.5 Gyr old Globular Cluster M92 corresponds to different redshifts z_{EDE} ≈ 10 vs. z_{Pl} ≈ 5.4:

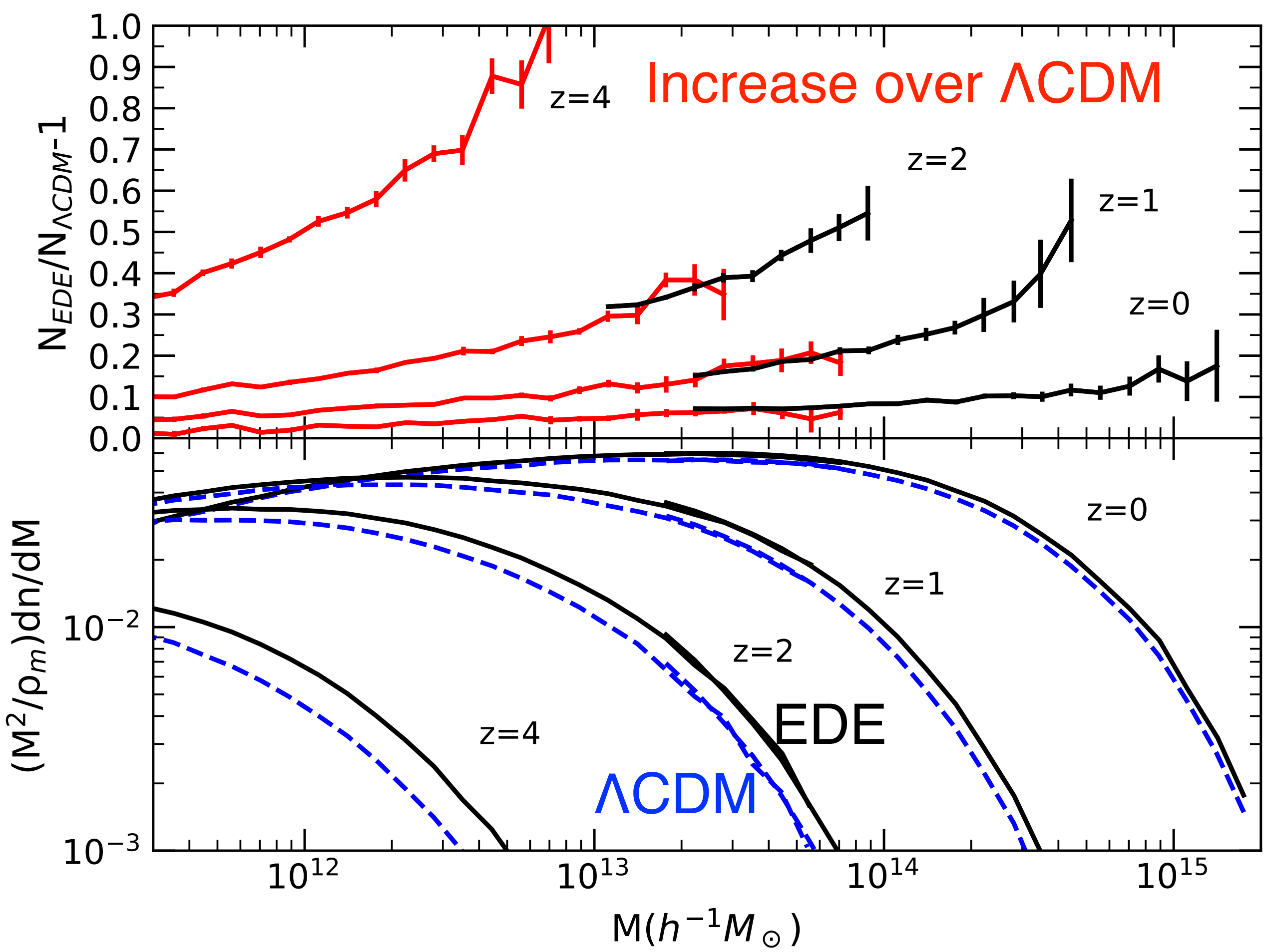


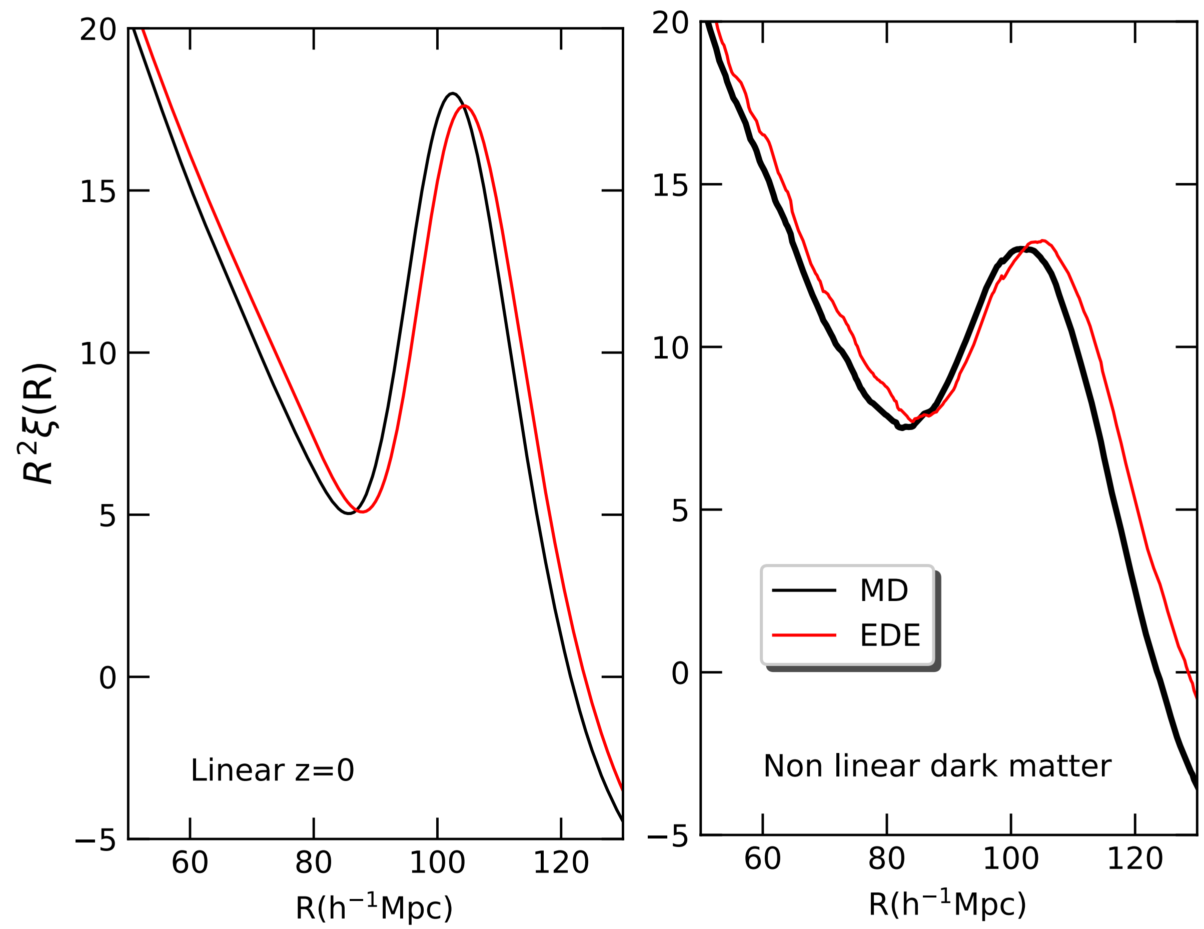


The EDE cosmology results in significantly earlier structure formation than standard ΛCDM, for example increasing the abundance of cluster-mass halos at $z \sim 1$ by ~ 50% and massive galaxies at $z \sim 4$ by ~ 2x. EDE also changes galaxy clustering, including increasing the baryon acoustic oscillation length scale but decreasing the correlations of nearby galaxies (Klypin et al. 2021).

	Model	H_0	Ω_m	Ω_b	t_0 (Gyr)	$A_s(\times 10^{-9})$	n_s	σ_8
EDE Models	Smith+2020	72.8	0.2915	0.0425	13.05	2.191	0.9860	0.836
	Murgia+2021	72.0	0.3009	0.0441	13.08	2.135	0.9895	0.837
	Agrawal+2019	70.5	0.302	0.0461	13.34	2.200	0.981	0.841
	Lin+2019	70.2	0.2981	0.0461	13.45	2.178	0.9832	
	Niedermann+2020	71.5	0.2999	0.0444	13.18	2.150	0.9912	0.841
ΛCDM	Planck 2020	67.37	0.3147	0.0492	13.80	2.097	0.9652	0.810

Cosmological parameters of five recent EDE models that relieve the Hubble tension, compared with the best-fit ΛCDM parameters from the final Planck TT,TE,EE + lowE + lensing analysis. The simulation results below use the Smith et al. (2020) EDE model. The similar Murgia+2021 parameters are in no worse agreement with gravitational lensing than standard ΛCDM.



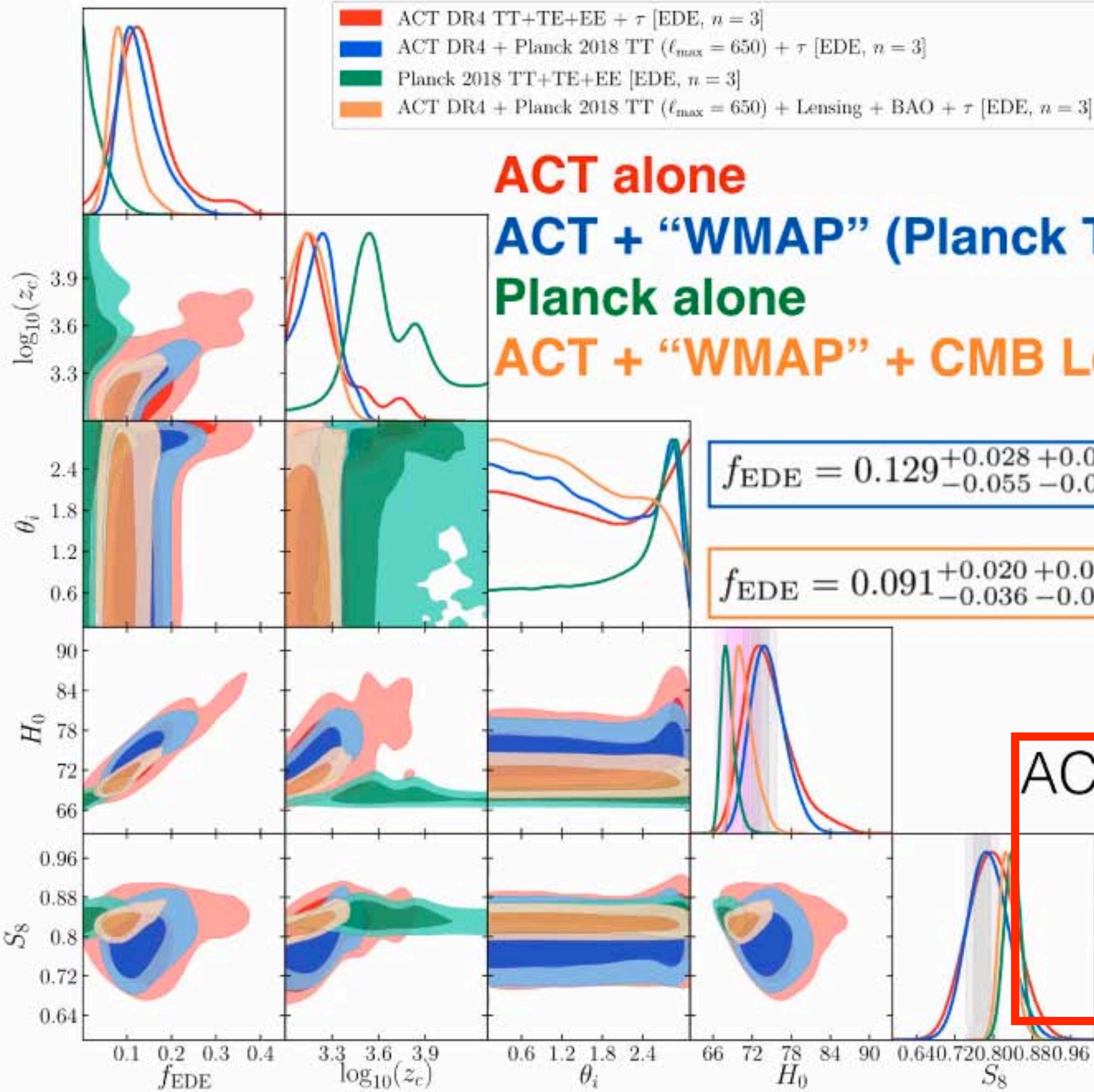


The EDE cosmology results in significantly earlier structure formation than standard ΛCDM, and it increases the BAO length scale but decreases the correlations of nearby galaxies (Klypin et al. 2021). Higher-resolution is needed for merger trees and substructure comparisons with ΛCDM. Tomo Ishiyama may be able to run paired (0.5 Gpc)³ ΛCDM and EDE simulations on the Fugaku supercomputer.

ACT DR4 EDE Results

Colin Hill
Columbia/CCA

Aspen Seminar
9 September 21



ACT alone

ACT + “WMAP” (Planck TT, ell<650)

Planck alone

ACT + “WMAP” + CMB Lensing + BAO

$$f_{\rm EDE} = 0.129^{+0.028\,+0.099\,+0.14}_{-0.055\,-0.076\,-0.084} \quad (68\%/95\%/99.7\%\ \rm CL)$$

$$f_{\rm EDE} = 0.091^{+0.020\,+0.069\,+0.11}_{-0.036\,-0.056\,-0.063} \quad (68\%/95\%/99.7\%\ \rm CL)$$

$$H_0 = 70.9^{+1.0}_{-2.0}\ \rm km/s/Mpc$$

ACT drives preference for non-zero EDE (>99.7% CL in joint fits)

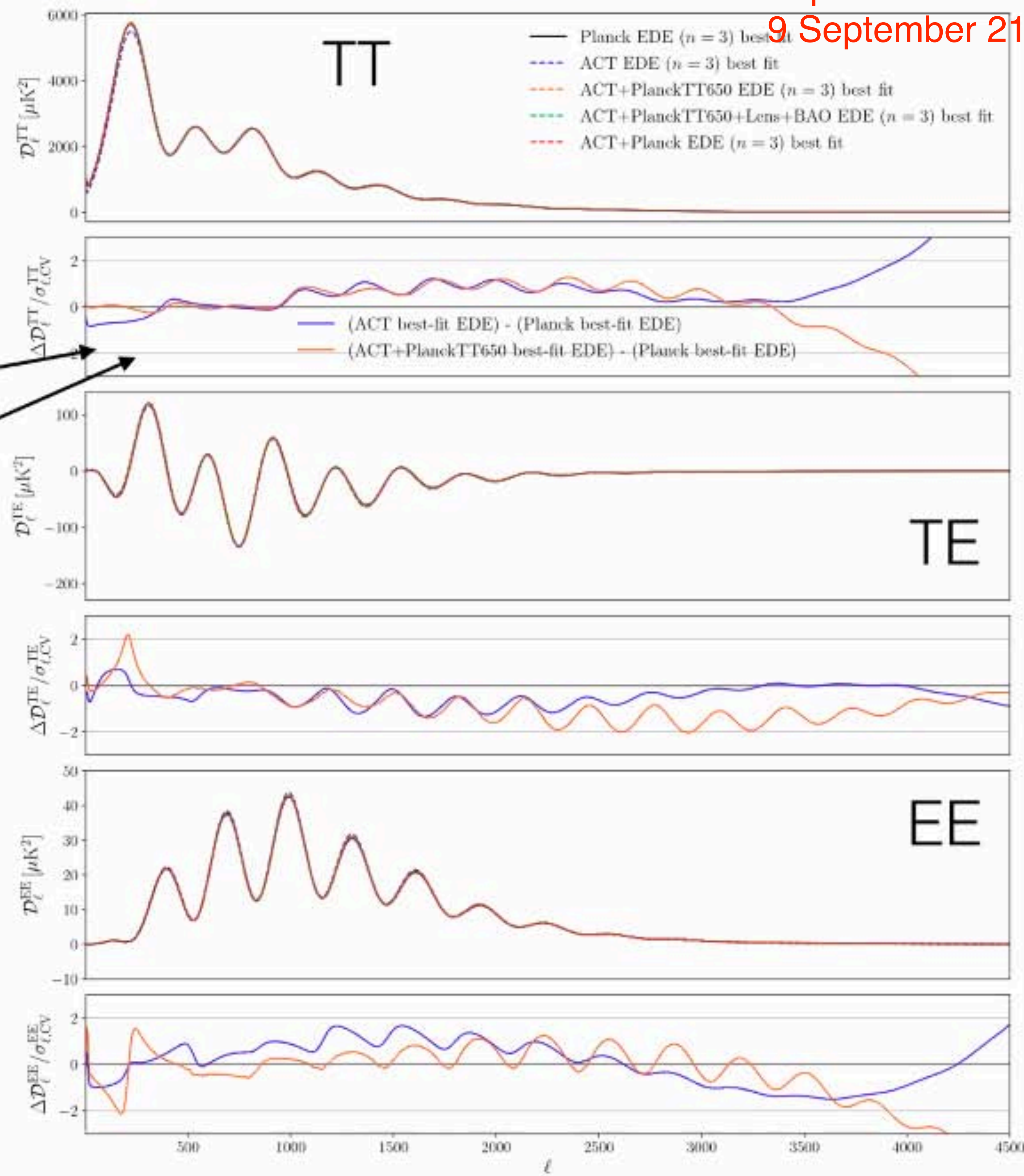
Discovering EDE in the CMB

Colin Hill
Columbia/CCA
Aspen Seminar
9 September 21

EDE (/other new physics) leaves signatures in small-scale CMB

ACT best-fit EDE - Planck EDE

ACT+"WMAP" EDE - Planck EDE



Imminent potential discovery with upcoming ACT DR6 (~2022): the models shown here can be distinguished at >20σ

JCH et al. (2021)

The Atacama Cosmology Telescope: Constraints on Pre-Recombination Early Dark Energy

J. Colin Hill,[1,2] Erminia Calabrese,[3] Simone Aiola,[2] Nicholas Battaglia,[4] Boris Bolliet,[1]

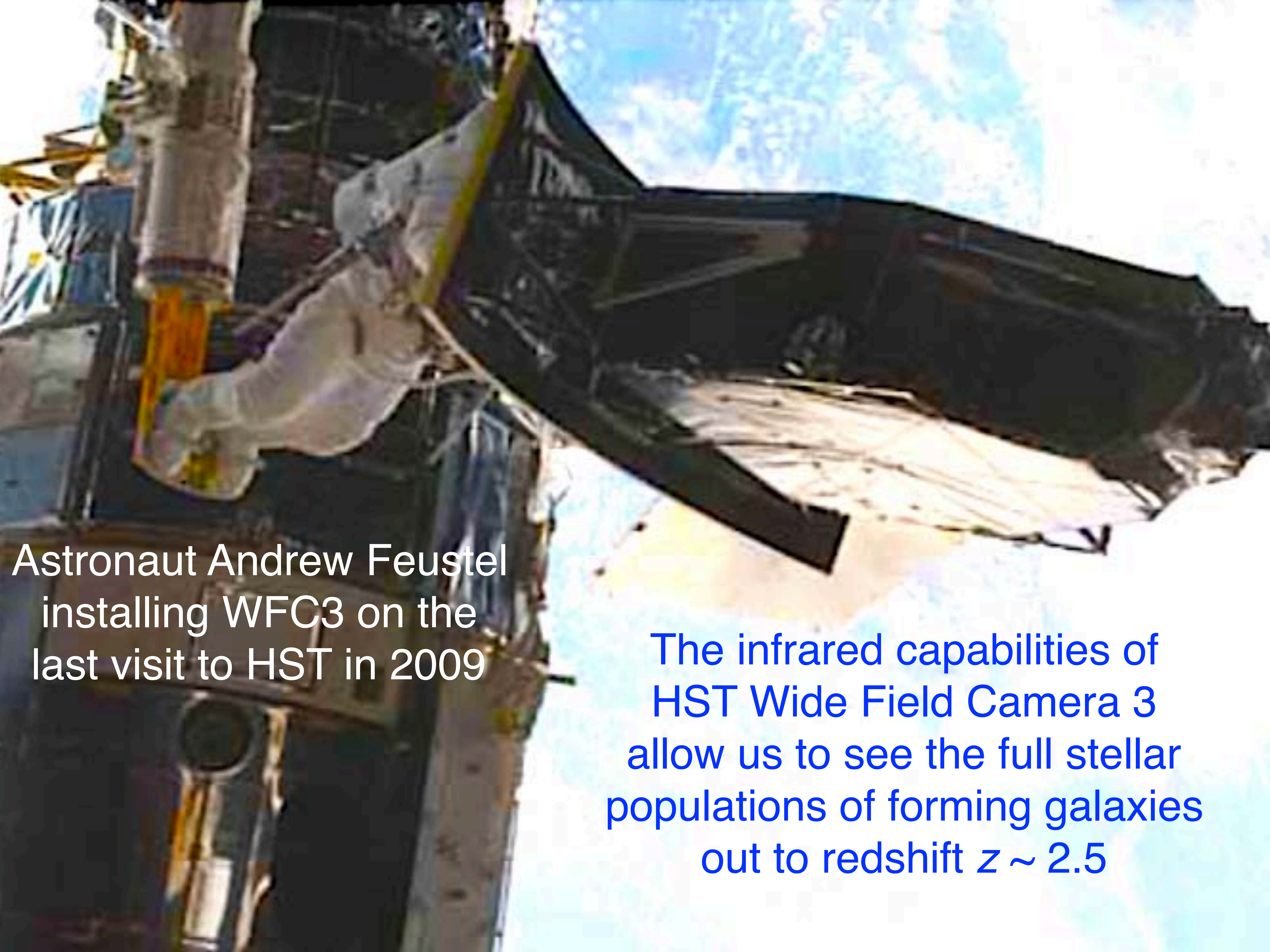
The early dark energy (EDE) scenario aims to increase the value of the Hubble constant (H_0) inferred from cosmic microwave background (CMB) data over that found in the standard cosmological model (ΛCDM), via the introduction of a new form of energy density in the early universe. The EDE component briefly accelerates cosmic expansion just prior to recombination, which reduces the physical size of the sound horizon imprinted in the CMB. Previous work has found that non-zero EDE is not preferred by *Planck* CMB power spectrum data alone, which yield a 95% confidence level (CL) upper limit $f_{\rm EDE} < 0.087$ on the maximal fractional contribution of the EDE field to the cosmic energy budget. In this paper, we fit the EDE model to CMB data from the Atacama Cosmology Telescope (ACT) Data Release 4. We find that a combination of ACT, large-scale *Planck* TT (similar to *WMAP*), *Planck* CMB lensing, and BAO data prefers the existence of EDE at $> 99.7\%$ CL: $f_{\rm EDE} = 0.091^{+0.020}_{-0.036}$, with $H_0 = 70.9^{+1.0}_{-2.0}$ km/s/Mpc (both 68% CL). From a model-selection standpoint, we find that EDE is favored over ΛCDM by these data at roughly 3σ significance. In contrast, a joint analysis of the full *Planck* and ACT data yields no evidence for EDE, as previously found for *Planck* alone. We show that the preference for EDE in ACT alone is driven by its TE and EE power spectrum data. The tight constraint on EDE from *Planck* alone is driven by its high-ℓ TT power spectrum data. Understanding whether these differing constraints are physical in nature, due to systematics, or simply a rare statistical fluctuation is of high priority. The best-fit EDE models to ACT and *Planck* exhibit coherent differences across a wide range of multipoles in TE and EE, indicating that a powerful test of this scenario is anticipated with near-future data from ACT and other ground-based experiments.


NEW CHALLENGES IN
COSMOLOGY
GALAXY FORMATION

Almost all the stars today are in large galaxies like our Milky Way. Nearby large galaxies are disk galaxies like our galaxy or big balls of stars called elliptical galaxies. But most galaxies in the early universe didn't look anything like our Milky Way. Many of them are pickle-shaped and clumpy.

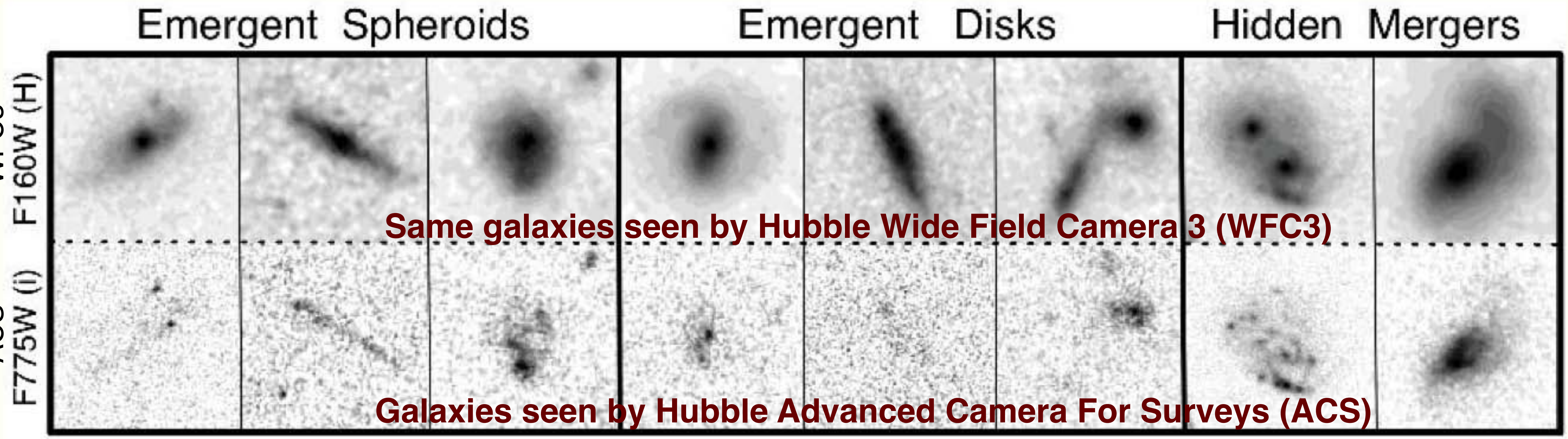



We are just now figuring out how galaxies form and evolve with the help of big ground-based telescopes, and Hubble and other space telescopes that let us see radiation clearly without interference from earth's atmosphere.


Astronaut Andrew Feustel installing WFC3 on the last visit to HST in 2009
The infrared capabilities of HST Wide Field Camera 3 allow us to see the full stellar populations of forming galaxies out to redshift $z \sim 2.5$

The CANDELS Survey shows shapes of z ≲ 2.5 galaxies

candels.ucolick.org



(blue 0.4 μm)(1+z) = 1.6 μm @ z = 3

(orange 0.6 μm)(1+z) = 1.6 μm @ z = 1.7

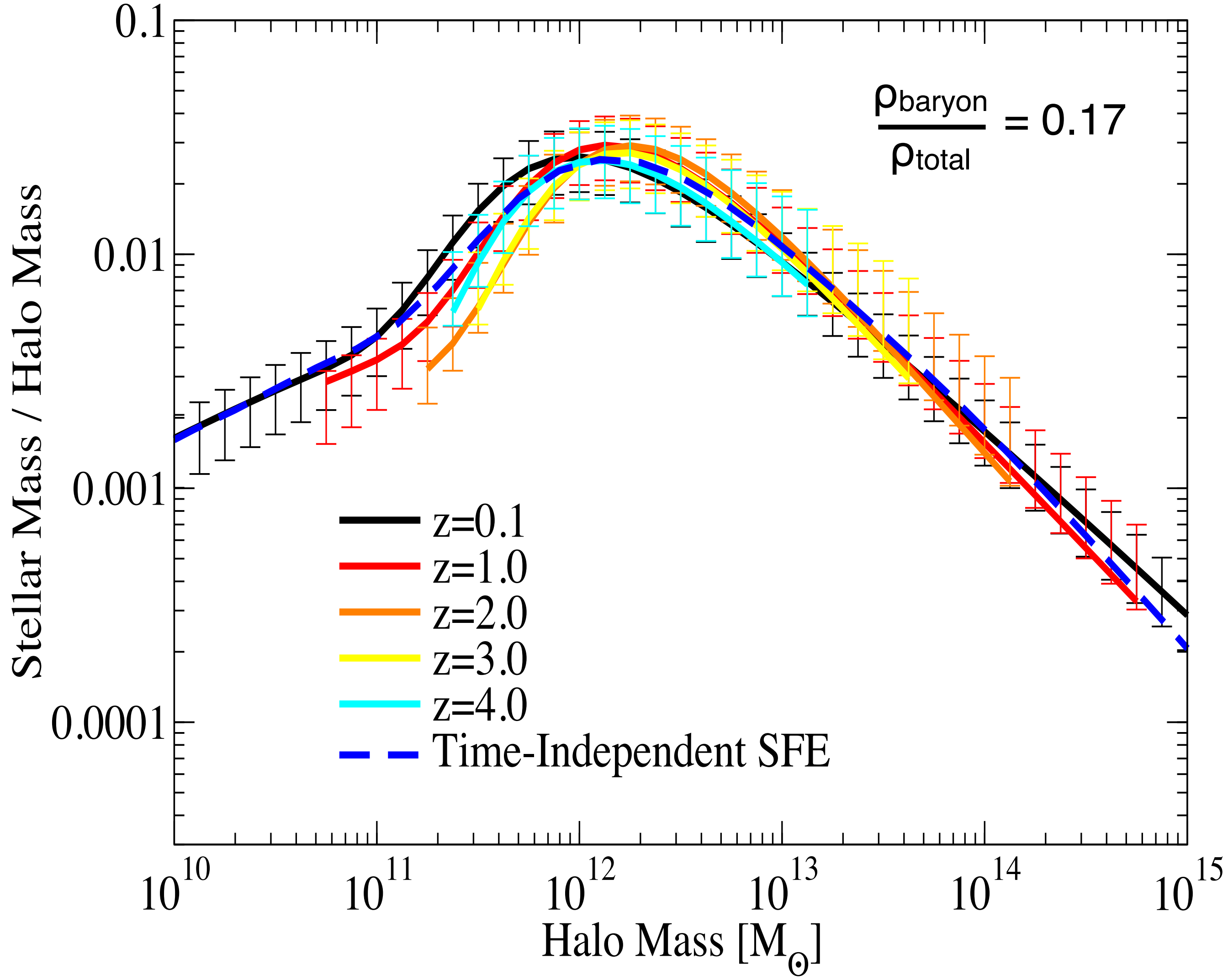
CANDELS: A Cosmic Odyssey

CANDELS is a powerful imaging survey of the distant Universe being carried out with two cameras on board the Hubble Space Telescope.

- CANDELS is the largest project in the history of Hubble, with 902 assigned orbits of observing time. This is the equivalent of four months of Hubble time if executed consecutively, but in practice CANDELS will take three years to complete (2010-2013).
- The core of CANDELS is the revolutionary near-infrared WFC3 camera, installed on Hubble in May 2009. WFC3 is sensitive to longer, redder wavelengths, which permits it to follow the stretching of lightwaves caused by the expanding Universe. This enables CANDELS to detect and measure objects much farther out in space and nearer to the Big Bang than before. CANDELS also uses the visible-light ACS camera, and together the two cameras give unprecedented panchromatic coverage of galaxies from optical wavelengths to the near-IR.
- CANDELS will exploit this new lookback power to construct a "cosmic movie" of galaxy evolution that follows the life histories of galaxies from infancy to the present time. This work will cap Hubble's revolutionary series of discoveries on cosmic evolution and bequeath a legacy of precious data to future generations of astronomers.

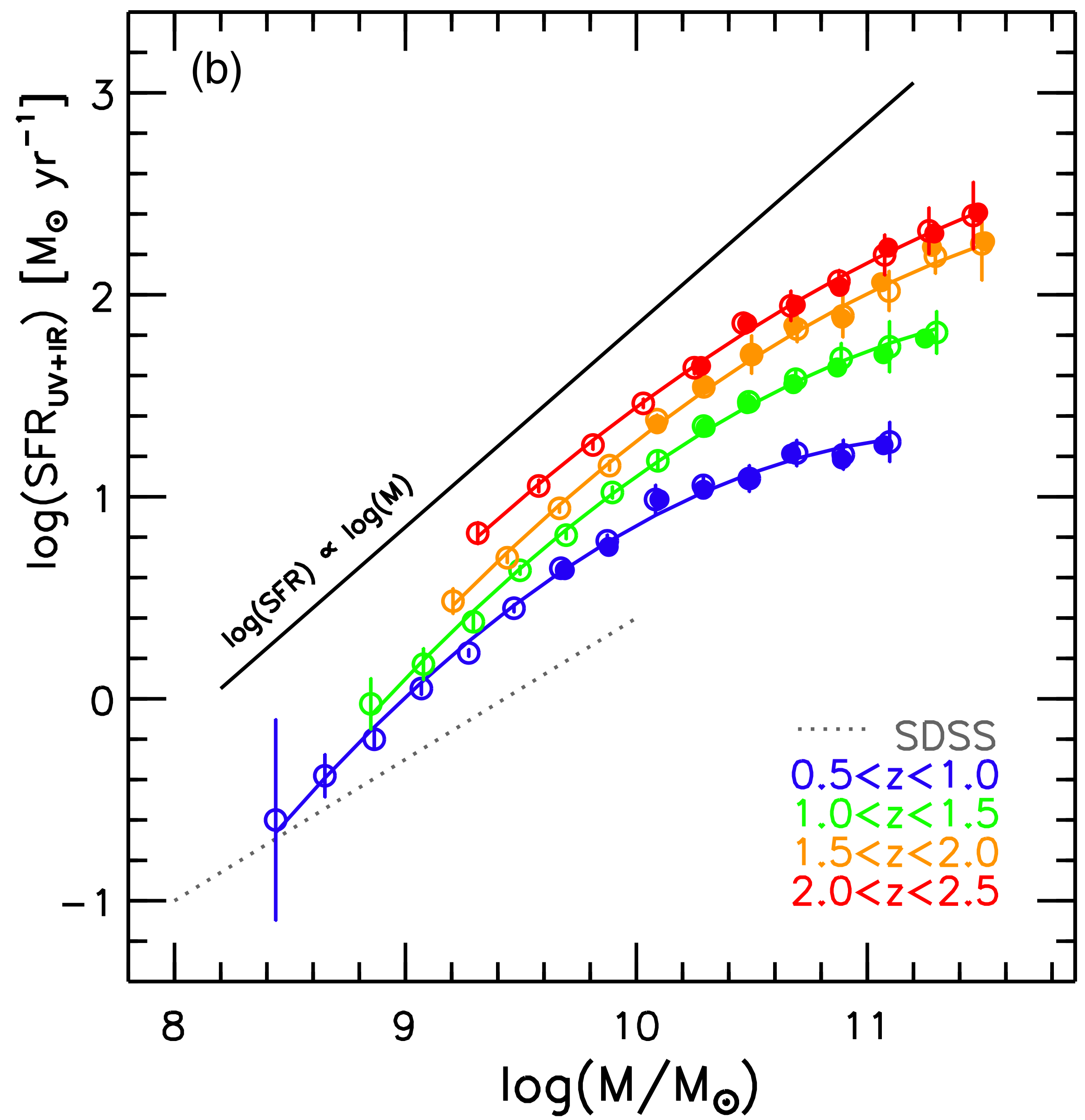
Two Key Discoveries About Galaxies

Relationship Between Galaxy Stellar Mass and Halo Mass



The stellar mass to halo mass ratio at multiple redshifts as derived from observations compared to the Bolshoi cosmological simulation. Error bars show 1σ uncertainties. A time-independent Star Formation Efficiency predicts a roughly **time-independent stellar mass to halo mass relationship**. (Behroozi, Wechsler, Conroy, ApJL 2013)

Star-forming Galaxies Lie on a "Main Sequence"

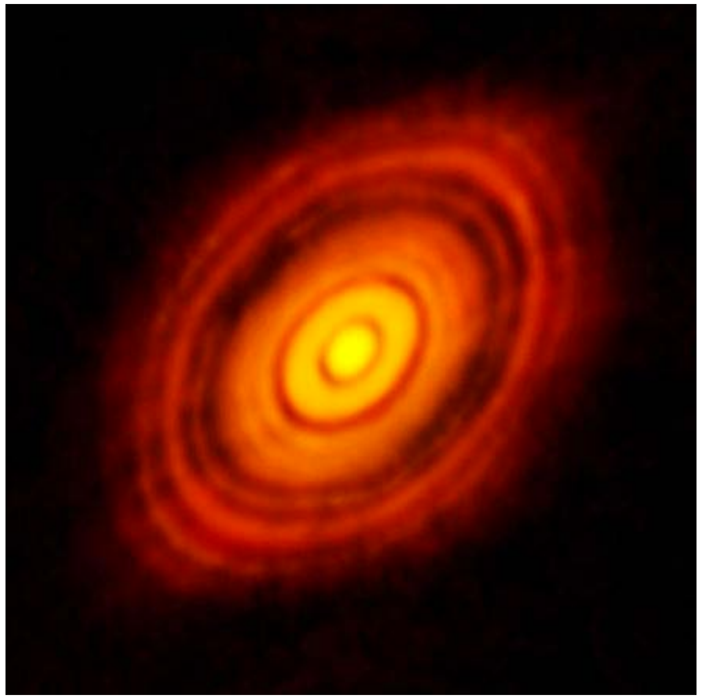


Just as the properties of hydrogen-burning stars are controlled by their mass, **the galaxy star formation rate (SFR) is approximately proportional to the stellar mass**, with the proportionality constant increasing with redshift up to about $z = 2.5$. (Whitaker et al. ApJ 2014)

Do Galaxies Start as Disks?

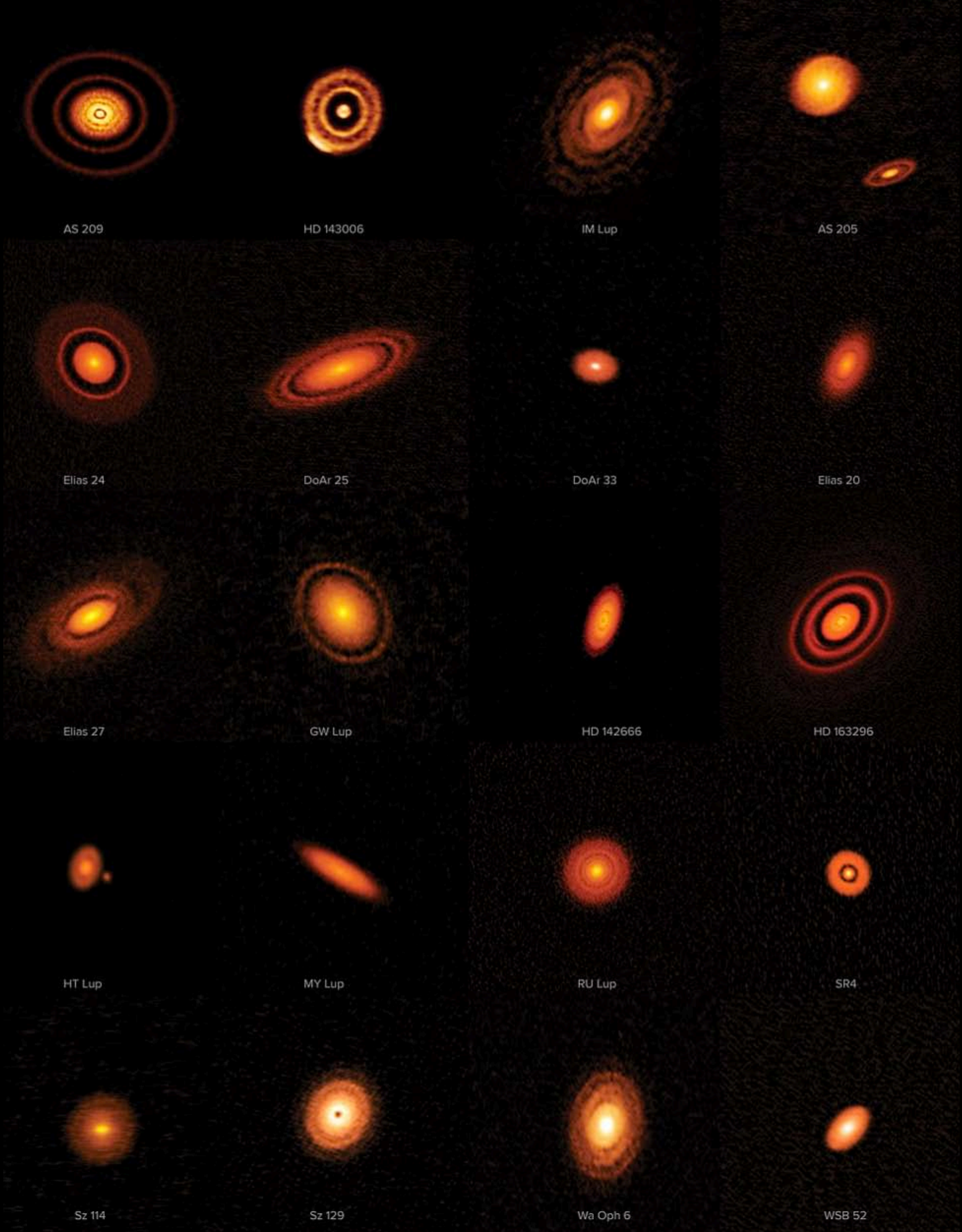
Newton's laws explained why planetary orbits are elliptical, but not why the planetary orbits in the solar system are nearly circular, in the same plane, and in the same direction as the sun rotates.

Laplace explained this as a consequence of angular momentum conservation as the sun and planets formed in a cooling and contracting protoplanetary gas cloud that formed a disk— like this one:



ALMA image of HL Tauri

20 Protoplanetary Disks from ALMA's High Angular Resolution Project DSHARP (2019)



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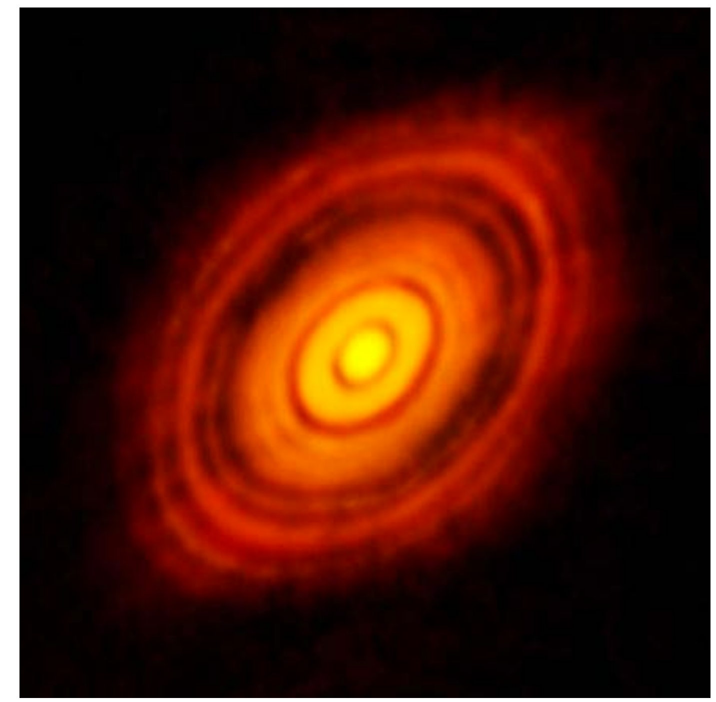


ALMA image of HL Tauri

For similar reasons, many astronomers once thought that galaxies would start as disks. But Hubble Space Telescope images of forming galaxies instead show that most forming galaxies are prolate – that is, pickle-shaped. As we will see, this is a consequence of most galaxies forming in prolate dark matter halos oriented along massive dark matter filaments.

MODERN COSMOLOGY

Ya B. Zeldovich

Institute of Physical Problem, Academy of Sciences of the USSR, Moscow.

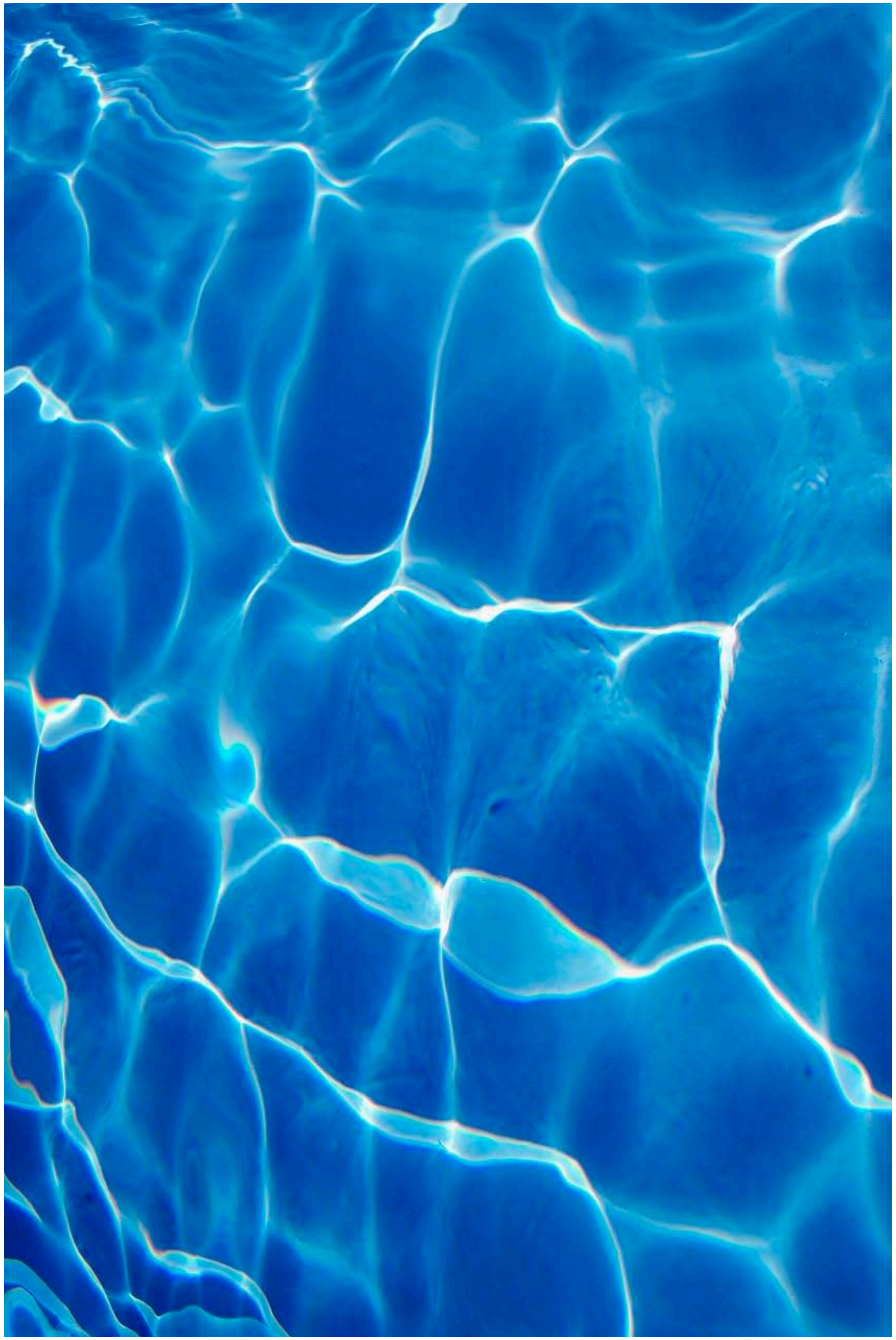
Cosmology, the study of the Universe as a whole, is perhaps the most difficult branch of astronomy, since there is always a danger of replacing true knowledge by prejudice, resulting from the impossibility of observing the whole Universe. The situation has changed during the last few decades.

Cosmology has become a respectable science, which was not so 50–60 years ago. However, the problems of the creation of the Universe, and with the reasons for its present form have not yet been solved. At the same time definite progress has been made in understanding the present state of the Universe and a number of its stages of evolution; this progress is a result of investigations carried out by many people, and joint efforts by numerous international groups of astronomers.

The pressure of netural gas can be neglected. Of course, this statement is not an absolute one: gas pressure can neglected in the case when the wavelength of density perturbations is sufficiently long. It is this legacy that we inherited from the radiation-dominated era. But then, if gas pressure does not play any role, the motion of gas turns out to be very specific: nothing prevents particles from coming close to each other to form high-density regions. In three-dimensional space gas can be compressed along each of the three independent directions perpendicular to each other. However, simultaneous compression along two or three axes occurs very rarely, and is not a typical phenomenon. As a rule, there is only one direction in each elementary volume which stands out among the rest.

Compression in this direction creates thin layers with a high density (they are jokingly called "pancakes"). Subsequent gas parcels colliding with a "pancake" heat up in the shock wave, i.e. "fly in". Besides, the "pancakes" grow along its plane. Of course, they are not absolutely flat, but that is not so important. At a later stage the "pancakes" begin to overlap, eventually forming a complex cell structure where compressed gas layers are surrounded by low-density regions.

Such a general picture of the cell structure of the Universe is supported by computer calculations, as well as by a rigorous mathematical analysis based on catastrophe theory and synergetics. An analogy has been established between gravitational instability and the laws of geometrical optics for light reflected from or refracted by stochastic waves at a water surface. (On a sunny day one can see patterns similar to those predicted by the "pancake" theory at the bottom of a swimming pool.) Obviously, galaxies should be created in compressed gas whose layers are still more exposed to the impact of further gravitational clustering.



Zel'dovich - *My Universe: Selected Reviews* 1992

MNRAS 367, 1781–1796 (2006)

The shape of dark matter haloes: dependence on mass, redshift, radius and formation

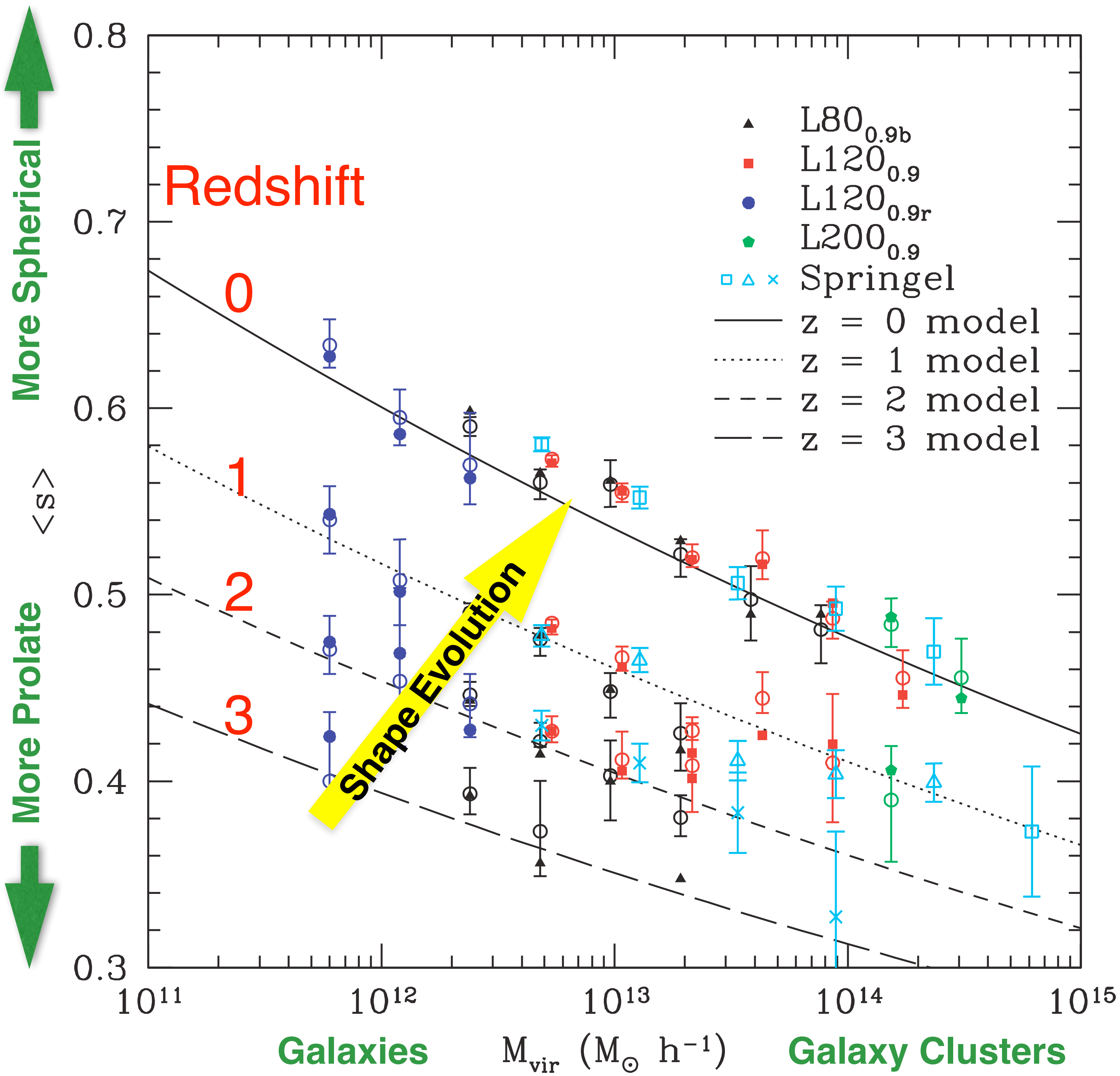
Brandon Allgood, Ricardo Flores, Joel R. Primack, Andrey V. Kravtsov, Risa Wechsler, Andreas Faltenbacher and James S. Bullock

$s = c/a$ = short axis / long axis

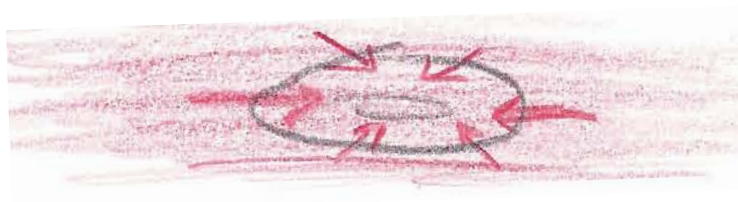
Halos are approximately triaxial ellipsoids

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} + \frac{z^2}{c^2} = 1 \qquad a \geq b \geq c$$

Halos start prolate, especially at low radius, and later become more spherical.



Low-redshift halo, accreting more spherically



High-redshift halo, accreting mainly along filament

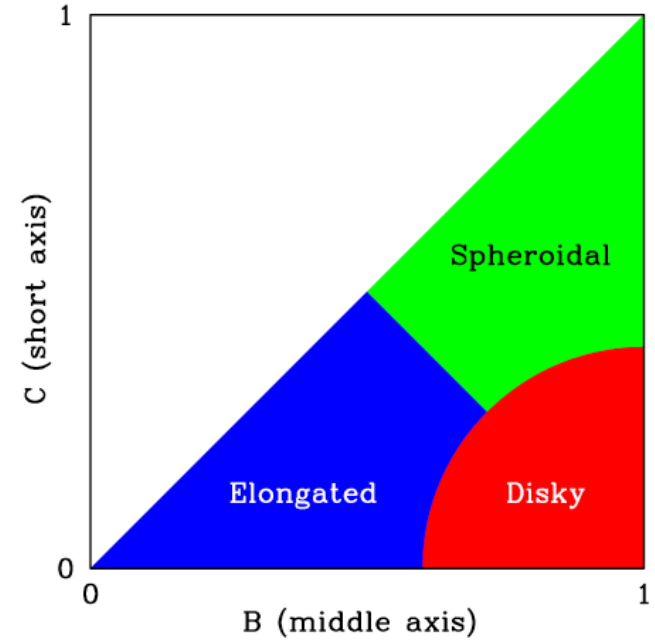
supported by anisotropic velocity dispersion, larger along principal axis

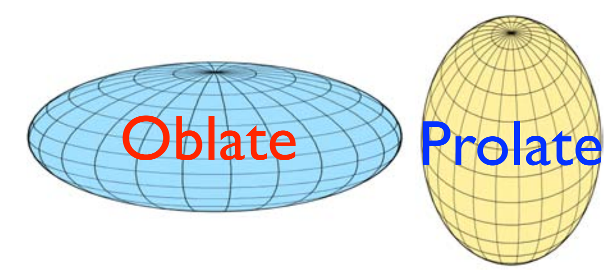
The Evolution of Galaxy Shapes in CANDELS: from Prolate to Oblate

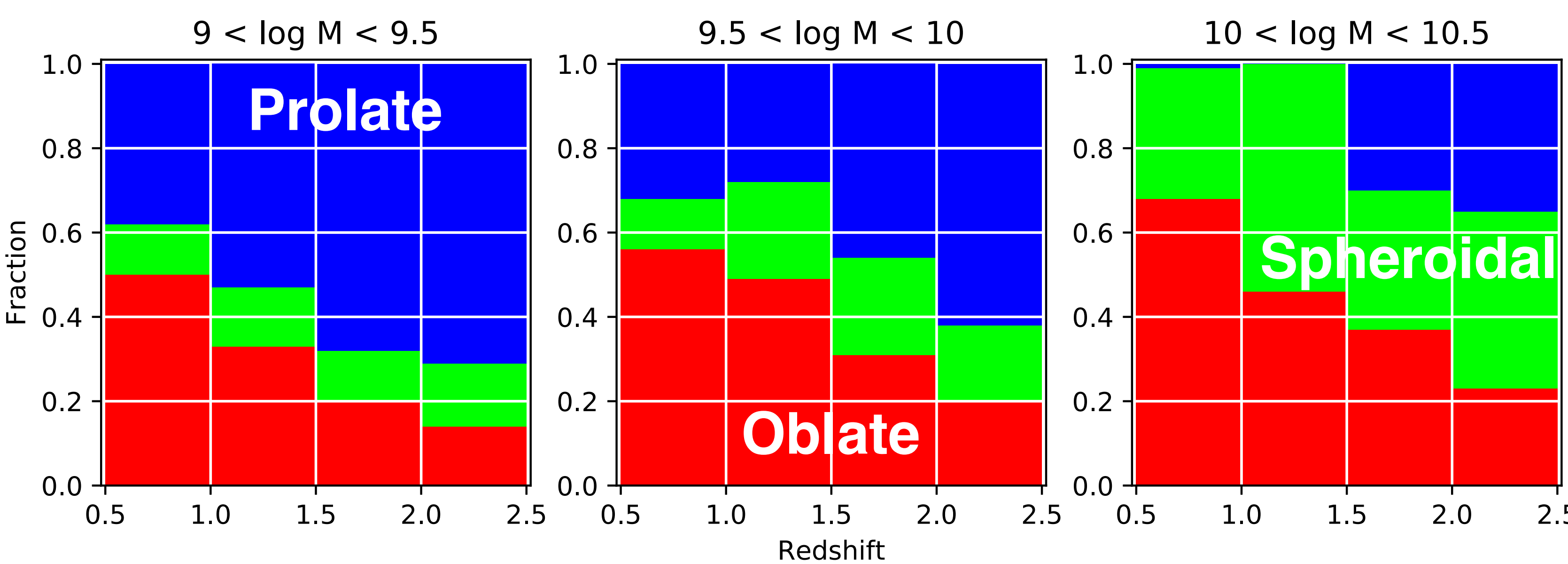
Haowen Zhang, Joel R. Primack, S. M. Faber, David C. Koo, Avishai Dekel, Zhu Chen, Daniel Ceverino, Yu-Yen Chang, Jerome J. Fang, Yicheng Guo, Lin Lin, and Arjen van der Wel MNRAS 484, 5170 (2019)

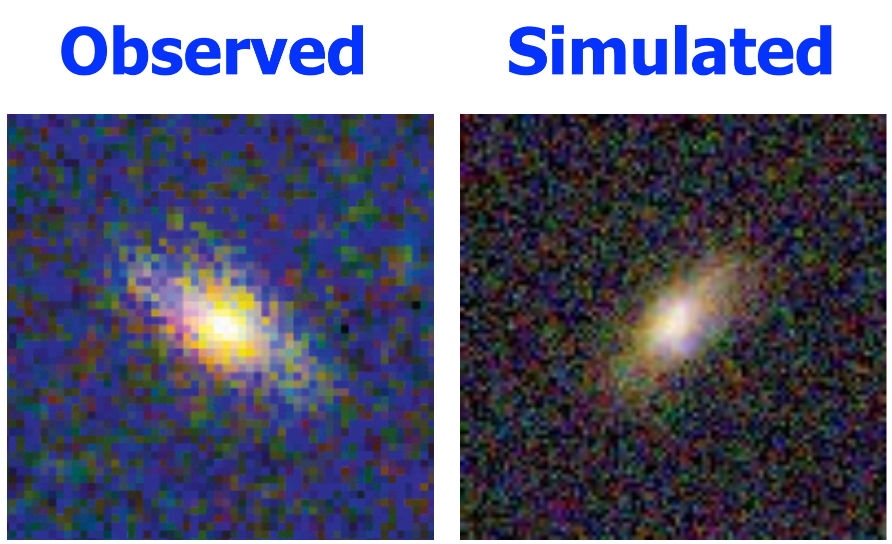
ABSTRACT

We model the projected b/a – $\log a$ distributions of CANDELS main sequence star-forming galaxies, where a (b) is the semi-major (semi-minor) axis of the galaxy images. We find that smaller-a galaxies are rounder at all stellar masses M_* and redshifts, so we include a when analyzing b/a distributions. Approximating intrinsic shapes of the galaxies as triaxial ellipsoids and assuming a multivariate normal distribution of galaxy size and two shape parameters, we construct their intrinsic shape and size distributions to obtain the fractions of prolate, oblate, and spheroidal galaxies in each redshift and mass bin. We find that galaxies tend to be prolate at low M_* and high redshifts, and oblate at high M_* and low redshifts, qualitatively consistent with van der Wel et al. (2014), implying that galaxies tend to evolve from prolate to oblate. These results are consistent with the predictions from simulations (Ceverino et al. 2015, Tomassetti et al. 2016) that the transition from prolate to oblate is caused by a compaction event at a characteristic mass range, making the galaxy center baryon dominated. We give probabilities of a galaxy's being prolate, oblate, or spheroidal as a function of its M_*, redshift, and projected b/a and a, which can facilitate target selections of galaxies with specific shapes at high redshifts. We also give predicted optical depths of galaxies, which are qualitatively consistent with the expected correlation that A_V should be higher for edge-on disk galaxies in each $\log a$ slice at low redshift and high mass bins.









(a) CANDELS galaxy

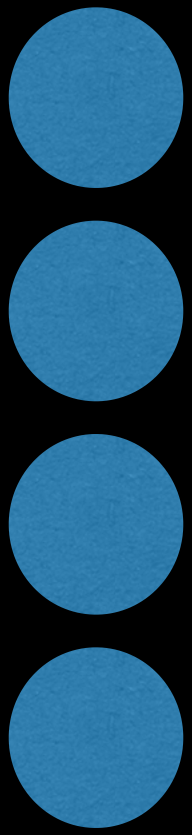
(b) VELA galaxy

Nearby large galaxies are mostly spheroids and disks — but they start out looking more like pickles.

How Can We Determine 3D Galaxy Shapes from 2D Telescope Images? Statistics!

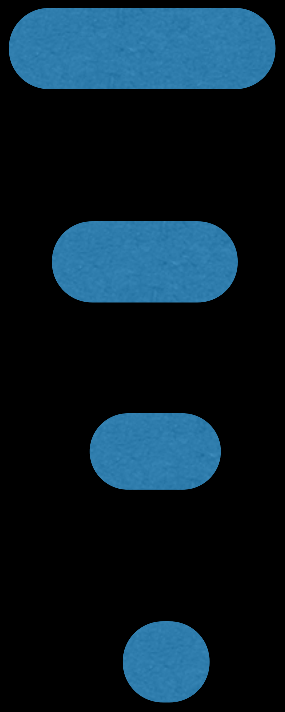
We see galaxies in all possible orientations

Let's orient them with their long axes horizontal and see the short/long axis ratio distribution



Spheroidal galaxies always have large axis ratio

Disk galaxies have even distribution of axis ratios

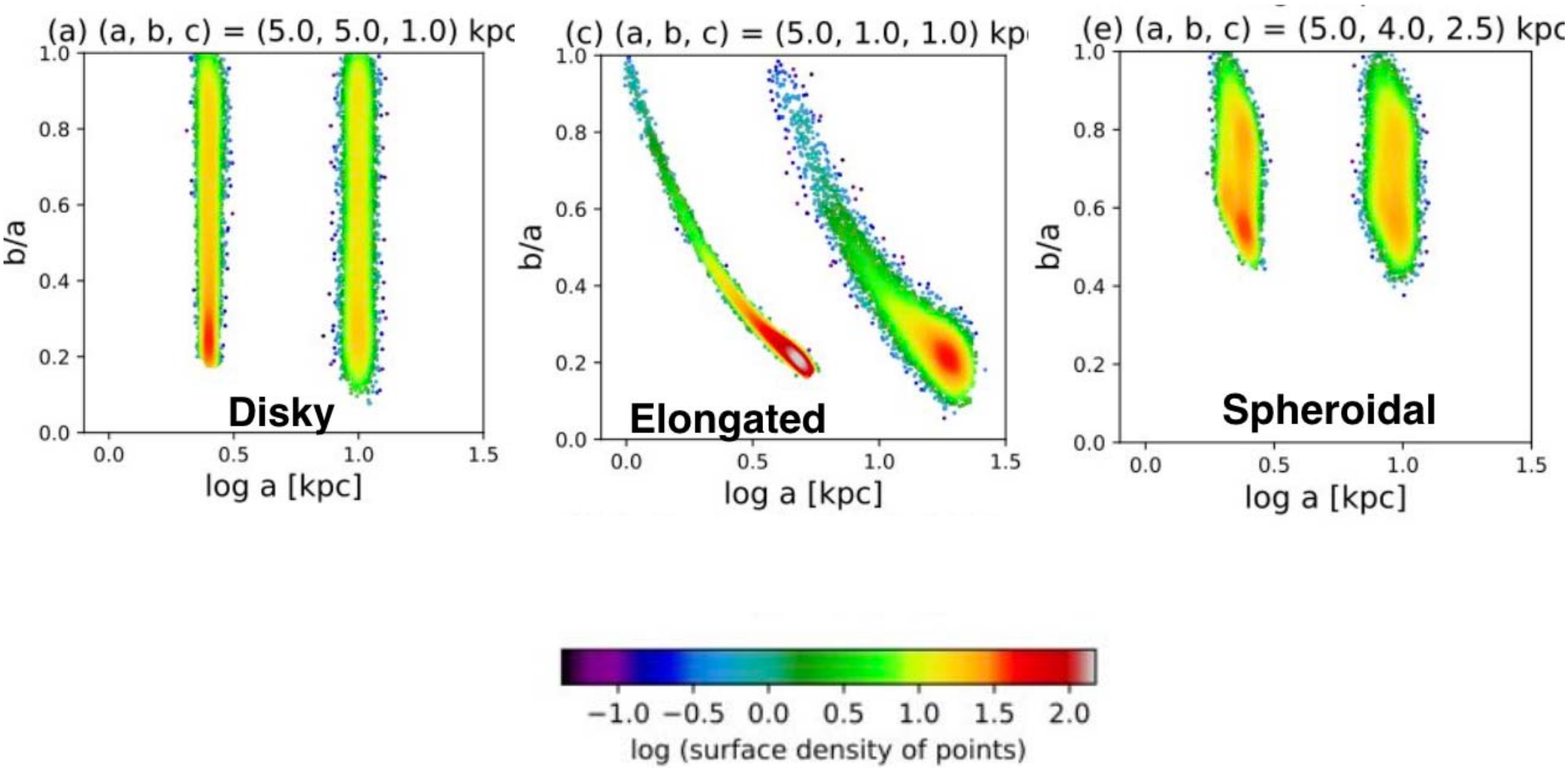


Prolate galaxies have small axis ratios except end-on

The Evolution of Galaxy Shapes in CANDELS: from Prolate to Oblate

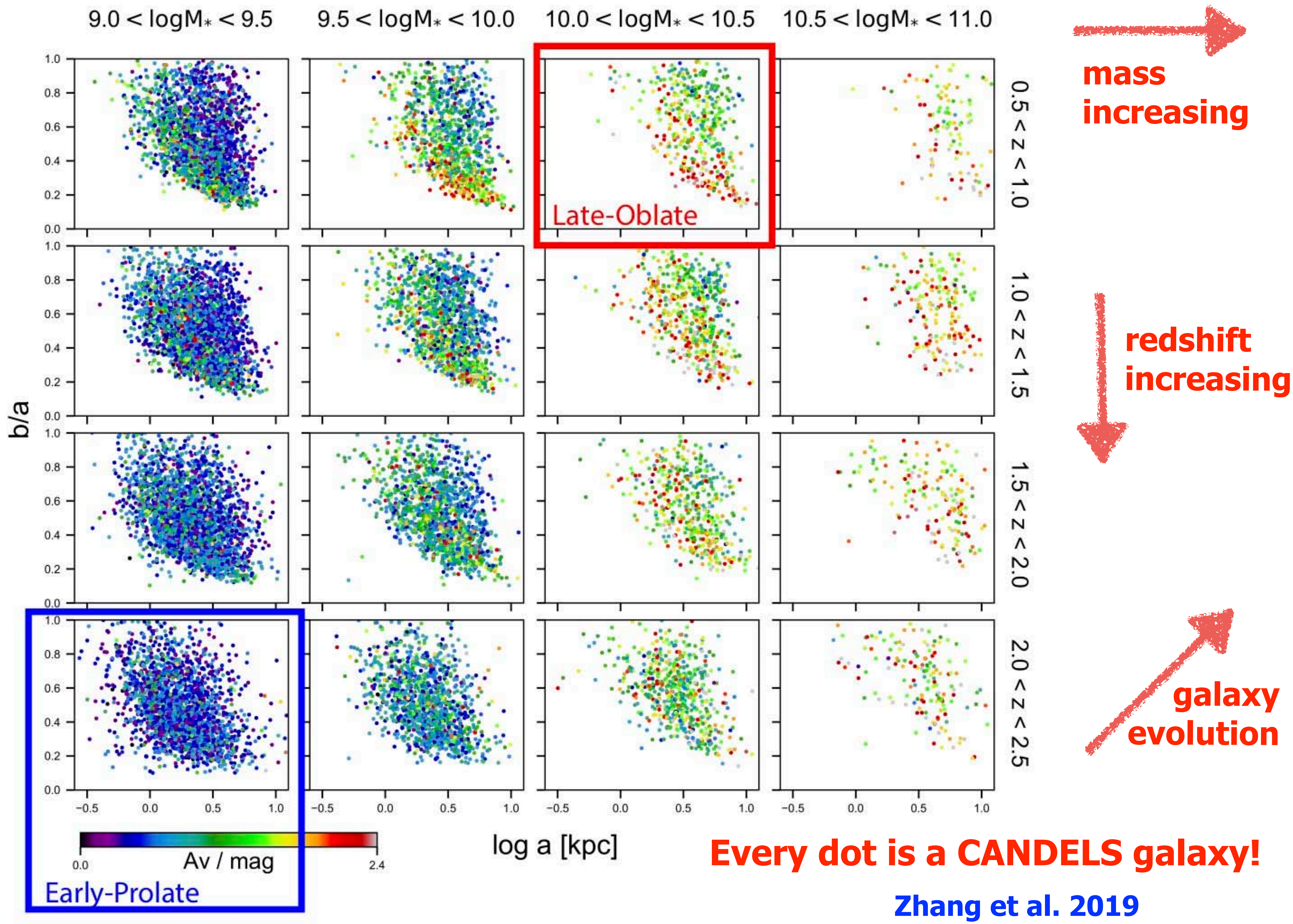
Haowen Zhang, Joel R. Primack, S. M. Faber, David C. Koo, Avishai Dekel, Zhu Chen, Daniel Ceverino, Yu-Yen Chang, Jerome J.
Yicheng Guo, Lin Lin, and Arjen van der Wel MNRAS 484, 5170 (2019)

b/a - log a distribution modeling to determine the shape distribution statistics



The Evolution of Galaxy Shapes in CANDELS: from Prolate to Oblate

Projected b/a - log a distributions of CANDELS galaxies in redshift-mass bins



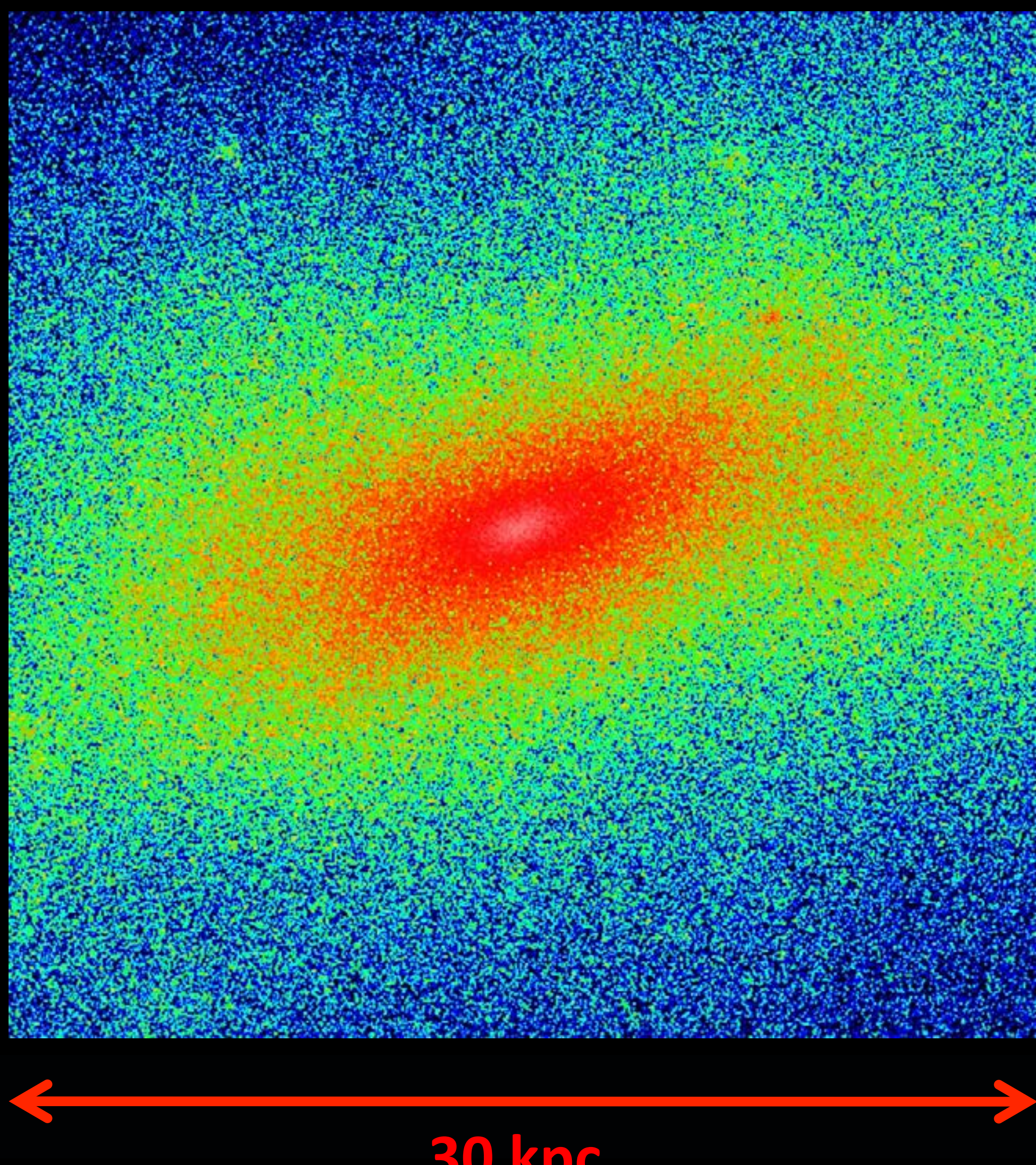
Every dot is a CANDELS galaxy!

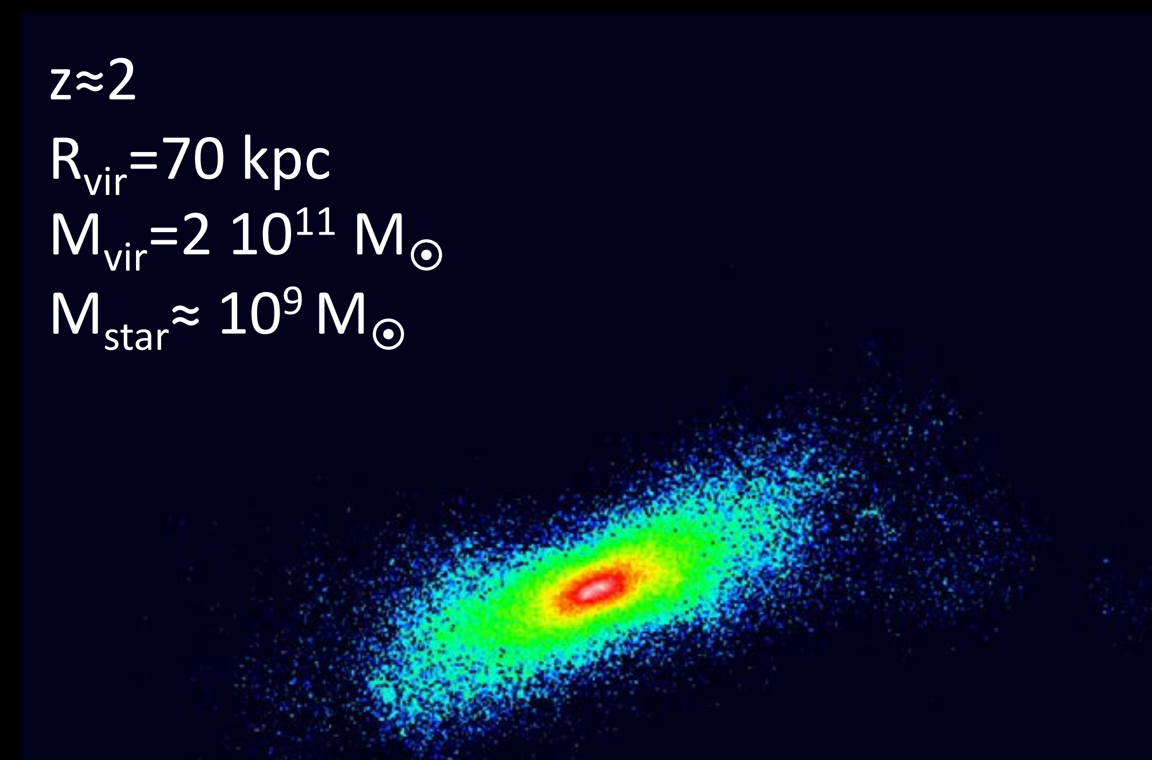
Zhang et al. 2019

Our cosmological zoom-in simulations often produce elongated galaxies like the observed ones. The elongated distribution of stars follows the elongated inner dark matter halo.

Prolate DM halo → elongated galaxy

DM **VELA28RP** **stars**

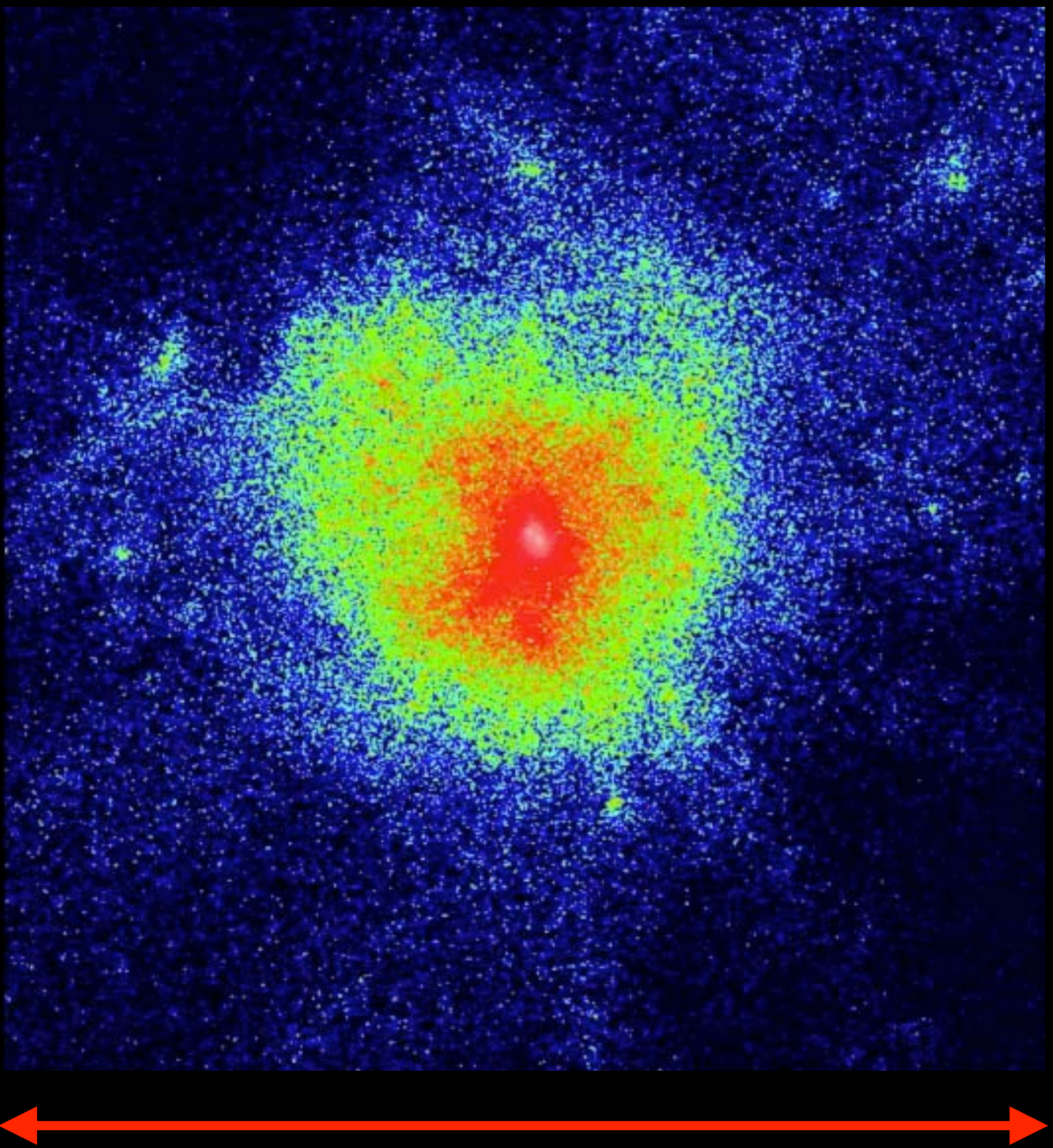


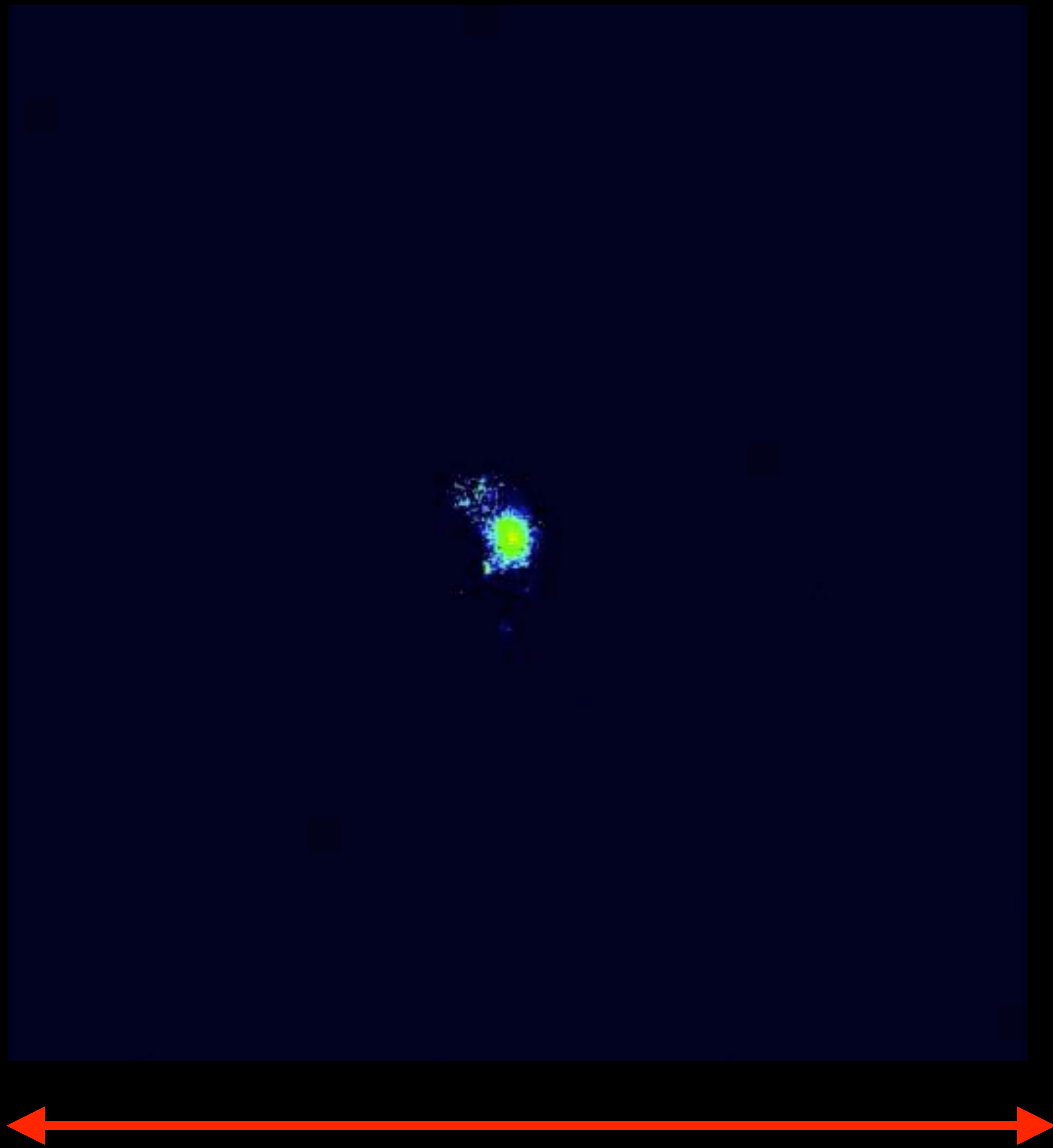


Dark matter halos are elongated, especially near their centers. Initially stars follow the gravitationally dominant dark matter, as shown. But later as the ordinary matter central density grows and it becomes gravitationally dominant, the star and dark matter distributions both become disky — as observed by Hubble Space Telescope (van der Wel+ ApJL Sept 2014).

Our cosmological zoom-in simulations often produce elongated galaxies like observed ones. The elongated distribution of stars follows the elongated inner dark matter halo. Here we show the evolution of the dark matter and stellar mass distributions in our zoom-in galaxy simulation VELA28, viewed from the same fixed vantage point.

DM **VELA28-gen3** **stars**

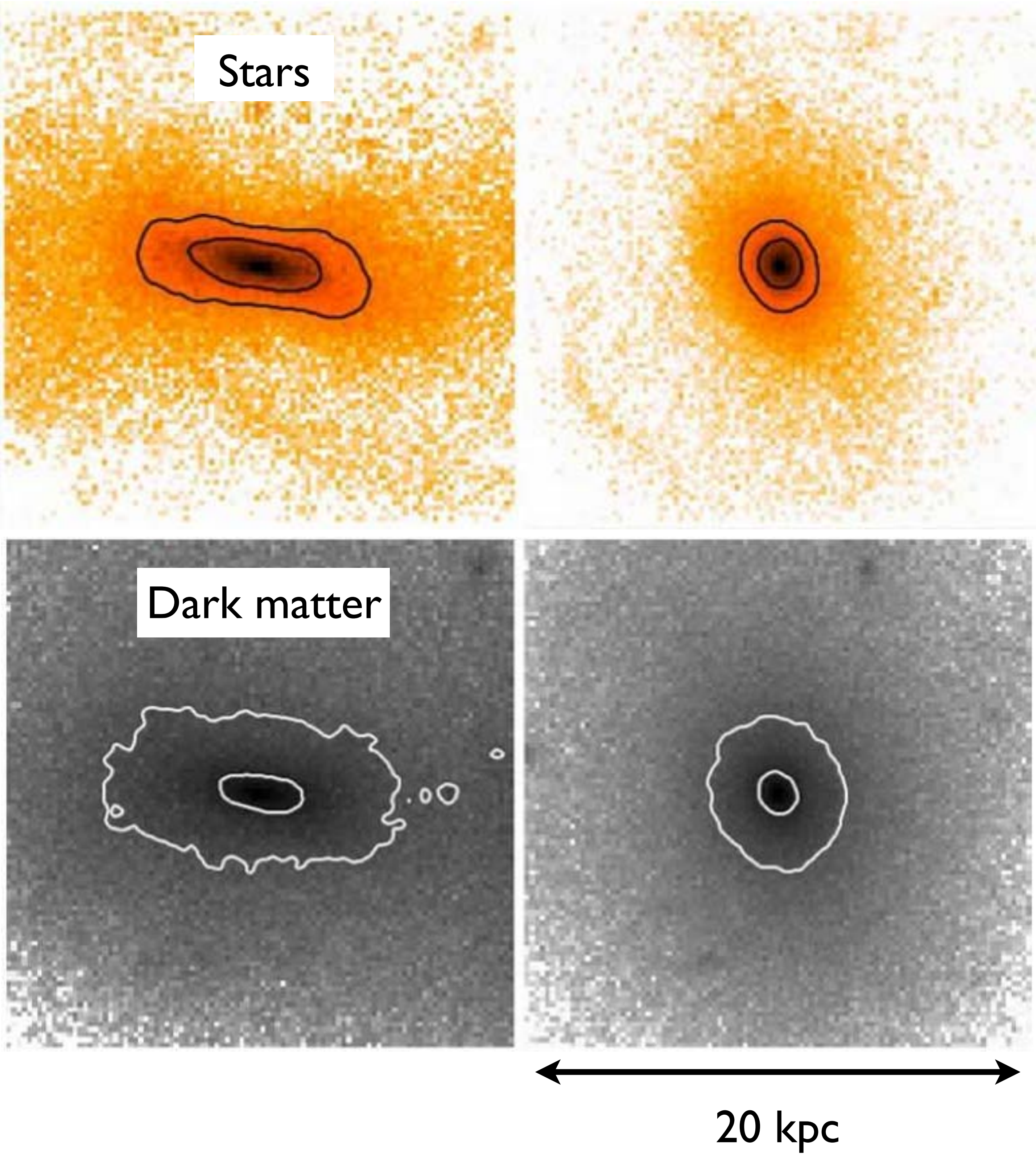




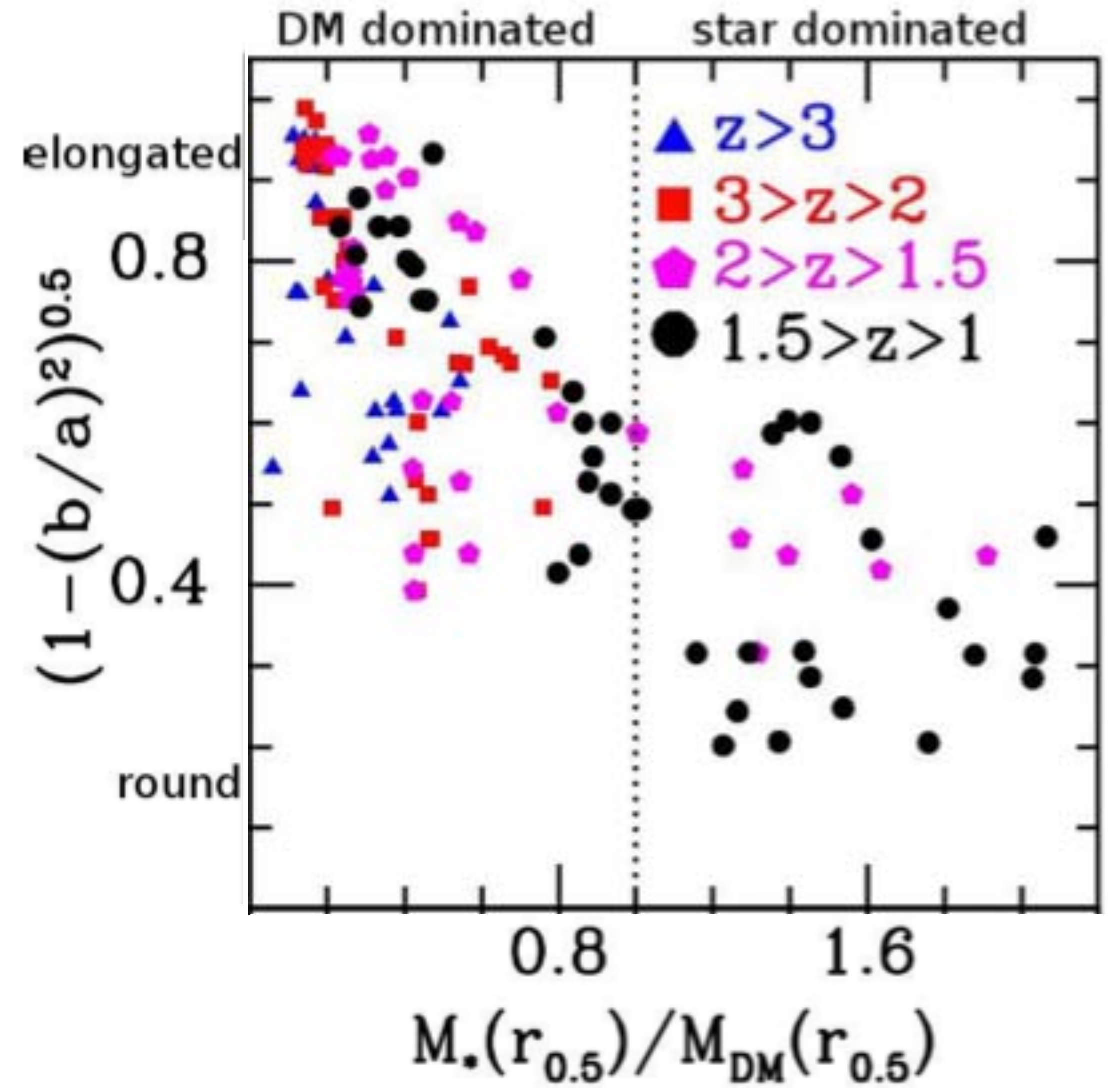
30 kpc **30 kpc**

Formation of elongated galaxies with low masses at high redshift

Daniel Ceverino, Joel Primack and Avishai Dekel MNRAS 2015



$M_* < 10^{10} M_\odot$ at z=2



Tomassetti et al. 2016 MNRAS

Simulated elongated galaxies are aligned with cosmic web filaments, become round after compaction (gas inflow fueling central starburst)

Pandya, Primack, et al. 2019 Alignments of prolate galaxies trace cosmic web?

“Face Recognition for Galaxies”

Deep Learning Identifies High-z Galaxies in a Central Blue Nugget Phase in a Characteristic Mass Range

Marc Huertas-Company, Joel Primack, Avishai Dekel, David Koo, Sharon Lapiner, Daniel Ceverino, Raymond Simons, Greg Snyder, et al. ApJ 2018

Cosmological zoom-in simulations model how individual galaxies evolve through the interaction of atomic matter, dark matter, and dark energy

Our VELA galaxy simulations agree with HST CANDELS observations that most galaxies start prolate, becoming spheroids or disks after compaction events

A deep learning code was trained with VELA galaxy images plus metadata describing whether they are pre-compaction, compaction, or post-compaction

The trained deep learning code was able to identify the compaction and post-compaction phases in CANDELized images

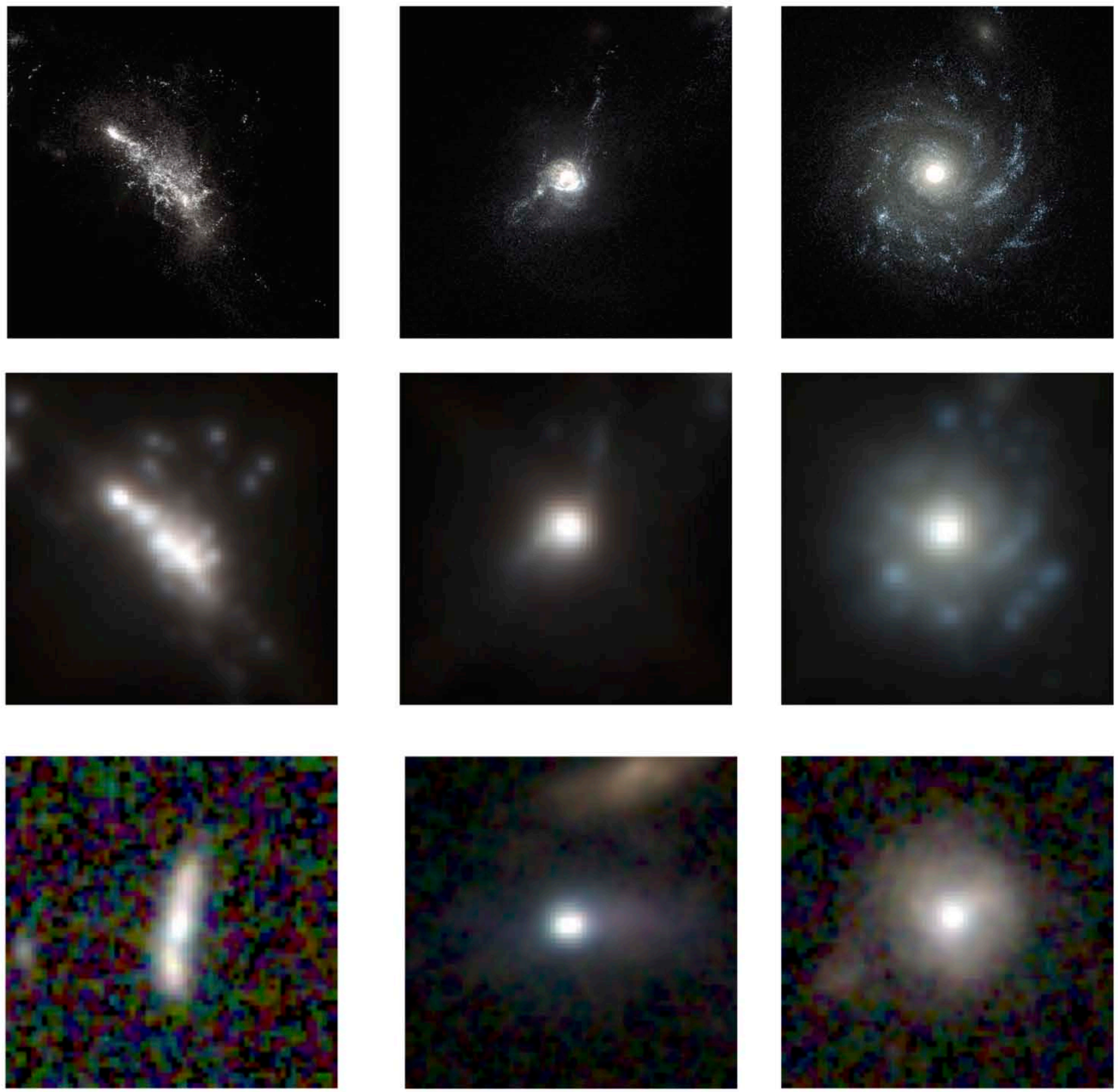
The trained deep learning code was also able to identify these phases in real HST CANDELS observations, finding that compaction occurred for stellar mass $10^{9.5-10.3}$ M_{sun}, as in the simulations

James Webb Space Telescope will allow us to do even better

"Face Recognition for Galaxies"

Huertas-Company, Primack, et al. ApJ 2018

Pre-BN	BN	Post-BN
Pre-Blue-Nugget-Stage	Blue-Nugget-Stage	Post-Blue-Nugget-Stage

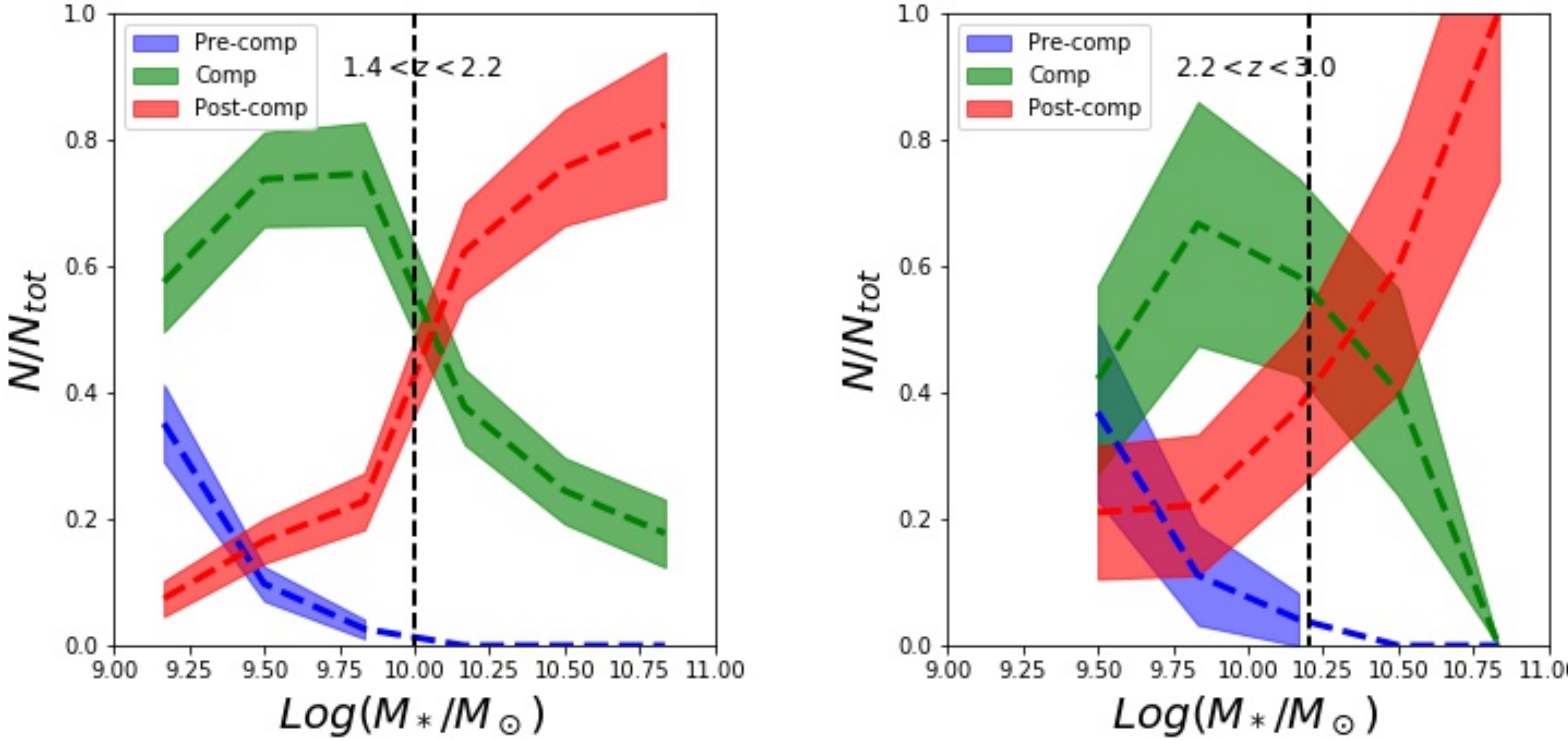


VELA High-Res Sunrise Images

VELA HST-Res Sunrise Images

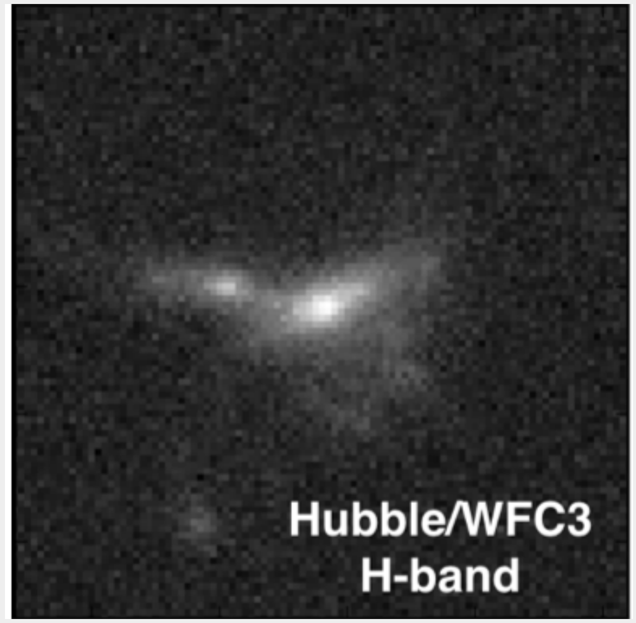
CANDELS HST Images

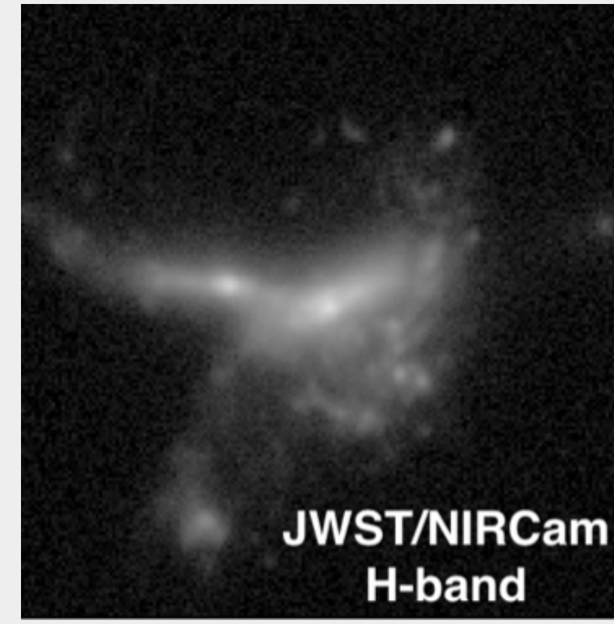
Applying the Trained Deep Learning Code to CANDELS Galaxies



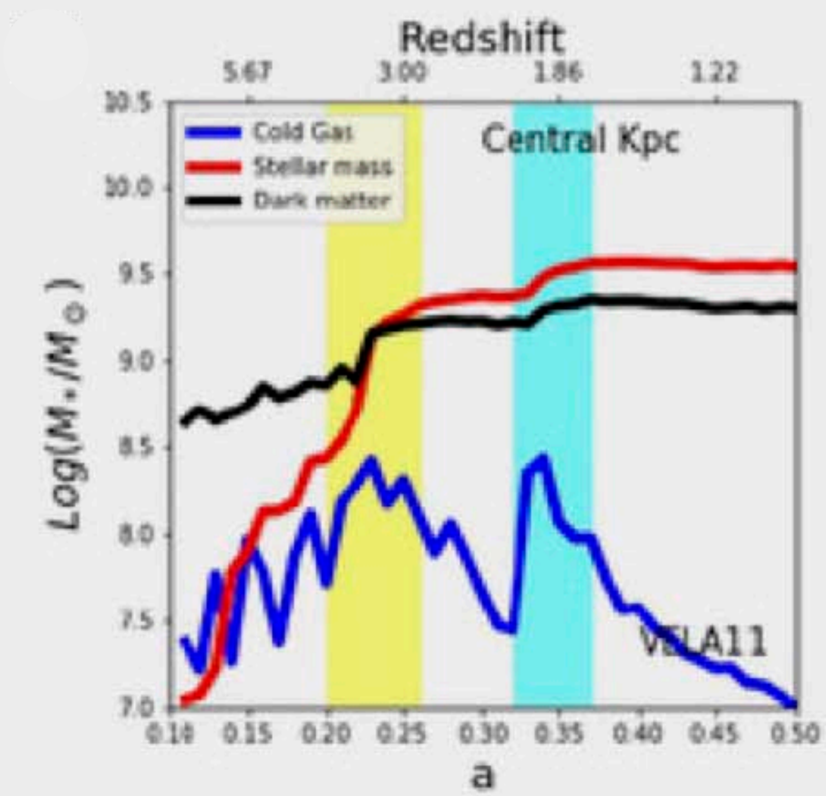
Stellar mass distributions of HST CANDELS galaxies in pre-compaction, compaction, and post-compaction phases in different redshift bins. The DL code correctly shows the temporal evolution. Galaxies in the compaction phase typically peak at stellar masses $10^{9.5-10}$ M_{sun} at all redshifts, as in the VELA simulations.

Convolutional Neural Net (Deep Learning) Galaxy Evolution Phase Determination: HST vs. JWST



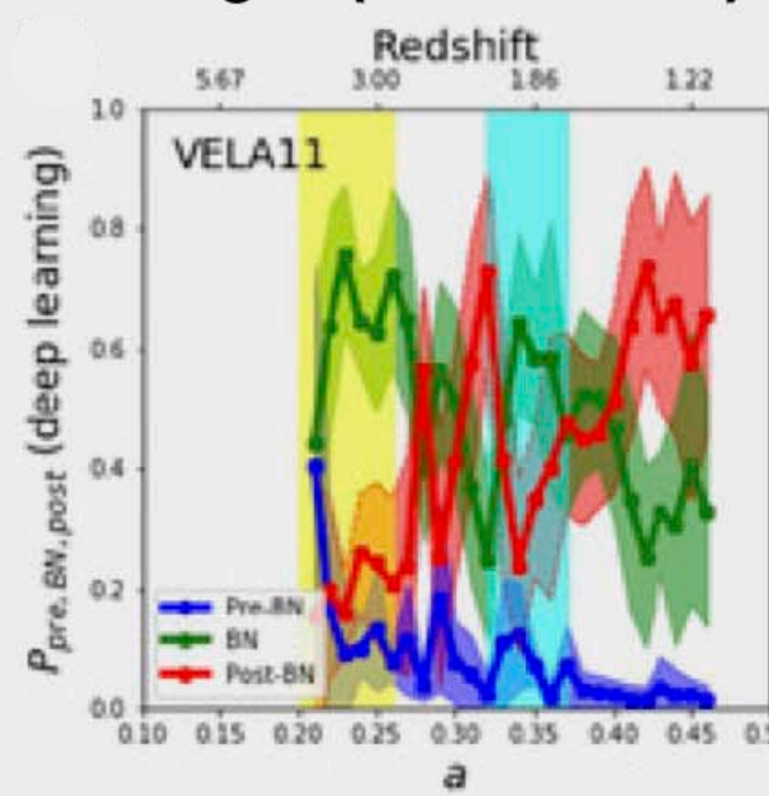


Simulation Metadata



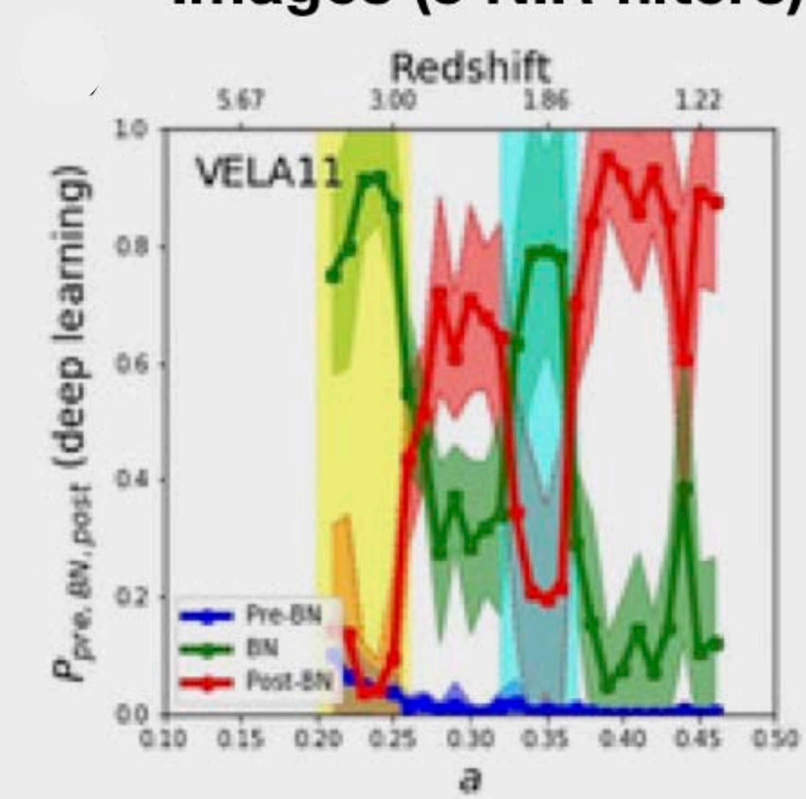
Simulated galaxy with two compaction events

CNN Trained with HST-like Images (3 NIR filters)



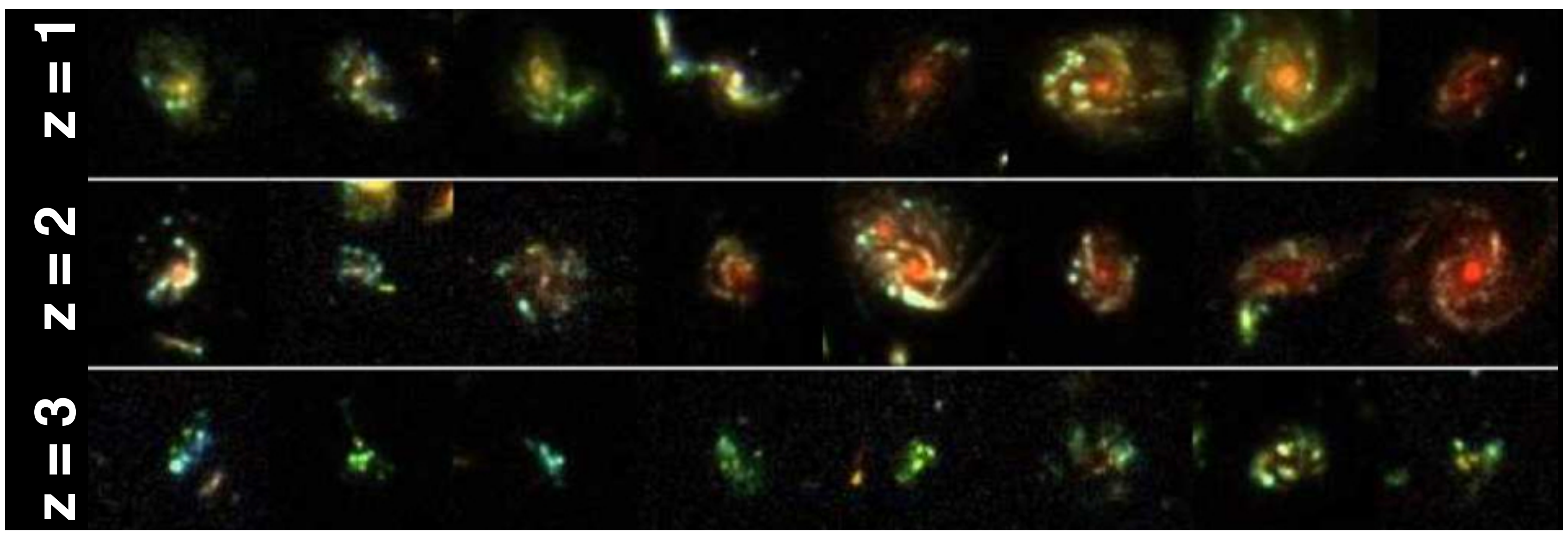
Deep learning struggles with mock HST images

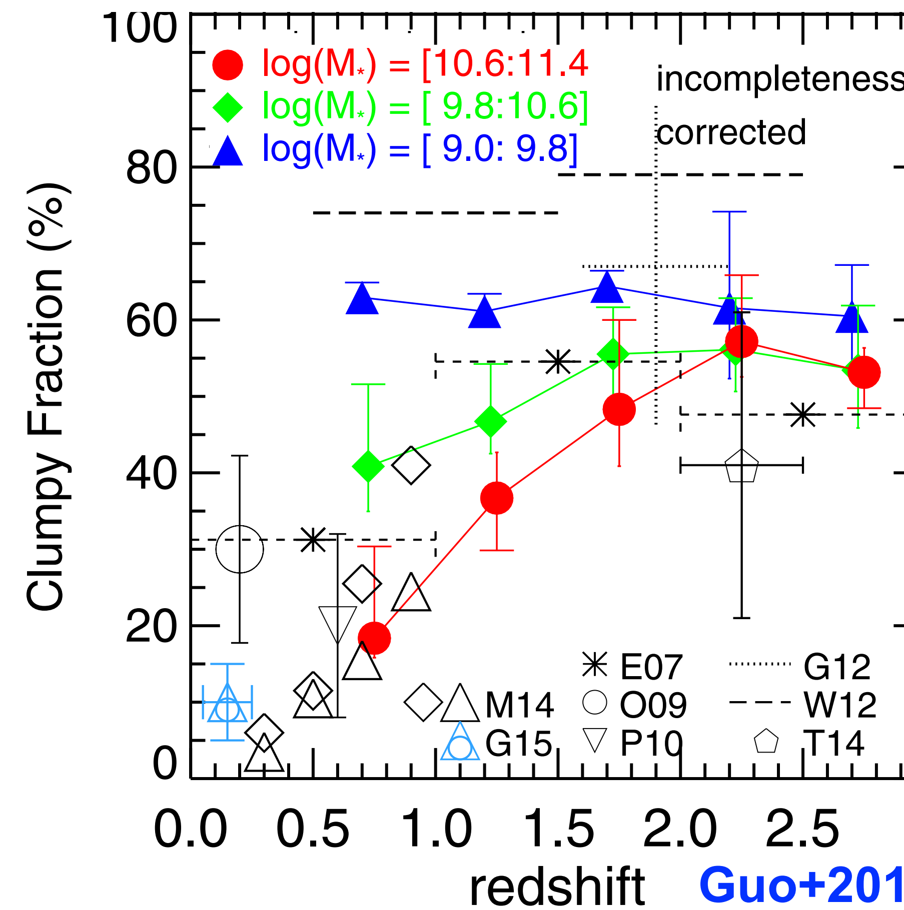
CNN Trained with JWST-like Images (3 NIR filters)



Deep learning does much better with JWST images

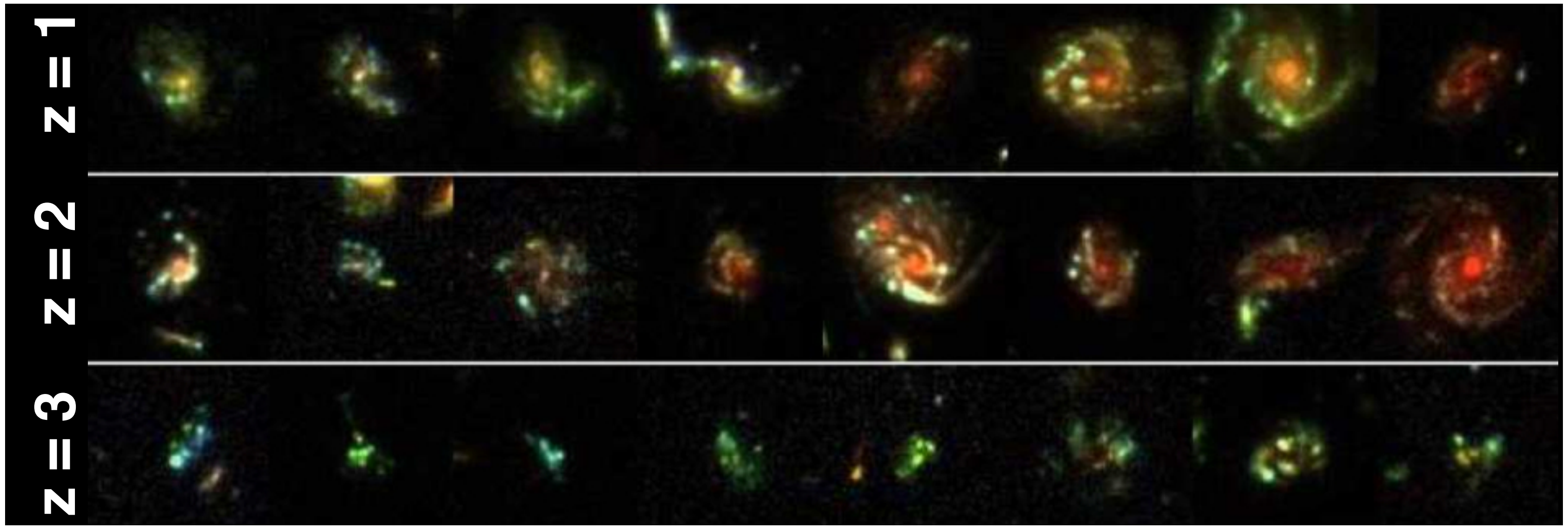
Massive stellar clumps are seen in a majority of star-forming galaxies at redshifts $z \geq 1$

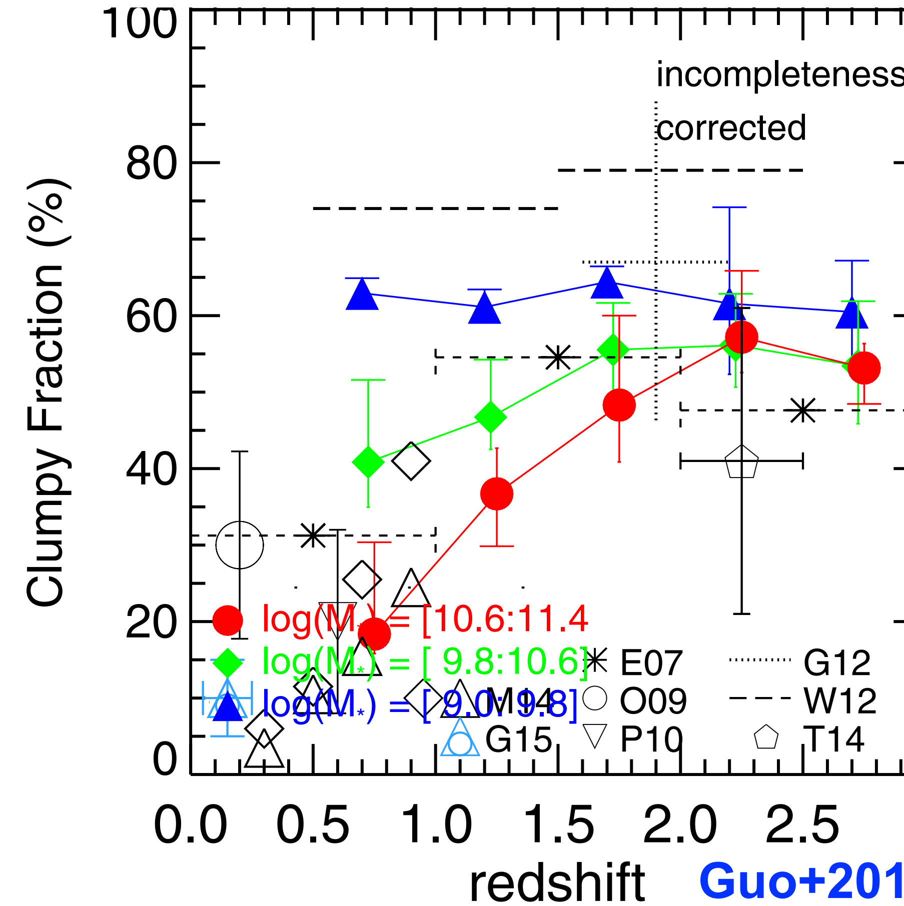




Guo+2015

Massive stellar clumps are seen in a majority of star-forming galaxies at redshifts $z \geq 1$

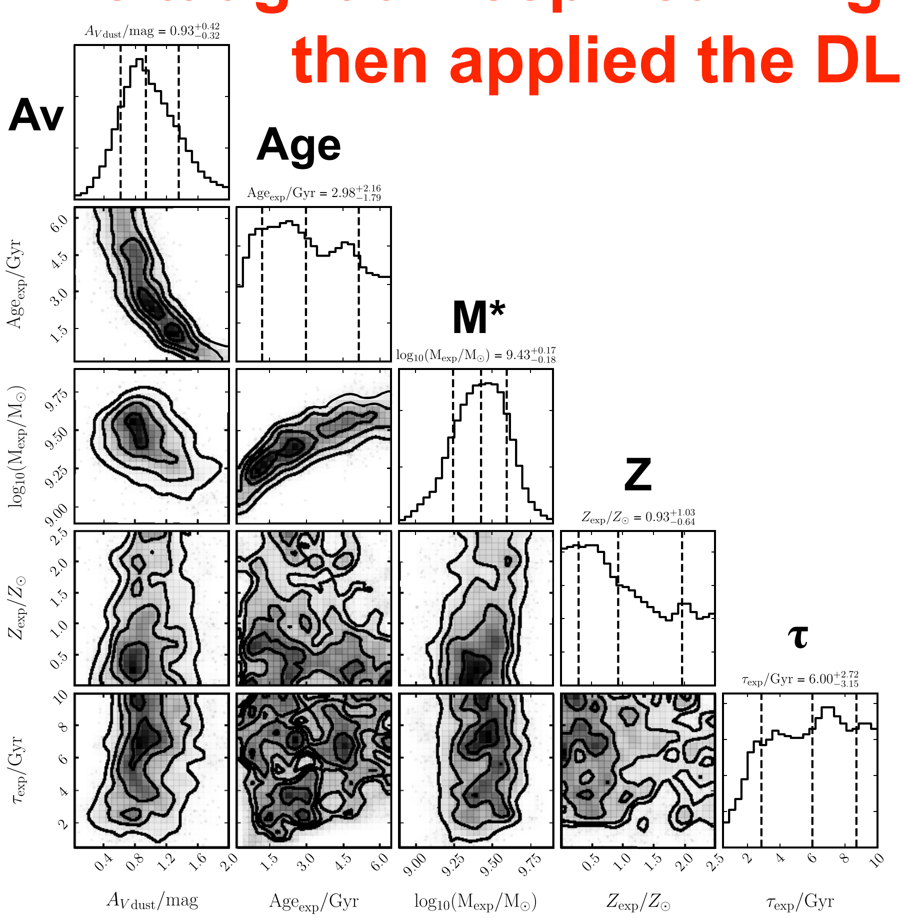


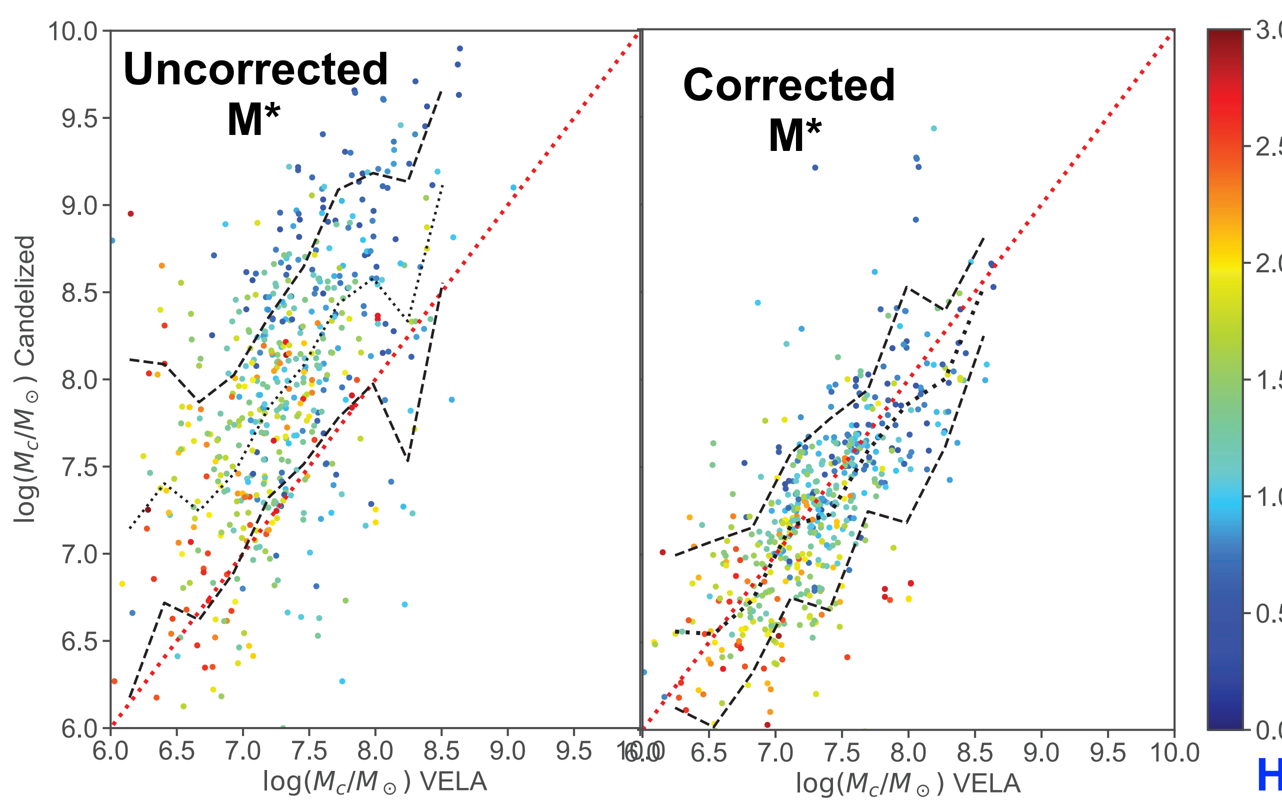


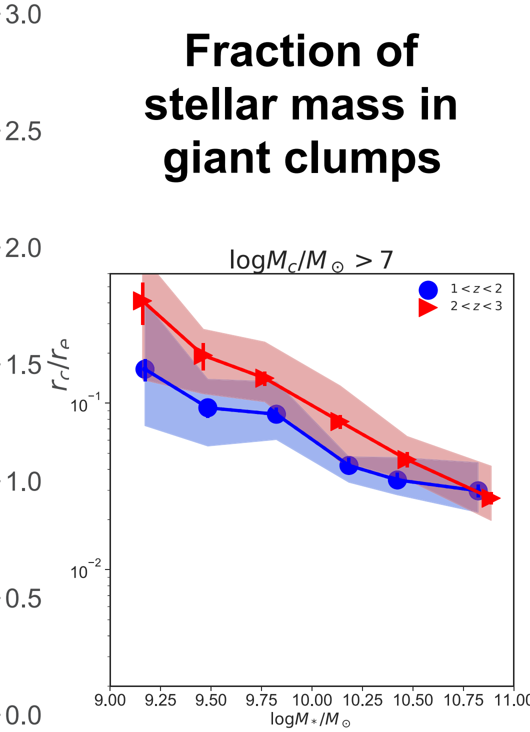
Guo+2015

Massive stellar clumps are often unresolved and blended with surrounding starlight

We taught a Deep Learning code to measure M* and ages of simulated clumps, and then applied the DL code to measure clumps in CANDELS galaxies

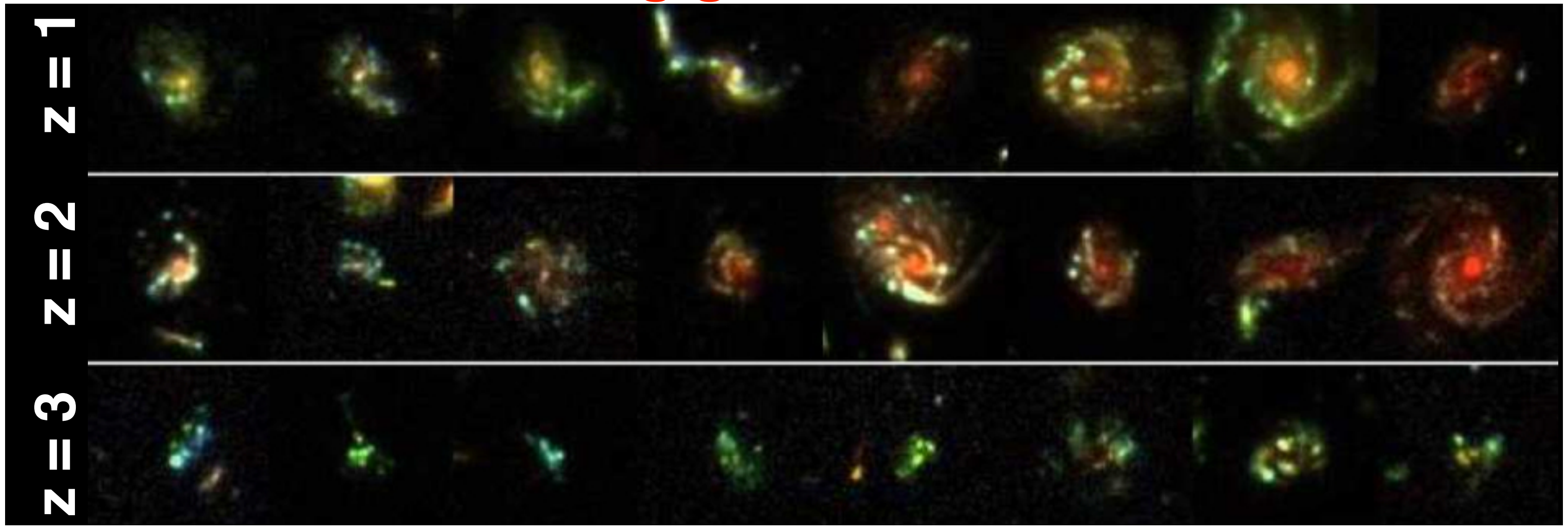


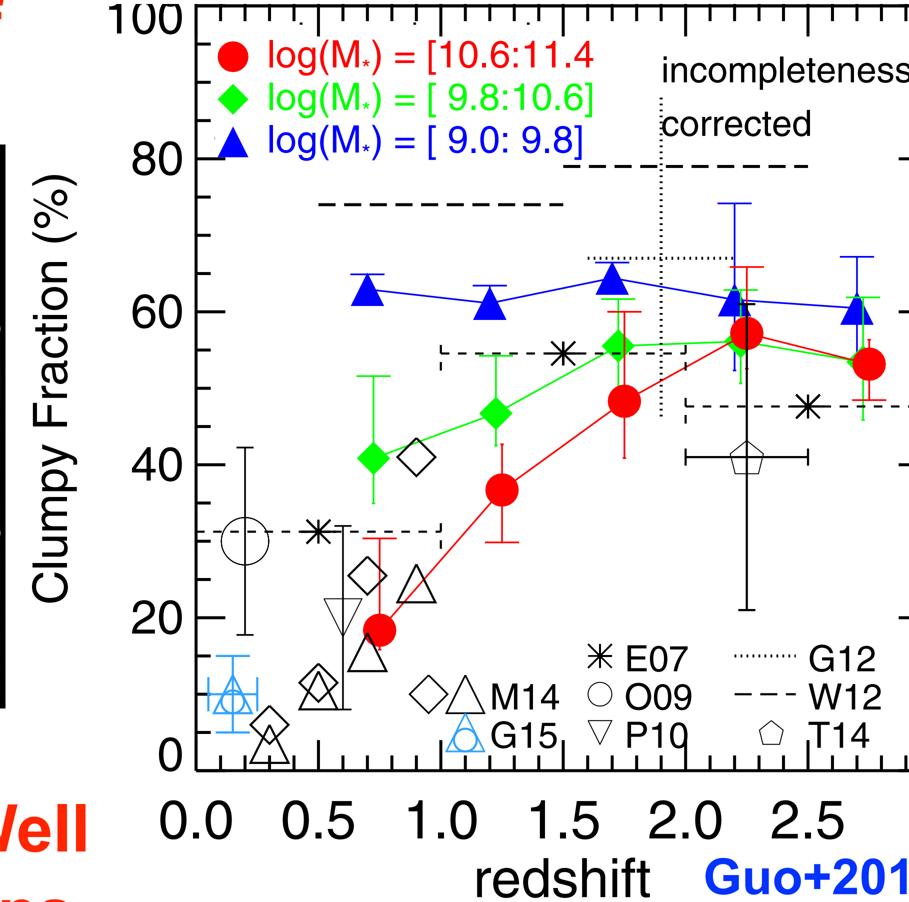




Huertas-Company+2020

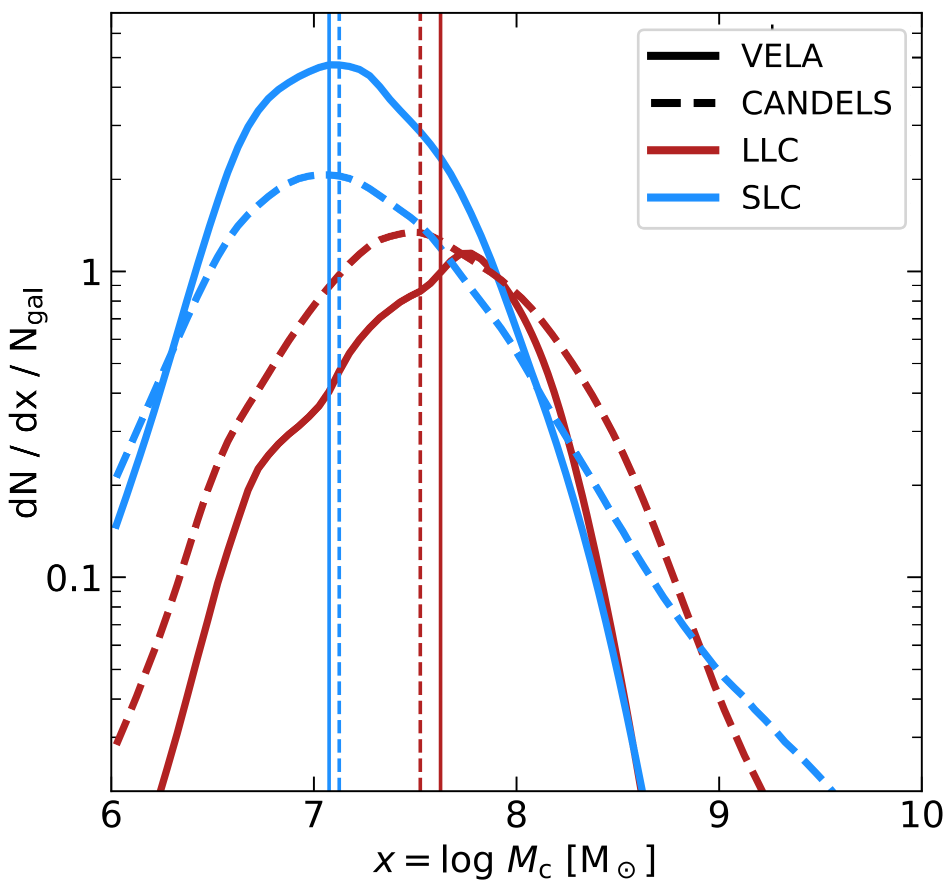
Massive stellar clumps are seen in a majority of star-forming galaxies at redshifts $z \geq 1$



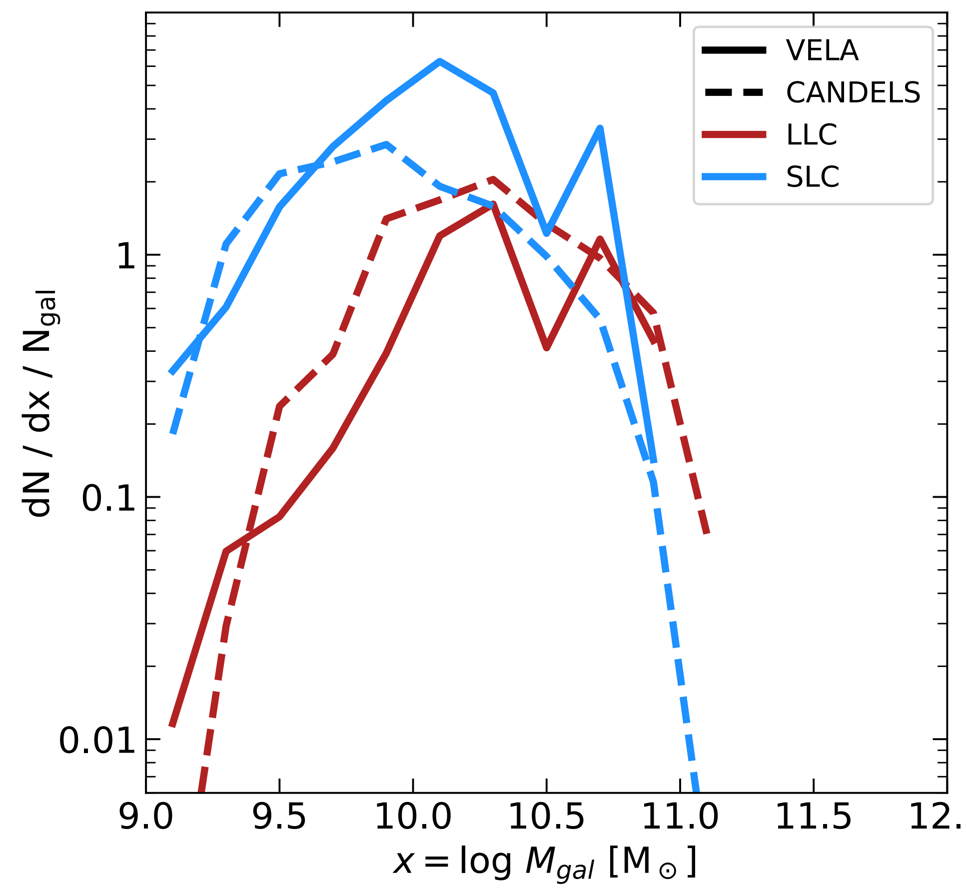


Guo+2015

VELA gen3 Simulations Agree Well with HST CANDELS Observations



Clump mass functions. Long-Lived Clumps (LLC) tend to be more massive. Vertical lines are medians.

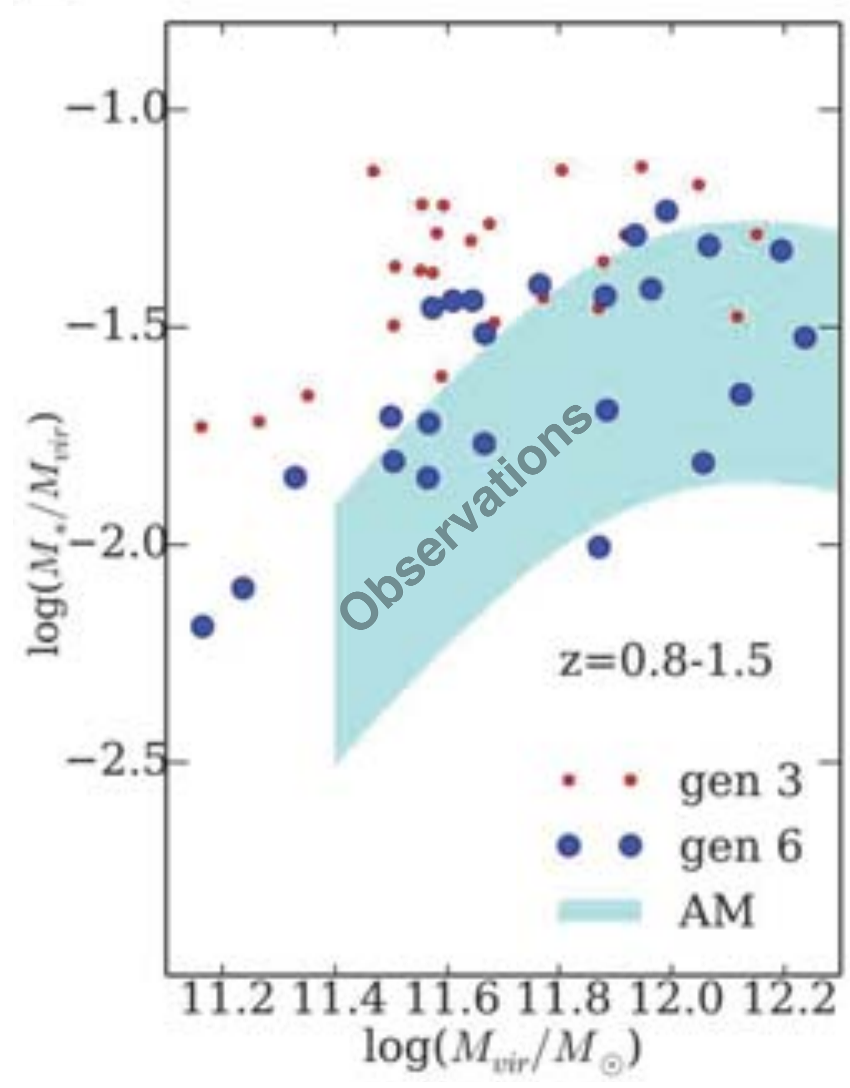


Clump host galaxy masses. Long-Lived Clumps (LLC) are found in in more massive galaxies.

Ginzburg+2021

FEEDBACK regulating star formation, both from supernovae and other stellar processes and from active galactic nuclei (AGN), is one of the greatest uncertainties in galaxy formation.

Daniel Ceverino has run our 25pc resolution VELA suite of 35 cosmological zoom-in hydro simulations five times with increasing stellar feedback. The results I've shown thus far are from VELA generation 3. But VELA gen6, with stronger feedback, is in much better agreement with the abundance matching (AM) stellar mass / halo mass relationship (shown in blue at right).



The extent to which high-resolution simulations produce massive clumps is a diagnostic for feedback.

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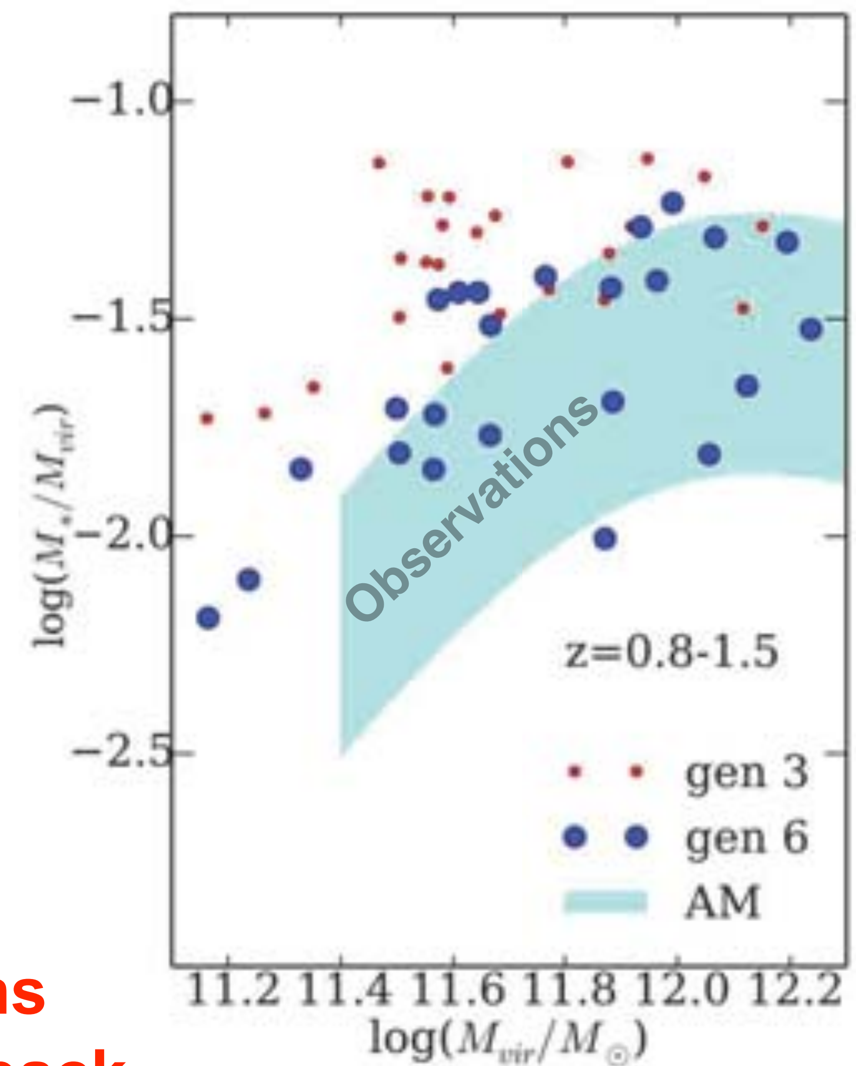
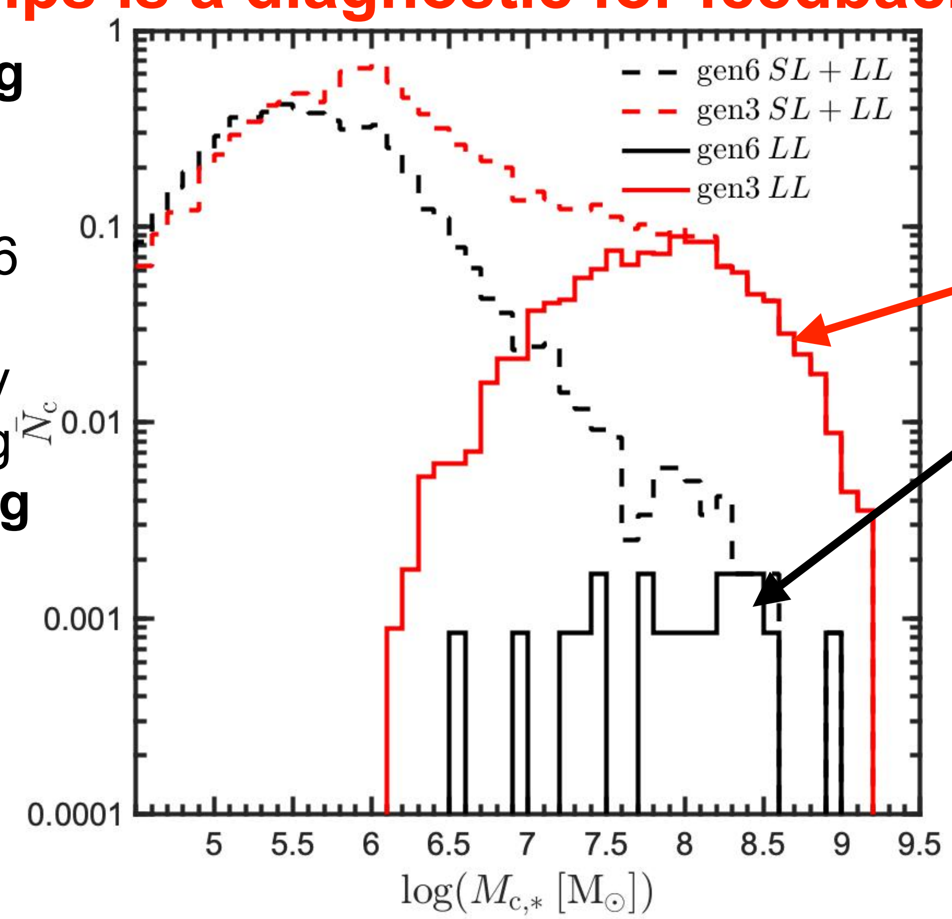


The extent to which high-resolution simulations produce massive clumps is a diagnostic for feedback.

Preliminary results demonstrating feedback effects on the clump stellar mass function. VELA gen3 (red lines, weak feedback) and gen6 (black lines, strong feedback) are two runs of the same suite of galaxy initial conditions. LL and SL are long lived and short lived clumps. **Strong feedback prevents the formation of very massive clumps and greatly decreases the abundance of long lived clumps.**

Try less ejective feedback and more preventive feedback?



This is also a problem for FIRE simulations - Oklopčić+2017

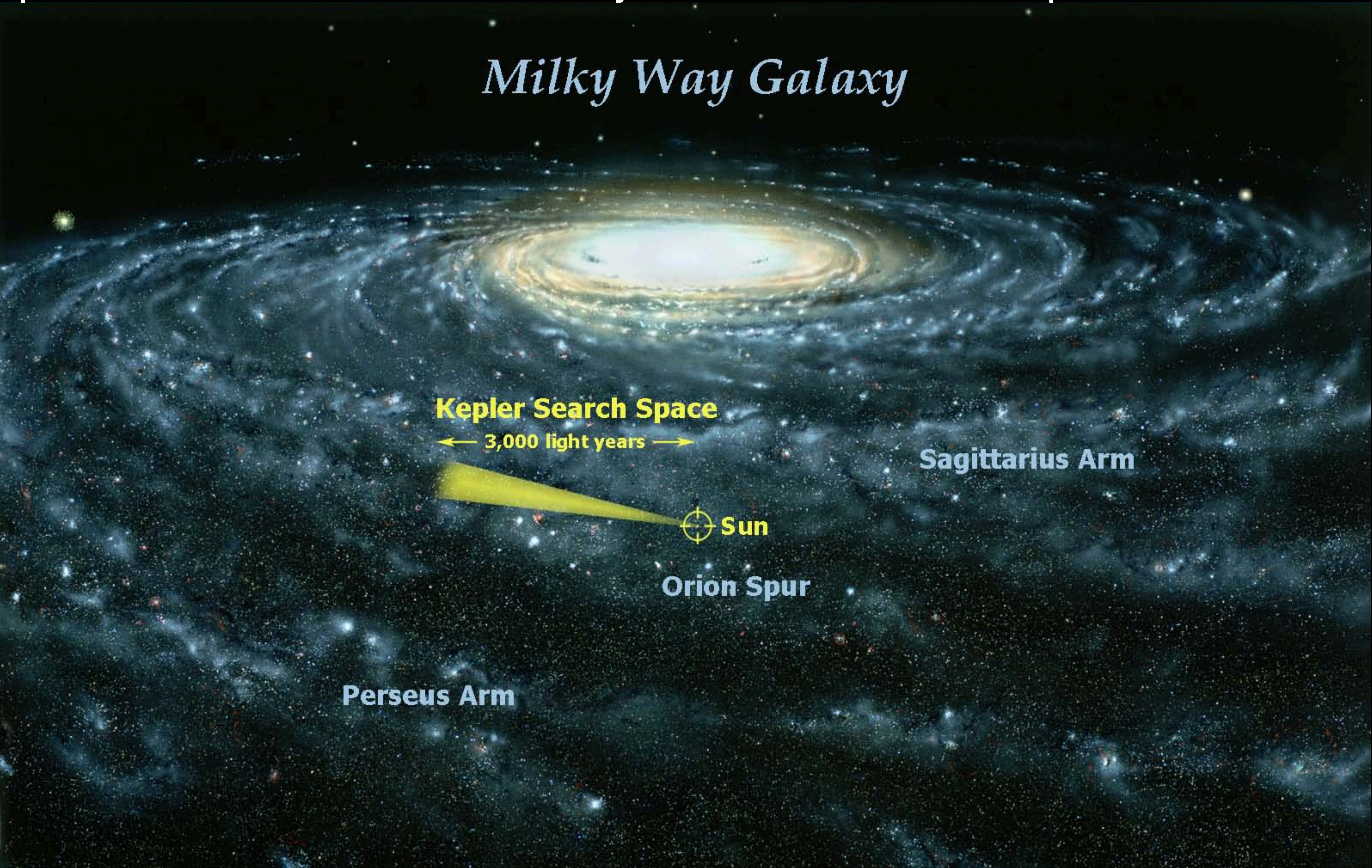
LL clumps with weak feedback (gen3)

LL clumps with strong feedback (gen6)

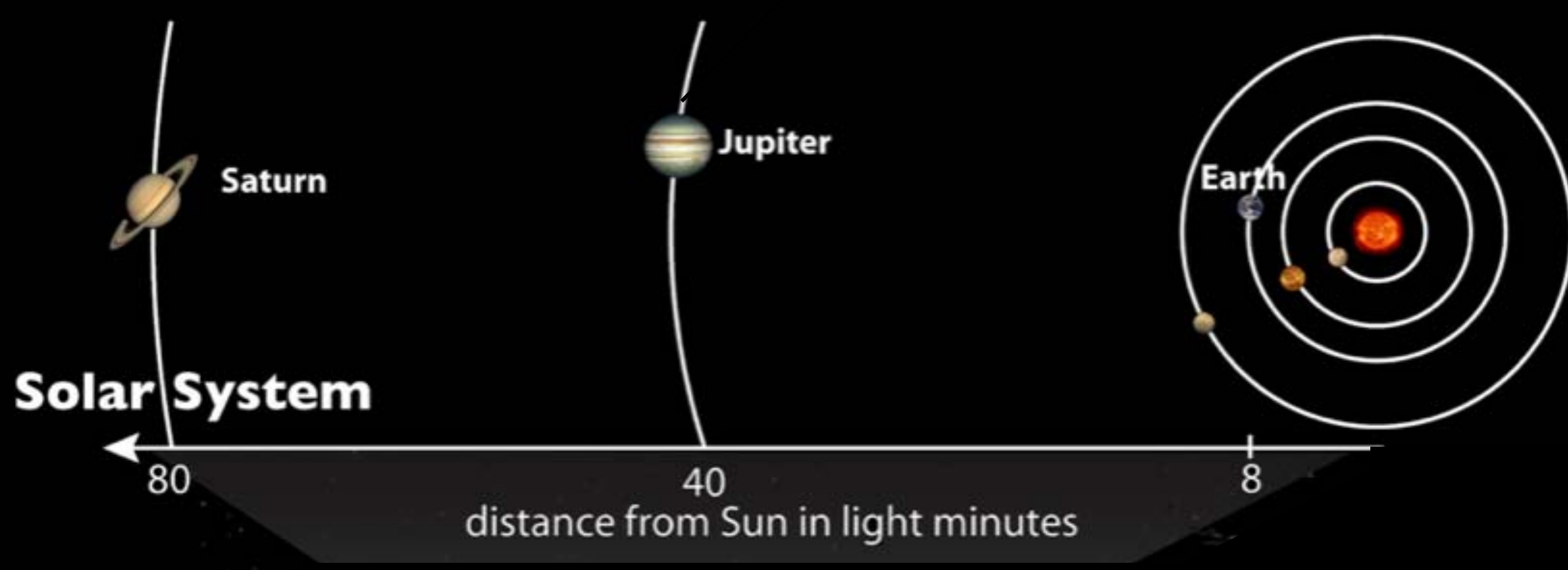
Analysis of HST and mock JWST images using Machine Learning: Huertas-Company+2020; Ginzburg+2021; Primack+ NASA ATP proposal; and work in prep.


NEW CHALLENGES IN
COSMOLOGY
GALAXY FORMATION
PLANET HABITABILITY

We have now discovered about 4000 planetary systems, mainly using star radial velocities from ground-based telescopes and planet-star transits observed by NASA's satellites Kepler and TESS.



We used to think that our system is typical, with rocky planets near our star and gas giants farther away.



Of the ~ 4000 planetary systems astronomers have discovered, there are very few like ours, with all the planets widely spaced in nearly circular orbits. Most planetary systems are much smaller.

The most common type of planet seems to be 2 to 6 times Earth's mass, a "**super-Earth**". No such planet exists in our Solar System.

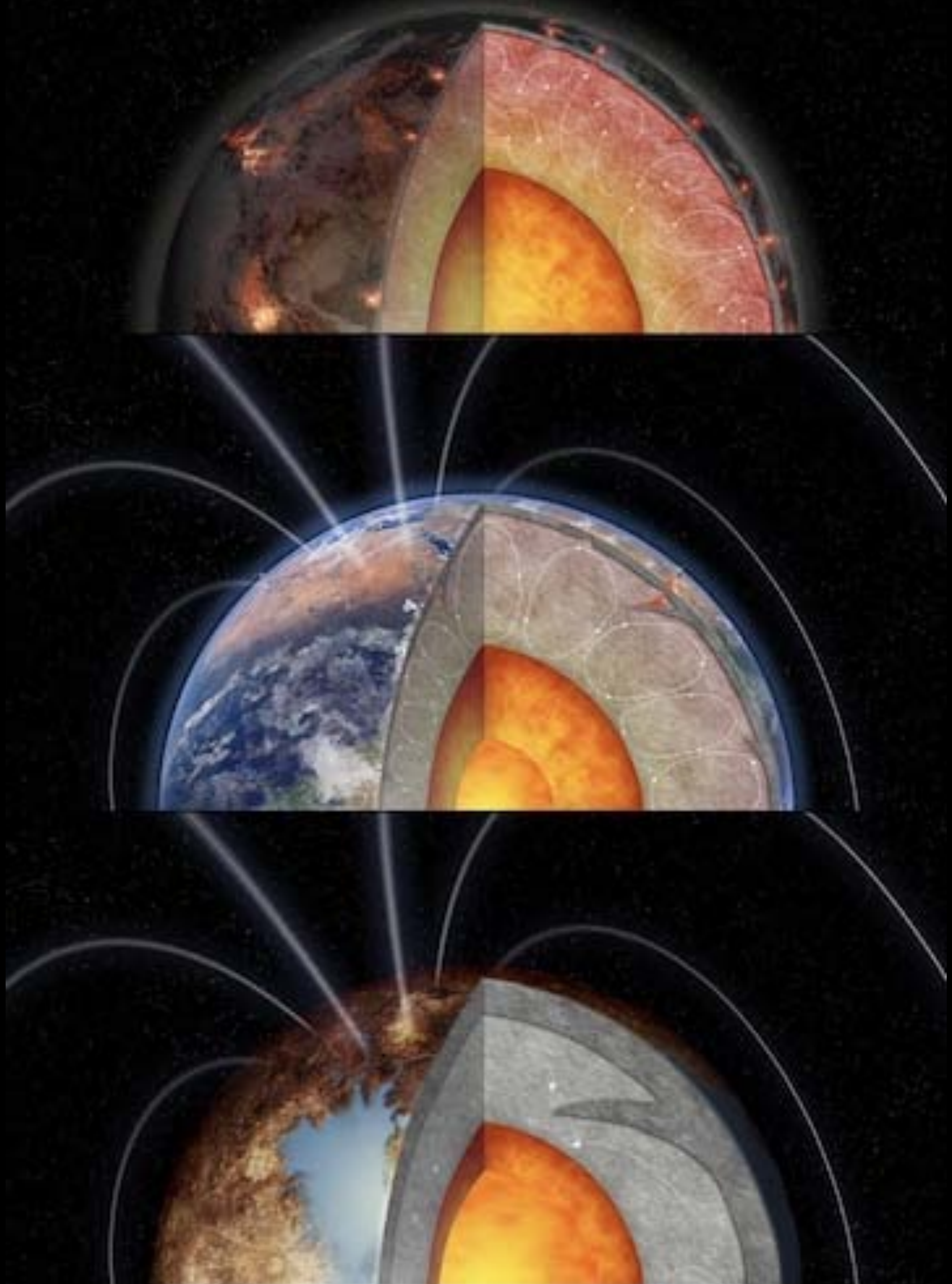
Some planets are in the **habitable zone around their stars** in which water would be in liquid form, but most of these planets are probably not hospitable to advanced forms of life. For one thing, they might not have an optimal abundance of the long-lived radioactive elements thorium and uranium to power plate tectonics and permit a magnetic dynamo. Too much Th and U would result in a lava world with frequent flood vulcanism, which caused the greatest mass extinction events on Earth. **Our living Earth may be a rare "Goldilocks" planet** with just the right amount of Th and U.

There may be **galactic habitable zones** — not too close to galaxy centers where there are frequent supernovae and AGN outbursts, nor too far where metals may be too rare to form rocky planets. However, recent measurements at $z > 0.6$ find flat or increasing gas metallicity with radius (Simons+2021).

3x Earth's Th and U
No magnetic dynamo &
frequent flood vulcanism

Earth's Th and U
Magnetic dynamo &
plate tectonics

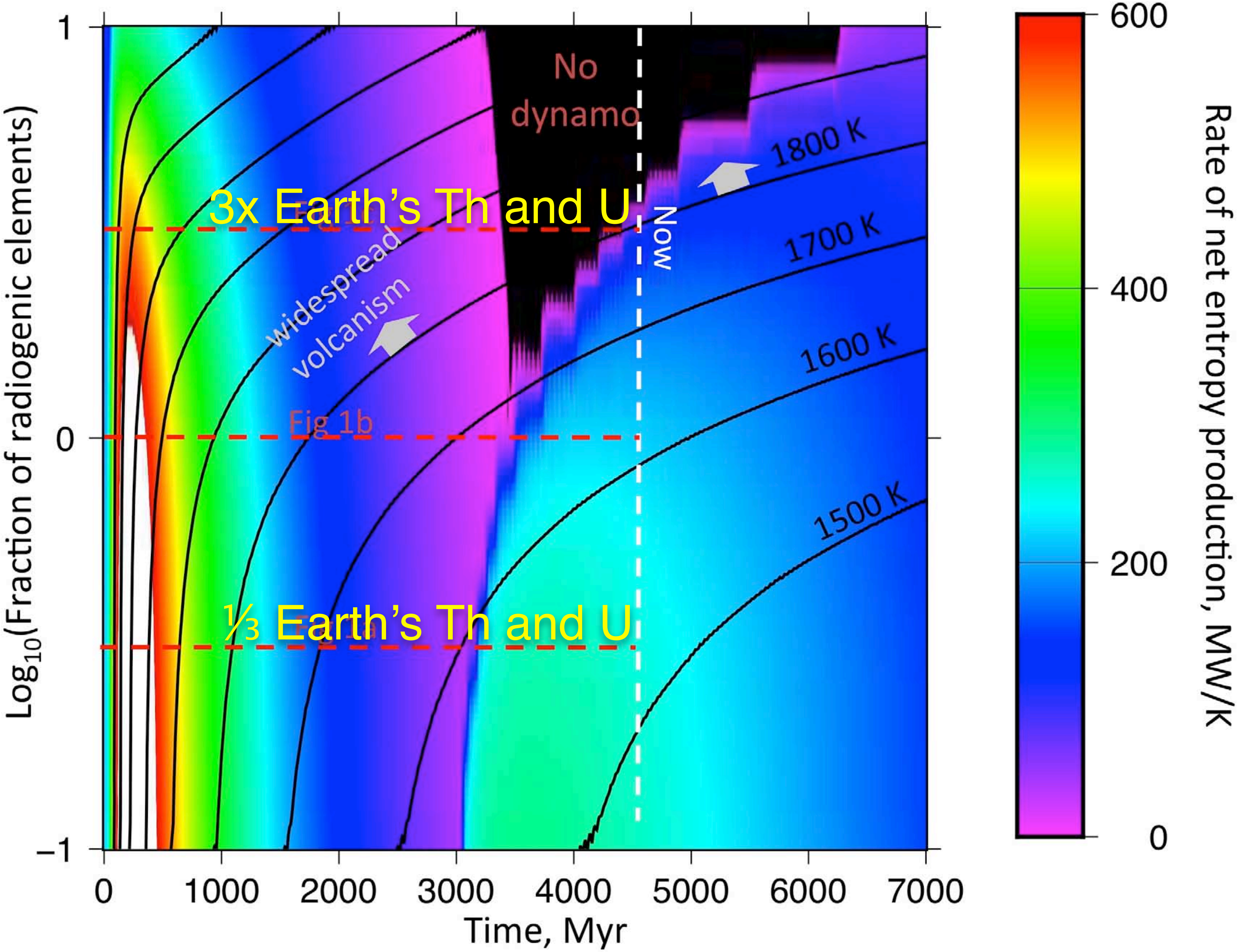
⅓ Earth's Th and U
Magnetic dynamo
but no plate tectonics



Radiogenic Heating and its Influence on Rocky Planet Dynamos and Habitability

Francis Nimmo, Joel Primack, S. M. Faber, Enrico Ramirez-Ruiz, and Mohammadtaher Safarzadeh

Astrophysical Journal Letters (November 2020)



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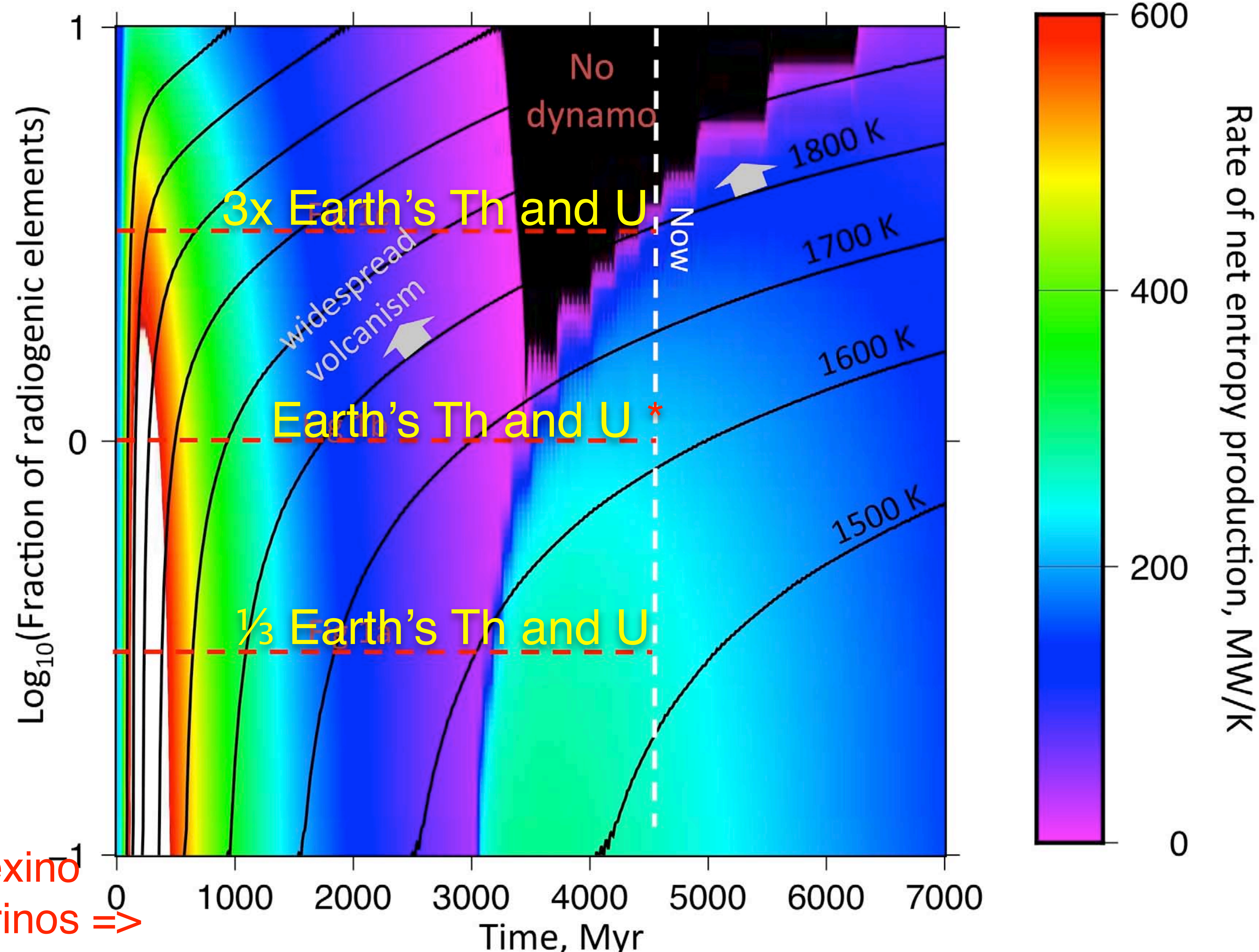
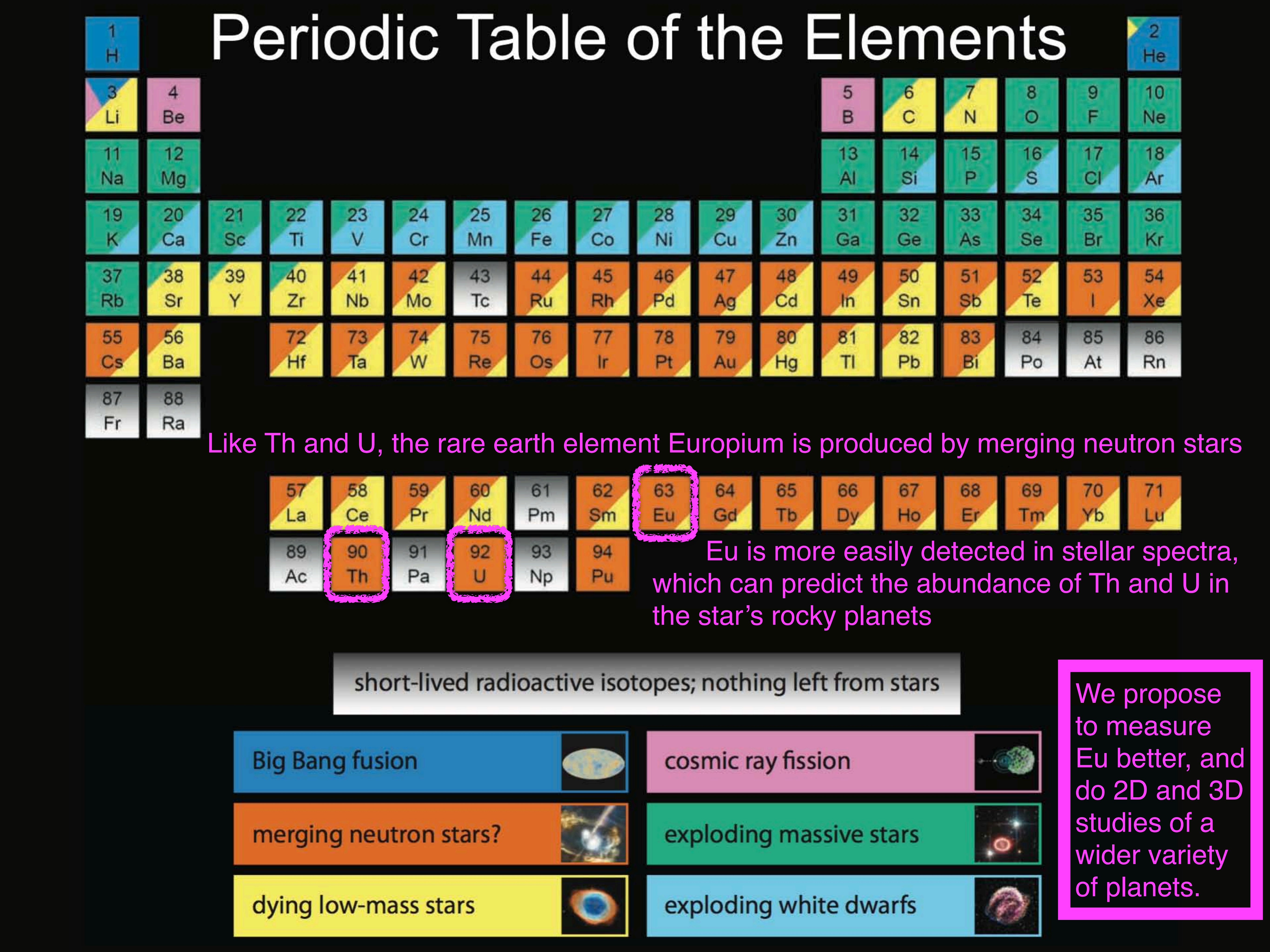


*Borexino geoneutrinos => higher Th & U

Periodic Table of the Elements



Periodic Table of the Elements
1 H
2 He
3 Li
4 Be
5 B
6 C
7 N
8 O
9 F
10 Ne
11 Na
12 Mg
13 Al
14 Si
15 P
16 S
17 Cl
18 Ar
19 K
20 Ca
21 Sc
22 Ti
23 V
24 Cr
25 Mn
26 Fe
27 Co
28 Ni
29 Cu
30 Zn
31 Ga
32 Ge
33 As
34 Se
35 Br
36 Kr
37 Rb
38 Sr
39 Y
40 Zr
41 Nb
42 Mo
43 Tc
44 Ru
45 Rh
46 Pd
47 Ag
48 Cd
49 In
50 Sn
51 Sb
52 Te
53 I
54 Xe
55 Cs
56 Ba
72 Hf
73 Ta
74 W
75 Re
76 Os
77 Ir
78 Pt
79 Au
80 Hg
81 Tl
82 Pb
83 Bi
84 Po
85 At
86 Rn
87 Fr
88 Ra
Like Th and U, the rare earth element Europium is produced by merging neutron stars
57 La
58 Ce
59 Pr
60 Nd
61 Pm
62 Sm
63 Eu
64 Gd
65 Tb
66 Dy
67 Ho
68 Er
69 Tm
70 Yb
71 Lu
89 Ac
90 Th
91 Pa
92 U
93 Np
94 Pu
Eu is more easily detected in stellar spectra, which can predict the abundance of Th and U in the star's rocky planets
short-lived radioactive isotopes; nothing left from stars
Big Bang fusion
cosmic ray fission
merging neutron stars?
exploding massive stars
dying low-mass stars
exploding white dwarfs
We propose to measure Eu better, and do 2D and 3D studies of a wider variety of planets.

11 December 2020

New Challenges in Cosmology

The Hubble tension between early universe and local measurements of H_0 can be resolved by a brief episode of dark energy at redshift z ~ 3500. New N-body simulations have shown that this Early Dark Energy scenario predicts earlier structure formation, e.g. ~ 50% more clusters than ΛCDM at redshift z ~ 1.

Galaxy Formation

Galaxies were long thought to start as disks, but HST images show that most galaxies instead start prolate (pickle shaped). Galaxy simulations can explain this as a consequence of the filamentary nature of the ΛCDM dark matter distribution. But comparisons between simulations and observations using novel machine learning methods reveal other potential challenges, including long-lived star-forming clumps seen in many high-redshift galaxies.

Planet Habitability

Earth may be a radioactively Goldilocks planet, with just the right amount of radiogenic heating by Th and U for a magnetic field and plate tectonics, both of which may be necessary for the evolution of complex life.

Some Concluding Thoughts

Without Dark Matter We Wouldn't Exist

With only the ordinary matter, the universe would be a low-density featureless soup

Dark matter started to form structures very early

Galaxies formed within bound "halos" of dark matter

Stars formed within galaxies, and stars made elements beyond hydrogen and helium: carbon, oxygen, …

Rocky planets formed from these heavier elements

Life began and evolved on one such planet

Dark matter is our ancestor and our friend!

Science Is Much Stranger Than Fiction

Before the discovery that most of the density of the universe is invisible, no one imagined this

What else remains to be discovered?